

THE USE OF HHG AT 4GLS

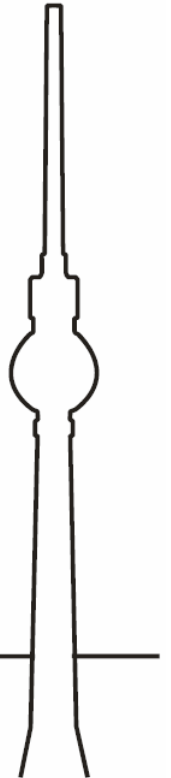
Brian McNeil,
Department of Physics,
University of Strathclyde, UK

David Dunning & Neil Thompson
CCLRC Daresbury Laboratory, UK

Brian Sheehy,
Sheehy Scientific Consulting



fel 06
BERLIN



August 27 – September 1, 2006
Berlin and Dresden, Germany



Outline

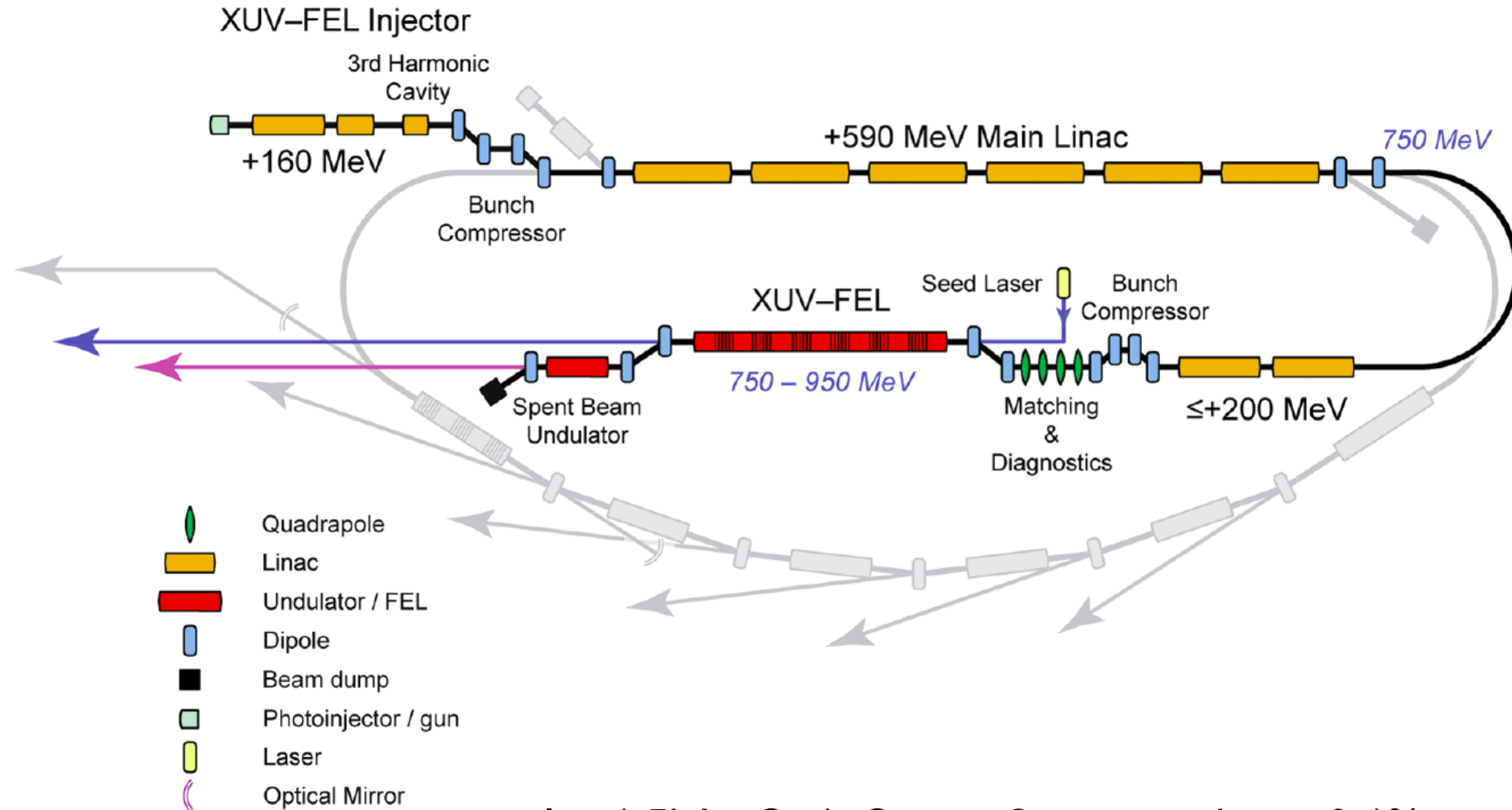
- Brief look at conceptual design of 4GLS XUV-FEL
- Benefits of amplifier seeding
- 4GLS HHG conceptual design seeding scheme
- Modelling HHG seeding
- Simulations of XUV-FEL conceptual design
- Further design work for XUV-FEL



Conceptual Design of 4GLS XUV-FEL



XUV-FEL layout



$$I_{pk}=1.5\text{kA} ; Q=1\text{nC} ; \varepsilon_n = 2 \text{ mm-mrad} ; \sigma_v=0.1\%$$



XUV-FEL parameter summary

General	
FEL design	High Gain Amplifier
Seeding mechanism	HHG source
Photon output	
Tuning Range	$\sim 8 - 100$ eV
Peak Power	$\sim 8 - 2$ GW
Repetition rate	~ 1 kHz
Polarisation	Variable elliptical
Min Pulse length FWHM	< 50 fs
Typical $\Delta f \Delta t$	~ 0.6
Max pulse energy	$400 \mu\text{J}$
Electron beam parameters	
Energy	750 - 950 MeV
Bunch Charge	1 nC
RMS bunch length	266 fs
Normalised emittance	2π mm mrad
RMS energy spread	0.1%
Undulator parameters	
Undulator Type	PPM & APPLE-II
No of Modules	8 & 5
Module lengths	~ 2 m
Period	45 mm & 51 mm
Focusing	FODO
Minimum magnetic gap	~ 10 mm



