DEVELOPMENT OF Nb₃Sn COATINGS FOR SUPERCONDUCTING RF CAVITIES AT FERMILAB*

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author(s). Nb₃Sn is a promising alternative material for superconducting RF cavities, with proven high quality factors at medium fields and predictions for increased superheating field as well. In this contribution, we describe the latest results $\frac{1}{2}$ from the Fermilab Nb₃Sn SRF program. Early experiments biefly review efforts to bring the parameters used in the coating process into a range where they produce uniform surfaces without regions showing signs of excess tin or thin/uncoated areas. We then present the latest cavity results, after modifications to the coating recipe based on feedback from film appearance and RF performance. must with the second These results show high Q₀ at medium fields and a maxi-

COATING DEVELOPMENT

of this Niobium has been the traditional material for SRF cavi- ξ ties, but new materials offer the potential for high quality factors (Q₀) at higher temperature to take advantage of E higher cryogenic efficiency. Nb₃Sn is a new SRF material [±] which already showed promise in pioneering research (e.g. \mathfrak{F} [1] and [2]), and has had significant progress recently after the research was revived at Cornell University [3].

At Fermilab, a Nb₃Sn SRF program recently began, and 201 the first cavities were coated in 2017. The early program has focused on achieving strong performance in single cell licence (cavities, so that similarly strong performance can be achieved after scaling up to the multicell and low frequency cavities that the 20" diameter coating chamber was \succeq designed to accommodate.

Cavities are coated via the vapor diffusion process. The 20 furnace raises the cavity temperature to ~1100 C, and a heater raises the tin source even higher to evaporate tin. erms of Excess or deficit of tin can be quickly apparent after coating. Several parameters have been found to impact important observables in the coating process such as evaporation rate, total tin transferred from the crucible, presence under of non-uniformities over the surface of the cavity, and presence of poorly or overly coated regions (see Fig. 1). Some used key parameters have been found to be:

- Crucible diameter (see Fig. 2)
- Heater power •
- Coating time
- Annealing time

this work may * This work was supported by the United States Department of Energy, from Offices of High Energy Physics. Fermilab is operated by Fermi Research Alliance, LLC under Contract No. DE-AC02-07CH11359 with the United States Department of Energy. † sposen@fnal.gov

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Figure 2: Four different tungsten crucible sizes.

Several 1.3 GHz single cell cavity coatings were carried out, and a strong correlation was observed between surface appearance and RF performance. For example, a cavity with apparent residual tin on the surface showed low Q₀ at 4.4 K, and hysteresis at 2.0 K (<T_C of tin \sim 3.7 K), roughly consistent with H_c of tin ~0.03 T. The Q vs E curves at 2.0 K and 4.4 K are shown in Fig. 3.

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Figure 3: RF measurement of the cavity with apparent residual tin from Fig. 2 (bottom). Note the hysteresis at 2 K appears to be consistent with residual tin on the surface.

RECENT RESULTS

To date, the best performance of a Nb₃Sn cavity at Fermilab has been achieved with the smallest diameter crucible tested, with the heater at maximum power (crucible temperature ~1200 C). The cavity was anodized prior to coating, as advocated by Siemens in the 1970s [4] and recently revived by Cornell [5]. The cavity surface, shown in Fig. 4, visually was highly uniform, showing the matte dark gray typical of Nb₃Sn. Both half-cells showed similar appearance.



Figure 4: Cavity coating after coating parameter development with no apparent non-uniformities

The cavity was cooled slowly and uniformly through the critical temperature to avoid generating thermocurrents [6]. The cooldown is shown in Fig. 5, including signals from 4 thermometers at various locations and 4 fluxgate magne-tometers, 3 of which were equally spaced the equator, and one located on the frame. Q_L vs T obtained with a network analyzer is also shown in Fig. 5, with transition very close to the expected $T_{c} \sim 18 \text{ K}.$



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Figure 5: Cooldown through of the cavity, including temperature and magnetic field at the transition and Q_L vs T measured with a network analyser.

Q vs E was measured at 4.4 K and <1.5 K up to quench, as shown in Fig. 6. 2.0 K measurement of Q vs E occurred up to 16 MV/m before quench and up to maximum field after quench. The post-quench 2.0 K data is corrected to match the lower field data to the pre-quench data.





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T-maps were measured at <1.5 K to localize regions of high heating (Fig. 7). Some hot spots were observed, but the heating was fairly widely distributed. Most regions showed non-linear heating in the O-slope region.



Figure 7: Temperature maps with bath temperature <1.5 K showing heating over the surface before and after quench.

CONCLUSION

distribution of this work After a program to explore parameters for coating Nb₃Sn cavities, a recipe was found that resulted in high quality Ffactors up to medium fields. Near future work on singlecell cavities will focus on reproducibility, as well as sys-18) tematic variation of individual parameters to further im-20] prove performance. The lessons learned will be applied to different frequency cavities (e.g. 650 MHz) and to multicell cavities to begin to scale up from R&D-scale cavities 5 to application-scale cavities. Some Q-slope is still present 0 above ~ 10 MV/m, but the cavity shows performance close to the highest measured for Nb₃Sn accelerator cavities, and В significantly improved compared to previous results from ^o Fermilab, as well as pioneering studies at University of the Wuppertal [7]. The Q vs E data is already promising for of scaling up to larger cavities for industrial accelerators opbe used under the terms erating at 5-10 MV/m with ~5 watts of dynamic load per cavity at 4 K.

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