XUV FEL Undulator Tuning: Variable Polarisation Options

● OPTION 1: baseline design

- ✱ **10-100eV**

- ✱ 10mm min gap

- ✱ W1 planar 45mm period

- ✱ W2 APPLE II 51mm period

“Gaps” are undulator gaps

⇒ Vacuum gaps will be

~ undulator gaps – 3mm

● OPTION 2: extension to 6eV

- ✱ **6-100eV**

- ✱ 8mm min gap

- ✱ W1 planar 45mm period

- ✱ W2 APPLE II 53mm period

● OPTION 3: extension to 6eV

- ✱ **6-100eV**

- ✱ 10mm min gap

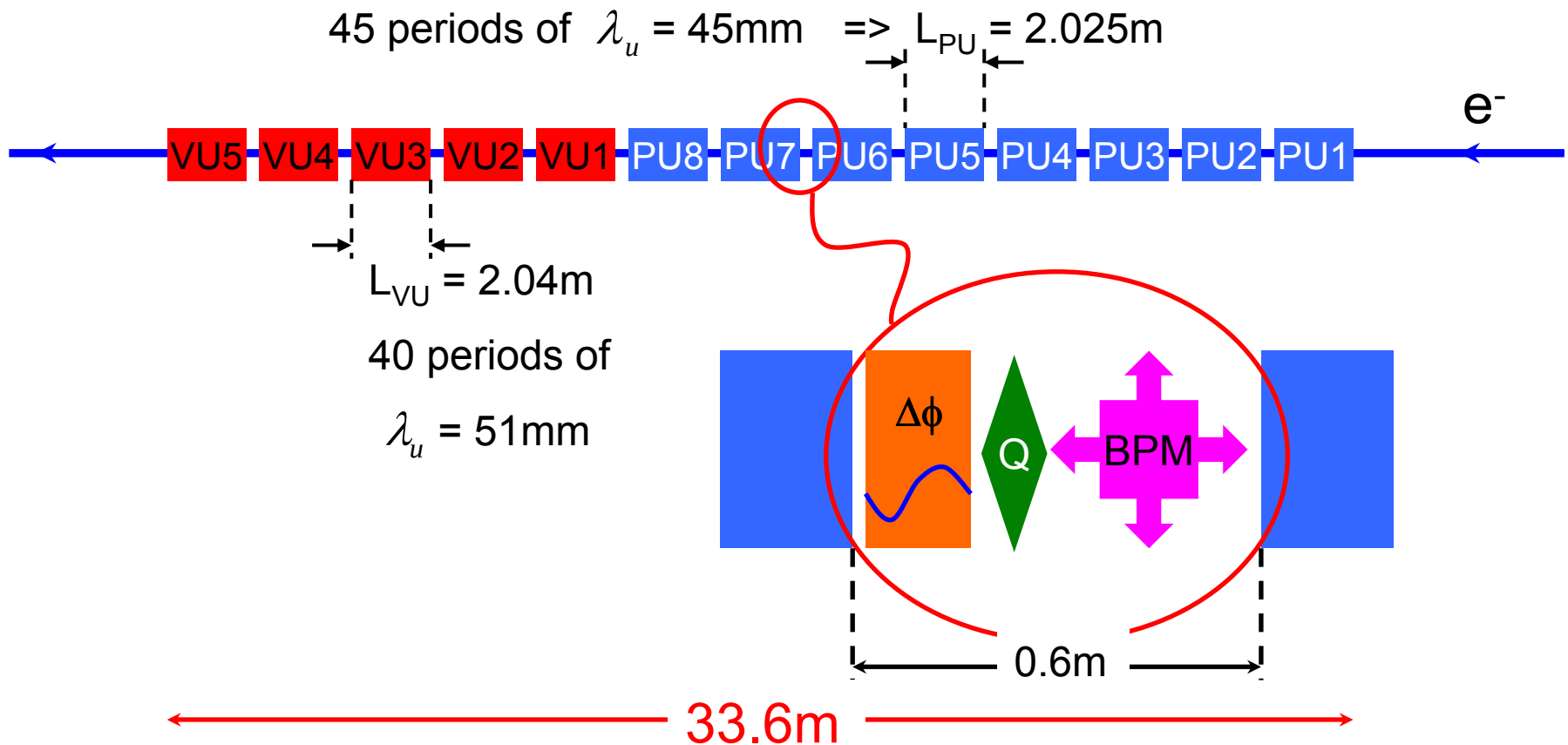
- ✱ W1 planar 45mm period

- ✱ W2 APPLE II 57mm period



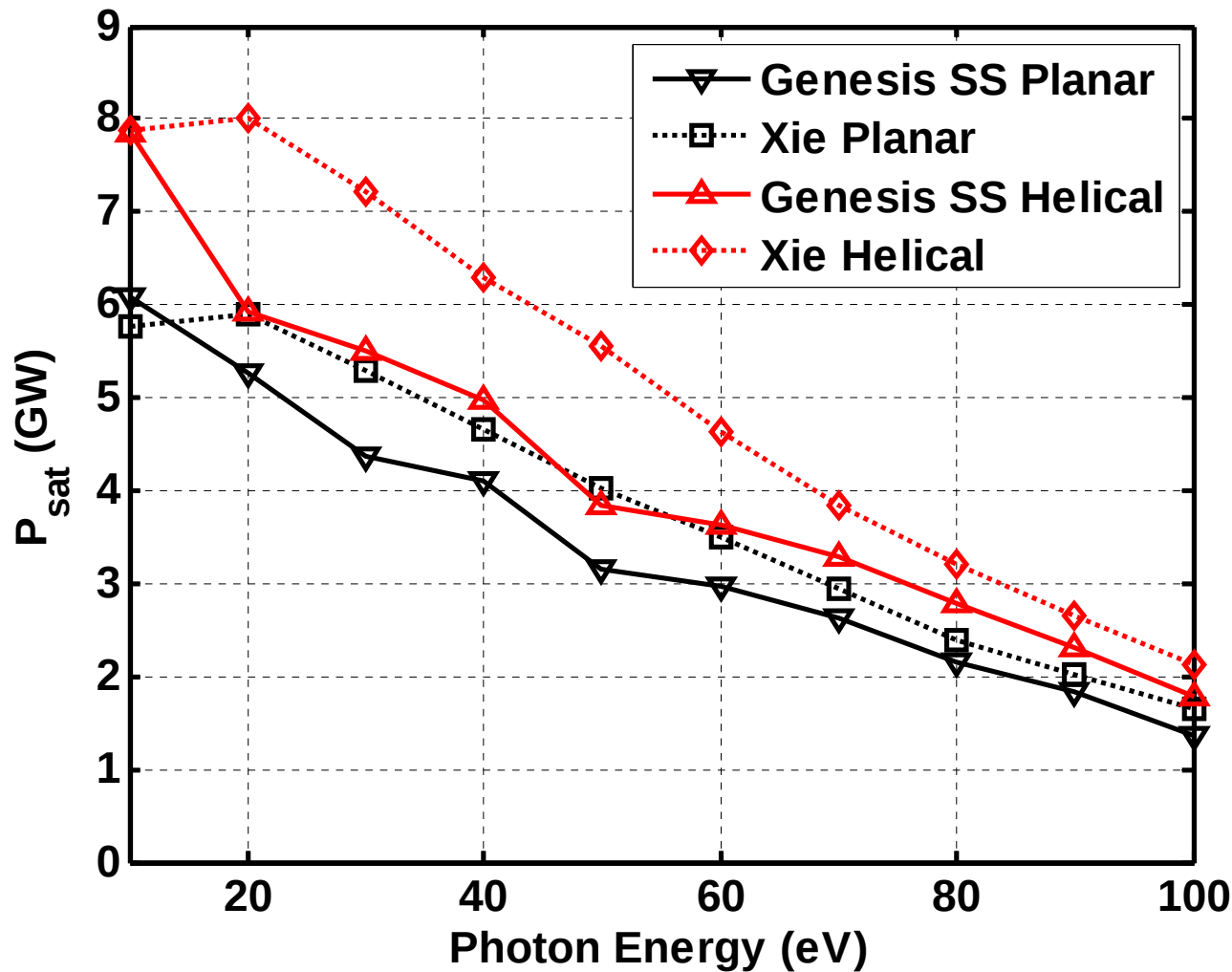
Baseline Option 1*

(10-100eV; 10mm gap)



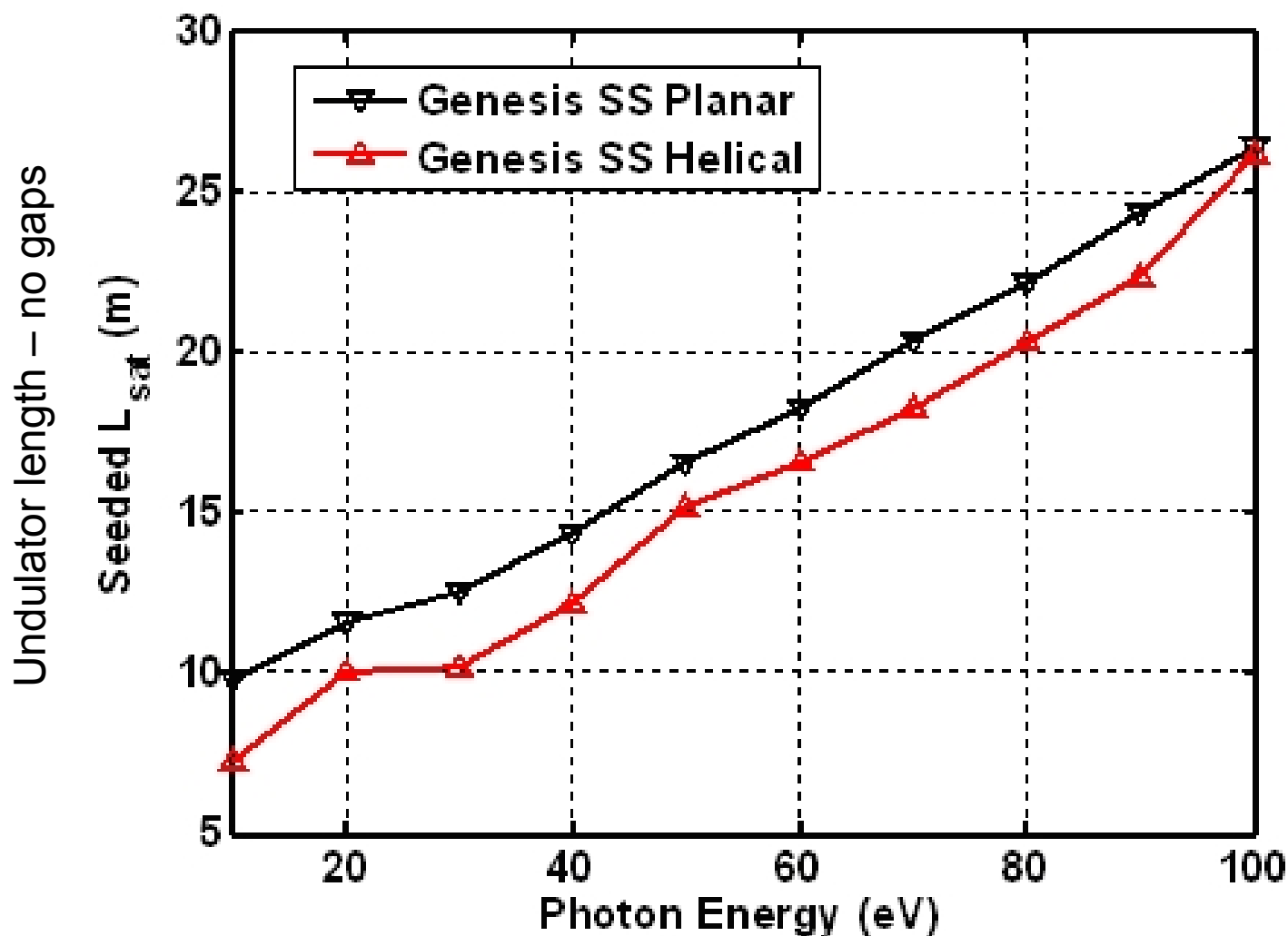


SS Simulations 10-100eV





SS Simulations 10-100eV



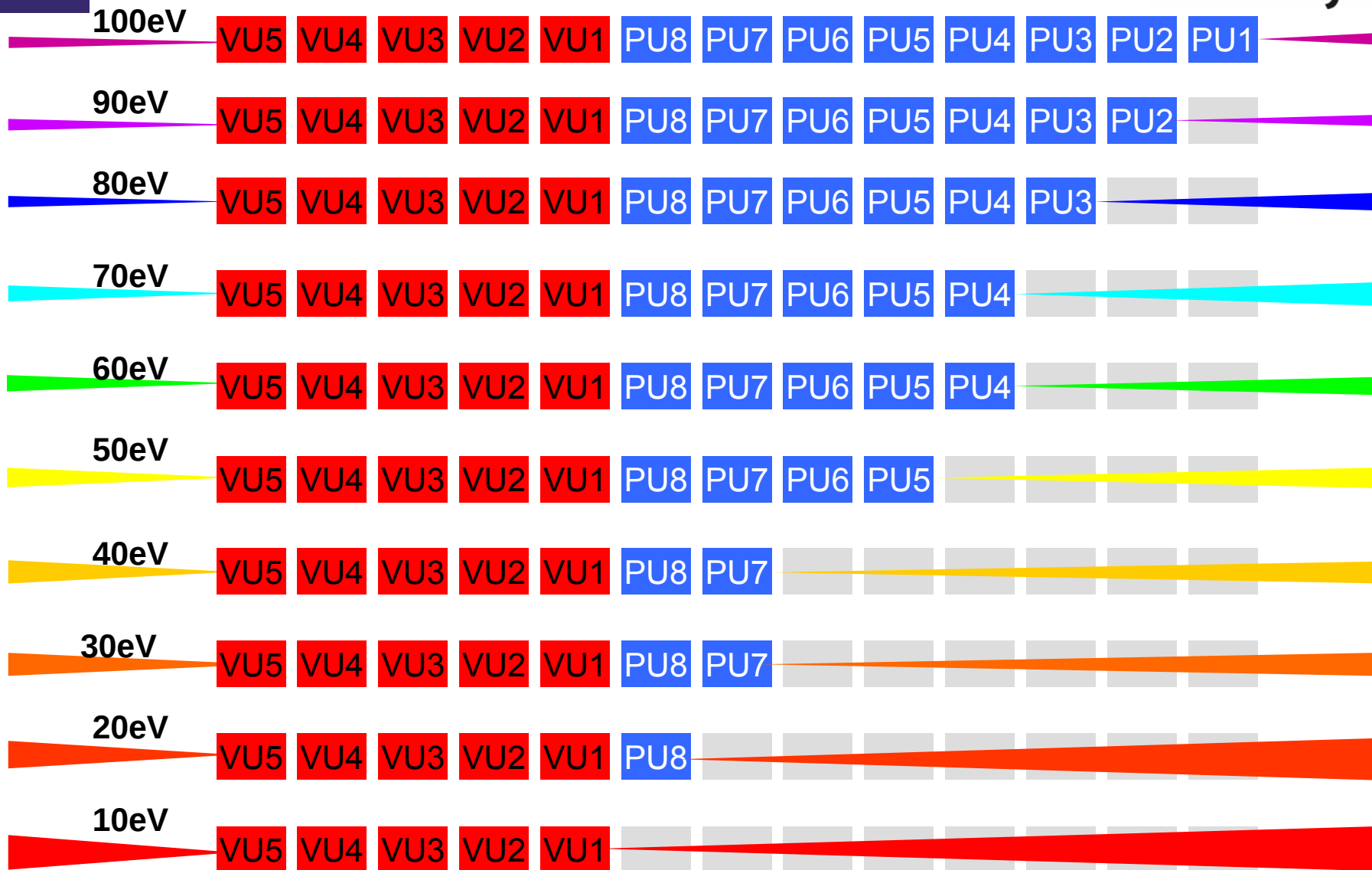


Benefits of amplifier seeding

- Improvement in the temporal coherence (and hence spectral brightness)
- Shot-to-shot reproducibility (same shape)
- Shot-to-shot stability (same power)
- Shorter pulses – not determined electron pulse
- Shorter interaction length (& cost)
- Amplify exotic pulses for post-amplification manipulation
- Seed may be used as source in its own right
- **Requires spatio-temporal synchronisation with electrons**

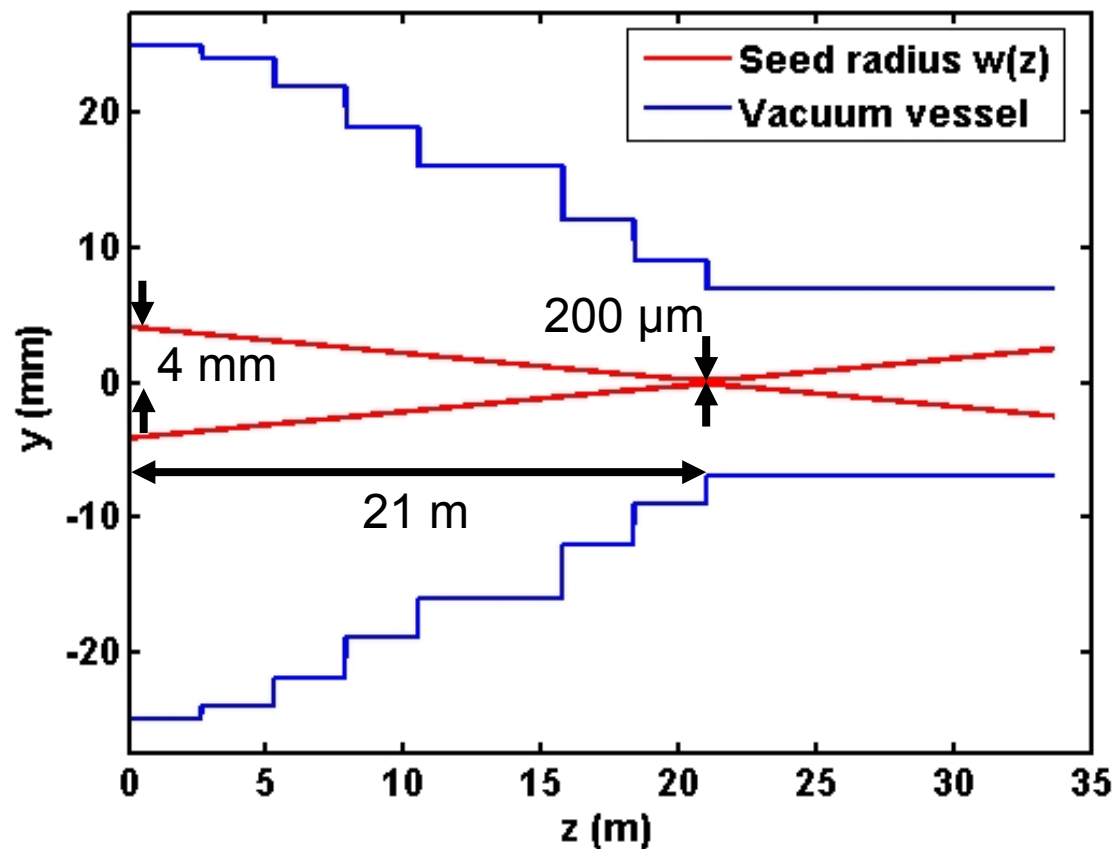


4GLS HHG conceptual design seeding scheme

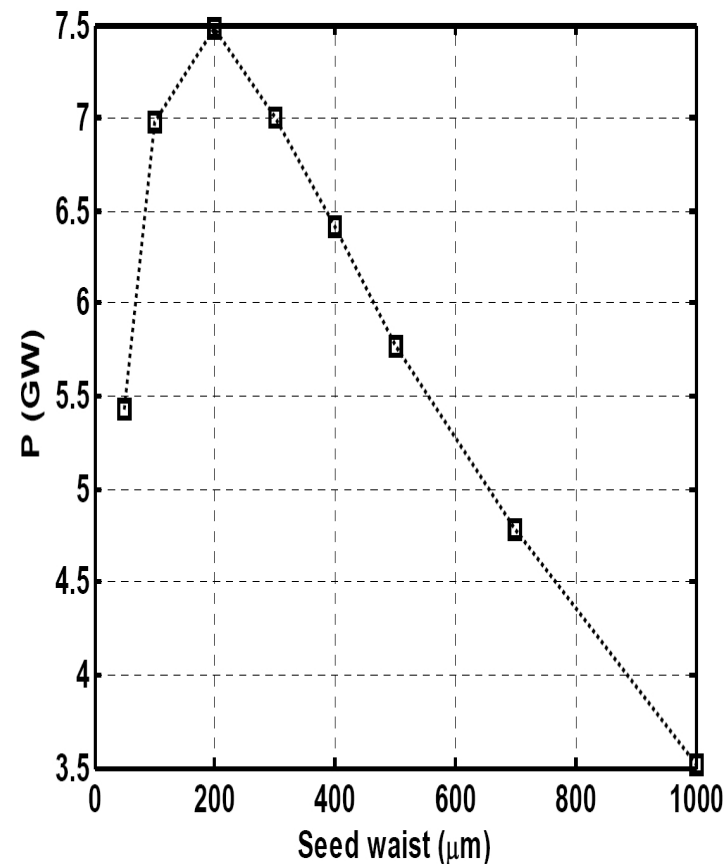




Tapered vacuum vessel



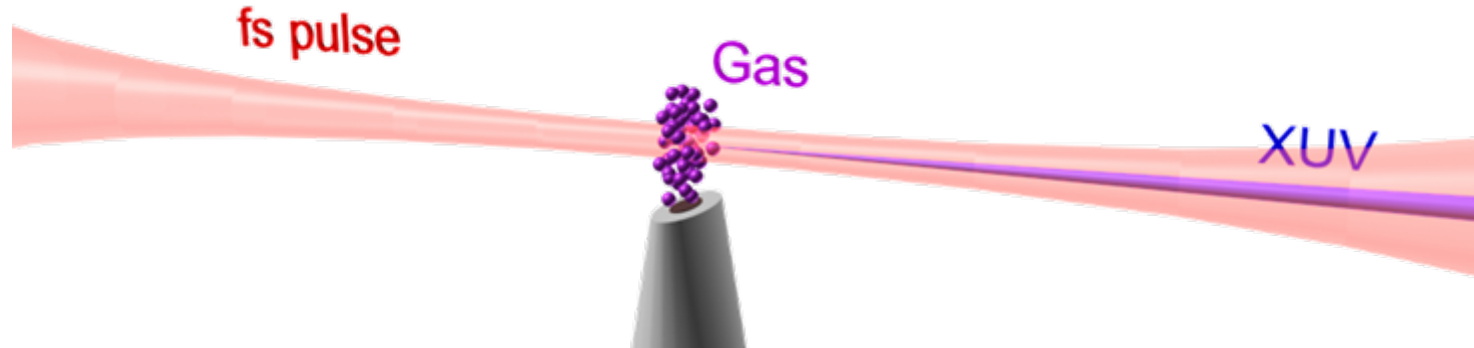
Saturation power



$$w(z) = w_0 \sqrt{1 + \left(\frac{z - z_0}{Z_R} \right)^2} \quad Z_R = \pi w_0^2 / \lambda$$



HHG Seed Generation



Good introductory source material at:

Henry C. Kapteyn et al., "EUV 'Photonics' of High-Harmonic Generation and Applications "

ICFA Beam Dynamics Mini Workshop (Future Light Source Working Group)

Workshop on the Physics of Seeded Free Electron Lasers

MIT, 17-19 June 2004 (<http://mitbates.mit.edu/xfel/conference.htm>)

and:

John W.G. Tisch, "Alternative Coherent X-ray Sources",

STI Round-Table Meeting DESY,

Hamburg 22-24 June 2004

(xfel.desy.de/content/e154/upload/upload_file/Meetings/STI_June04/Tisch.pdf)



HHG Energy Spectrum*

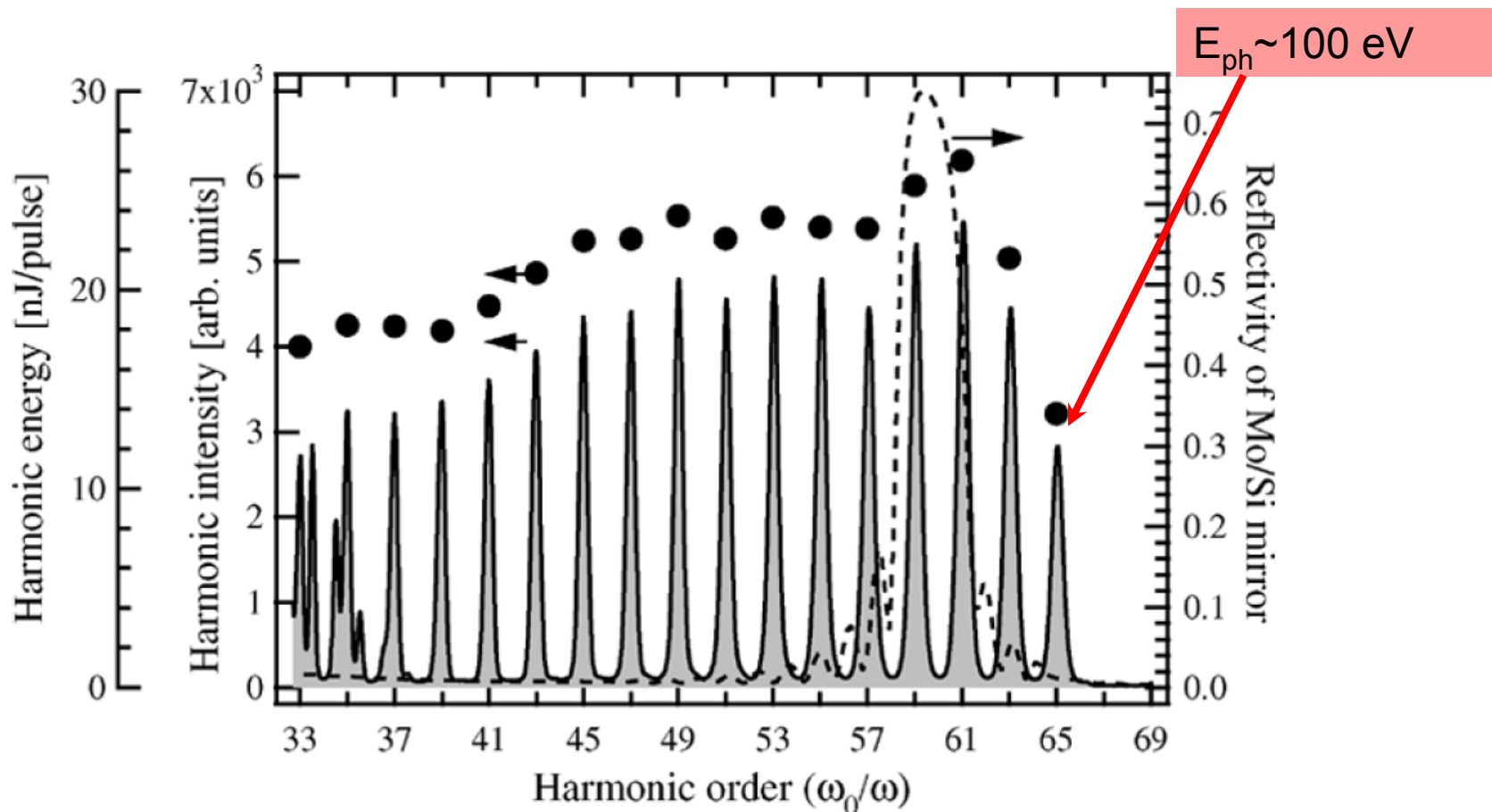
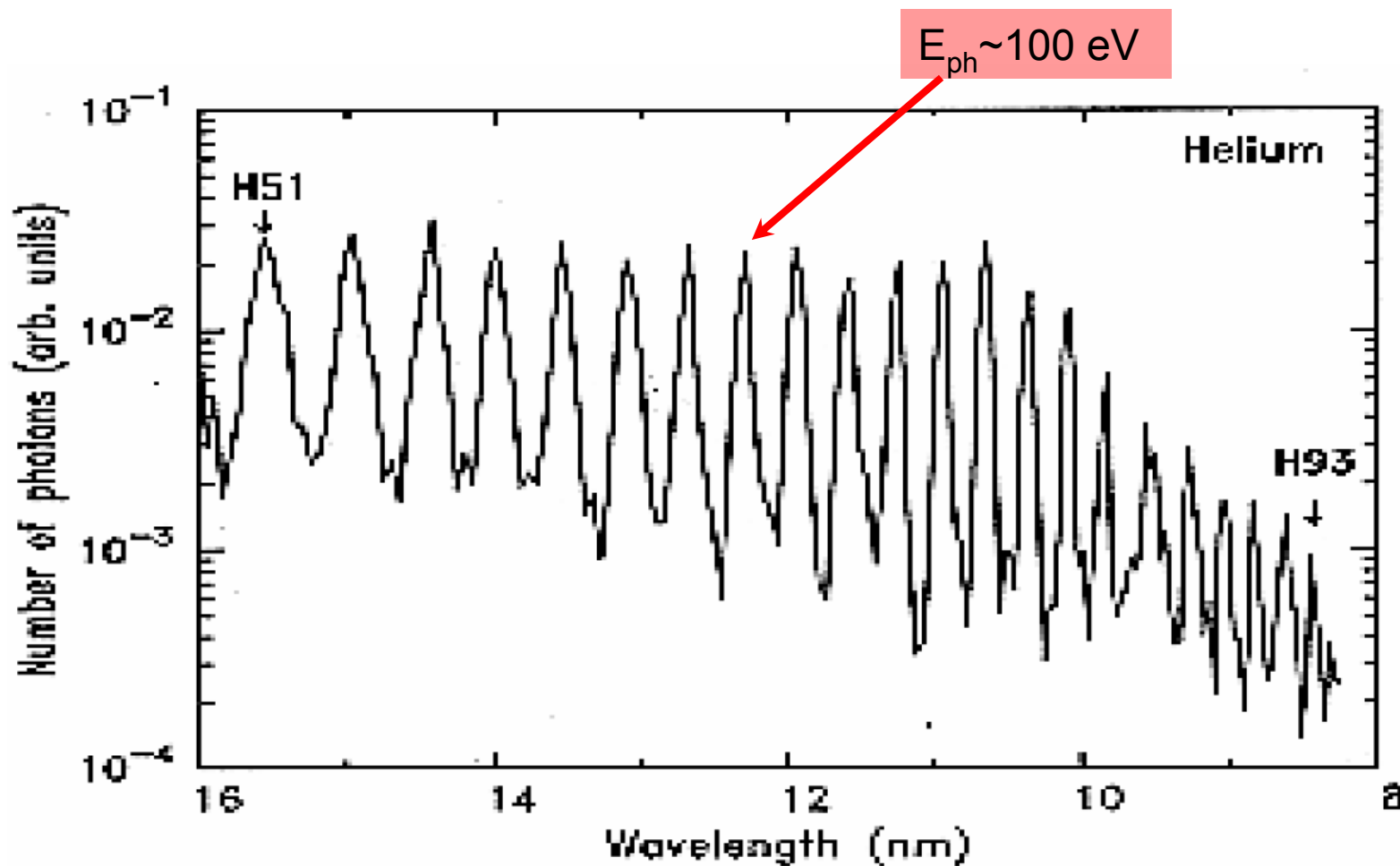


Fig. 9. Experimentally obtained harmonic spectra and energy yield in Ne gas. Dotted line shows the reflectivity of Mo–Si multilayer mirror.



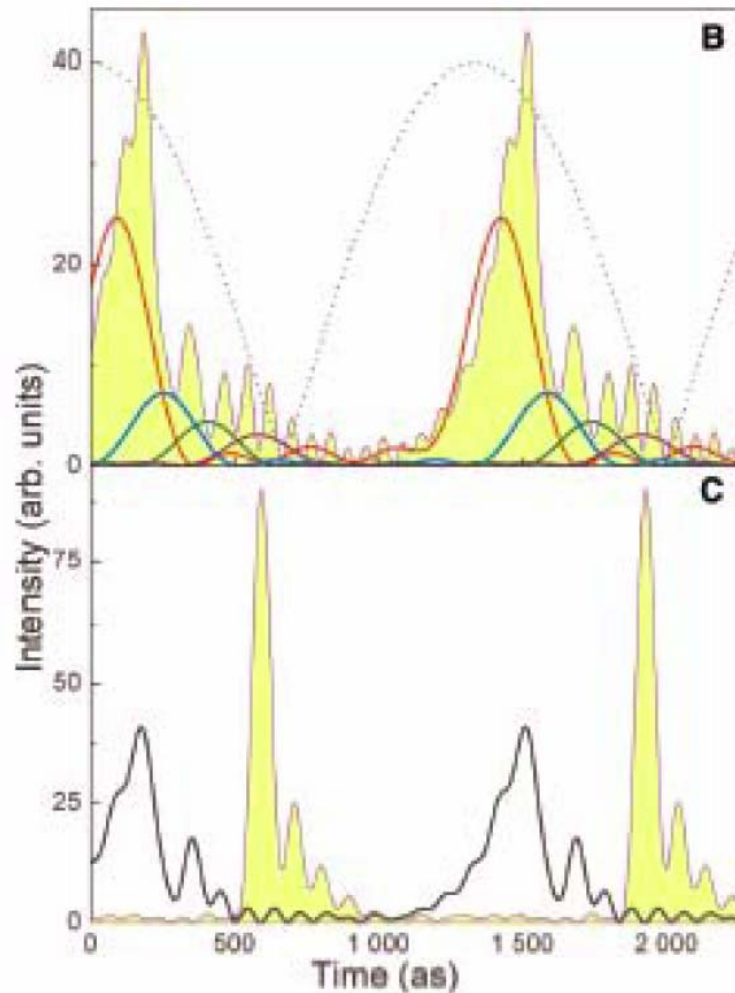
HHG Energy Spectrum*



*Y. Mairesse et al., Science **302**, 1540 (2003)



HHG temporal structure

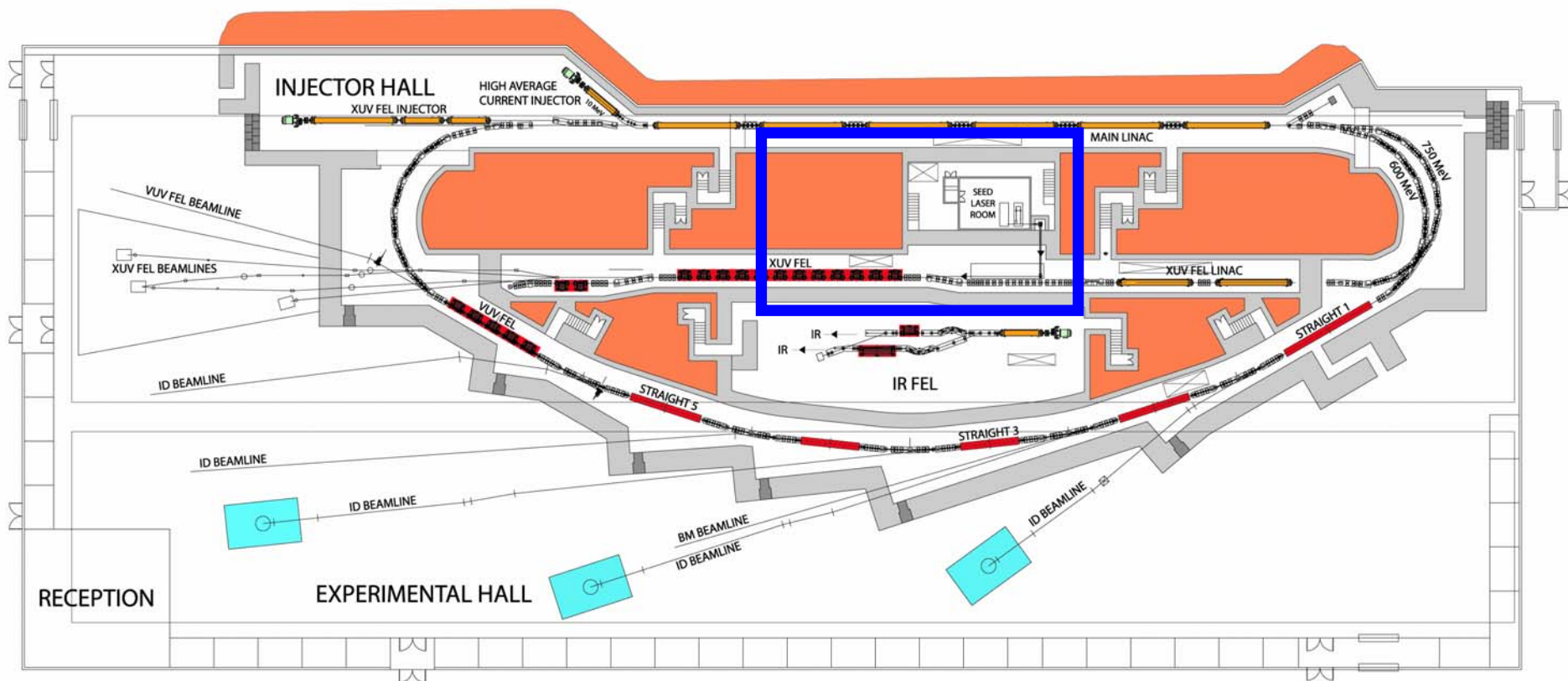


Temporal structure contains a series of atto-second spikes separated by $\frac{1}{2}$ of the drive laser period



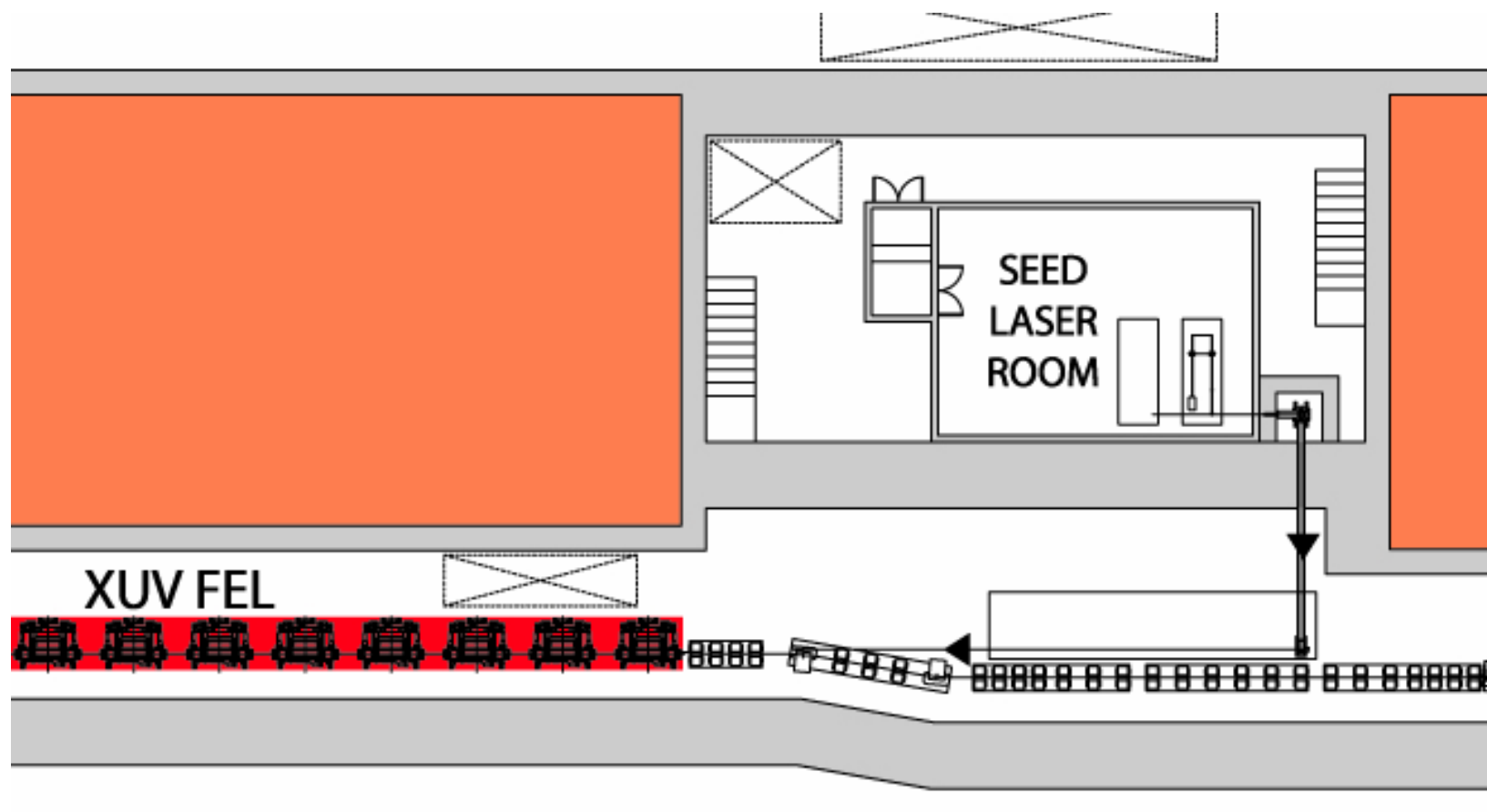
University of
Strathclyde

4GLS layout



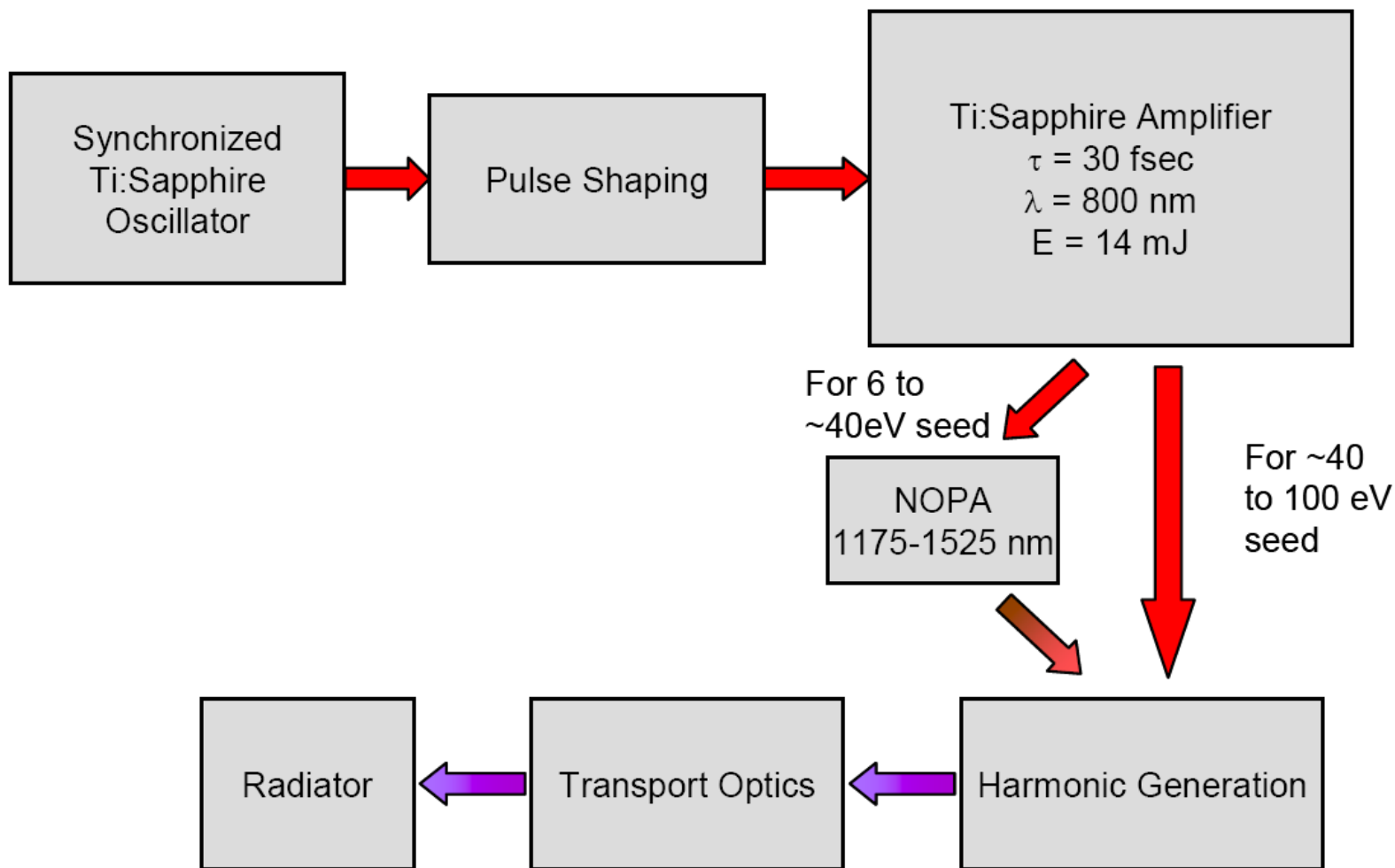


Seed laser room





HHG seed for XUV-FEL



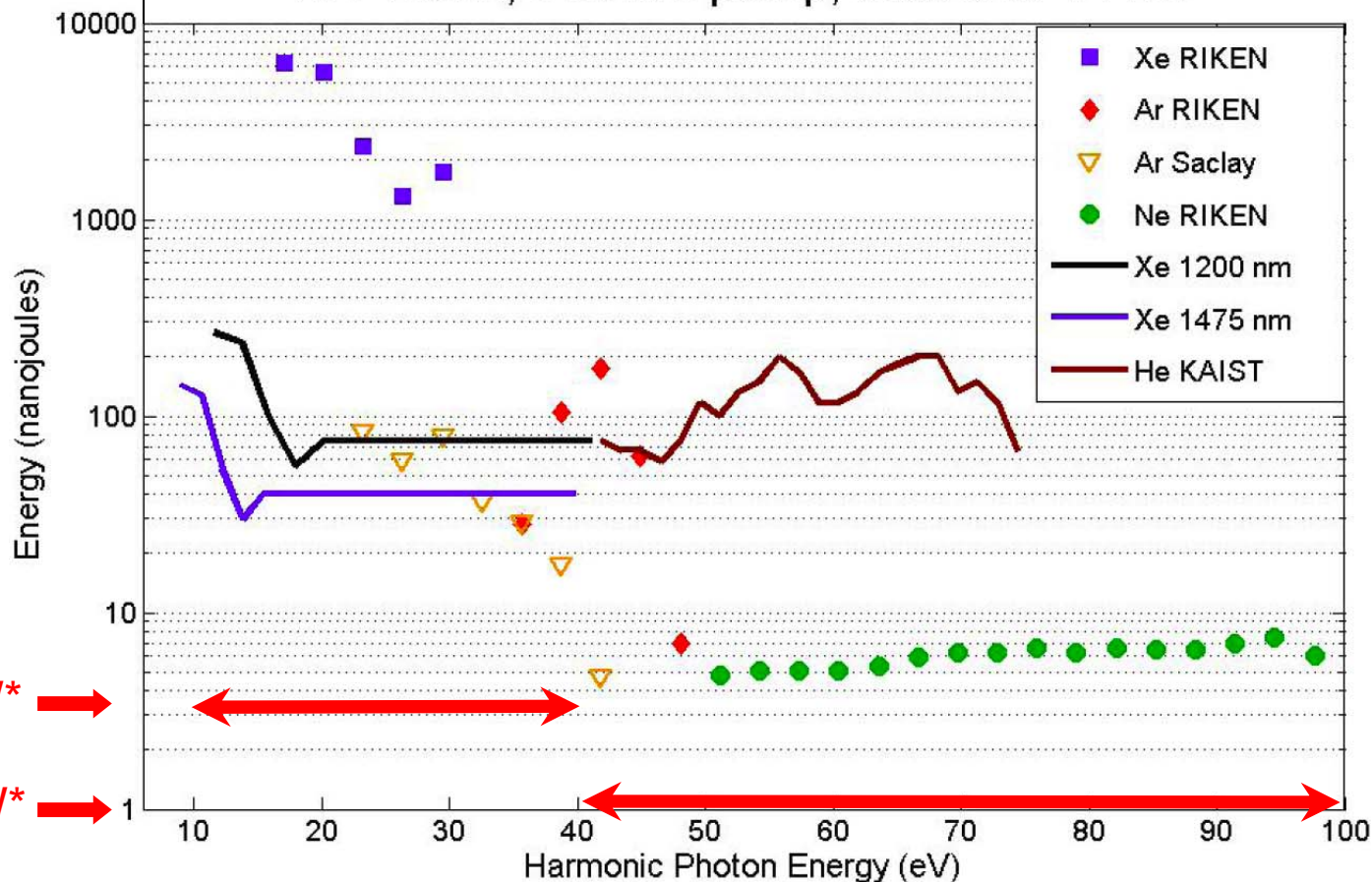
See: B. Sheehy et al., this conference: MOPPH067



~3 GW X

Seed energy

HH Yields, 800 nm pump, scaled to 14 mJ



*For 30 fs FWHM – The peak powers used for XUV simulations

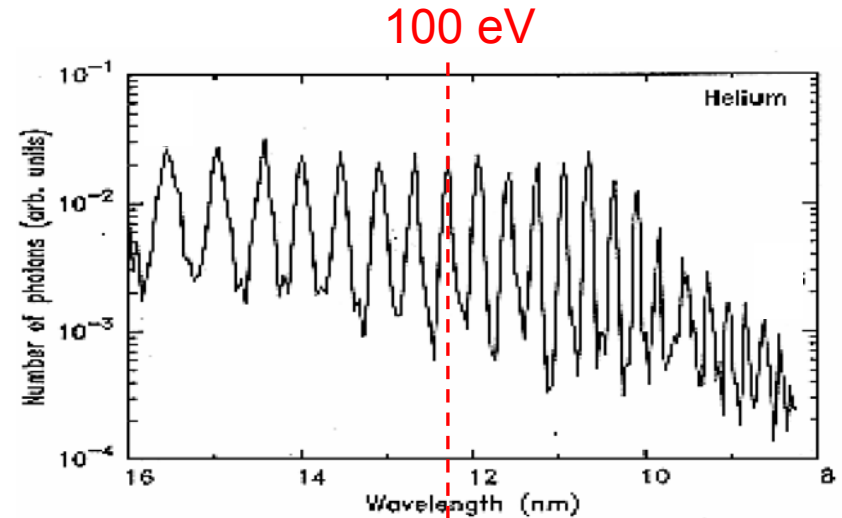


Modelling HHG seeding

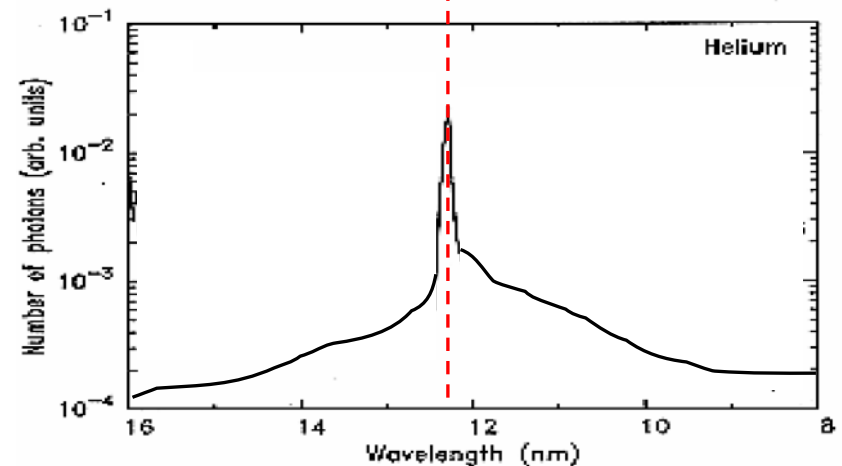


Seeding options

1) Un-filtered ?

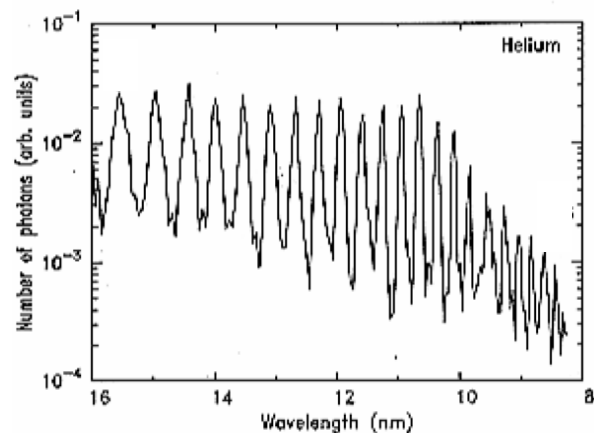


2) Filtered ?





Un-filtered expectations



1) Spectral:

FEL gain bandwidth < harmonic separation

$$\left. \frac{\Delta\lambda_n}{\lambda_n} \right|_{\text{HWHM}} \sim \rho$$

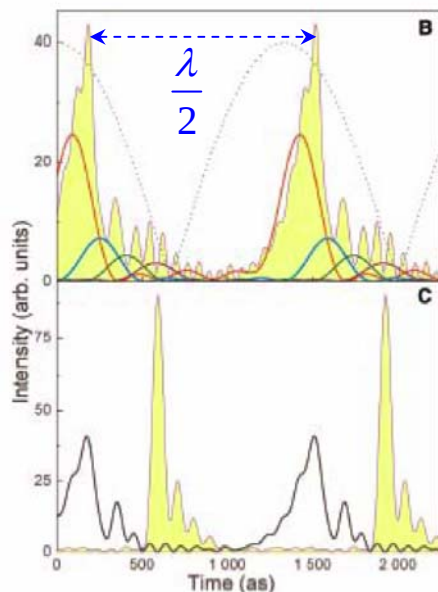
$\sim \rho$

$$\rho < \frac{1}{n}$$

$$\frac{\Delta\lambda_n}{\lambda_n} \approx \frac{1}{n}$$

$$\rho \approx 2 \times 10^{-3}$$

$$n^{-1} \approx 1.5 \times 10^{-2}$$



2) Spatio-temporal:

Distance between spikes < cooperation length

$$\frac{\lambda_1}{2}$$

$$\rho < \frac{1}{2\pi n}$$

$$\frac{\lambda_n}{4\pi\rho} = \frac{\lambda_1}{4\pi\rho n}$$



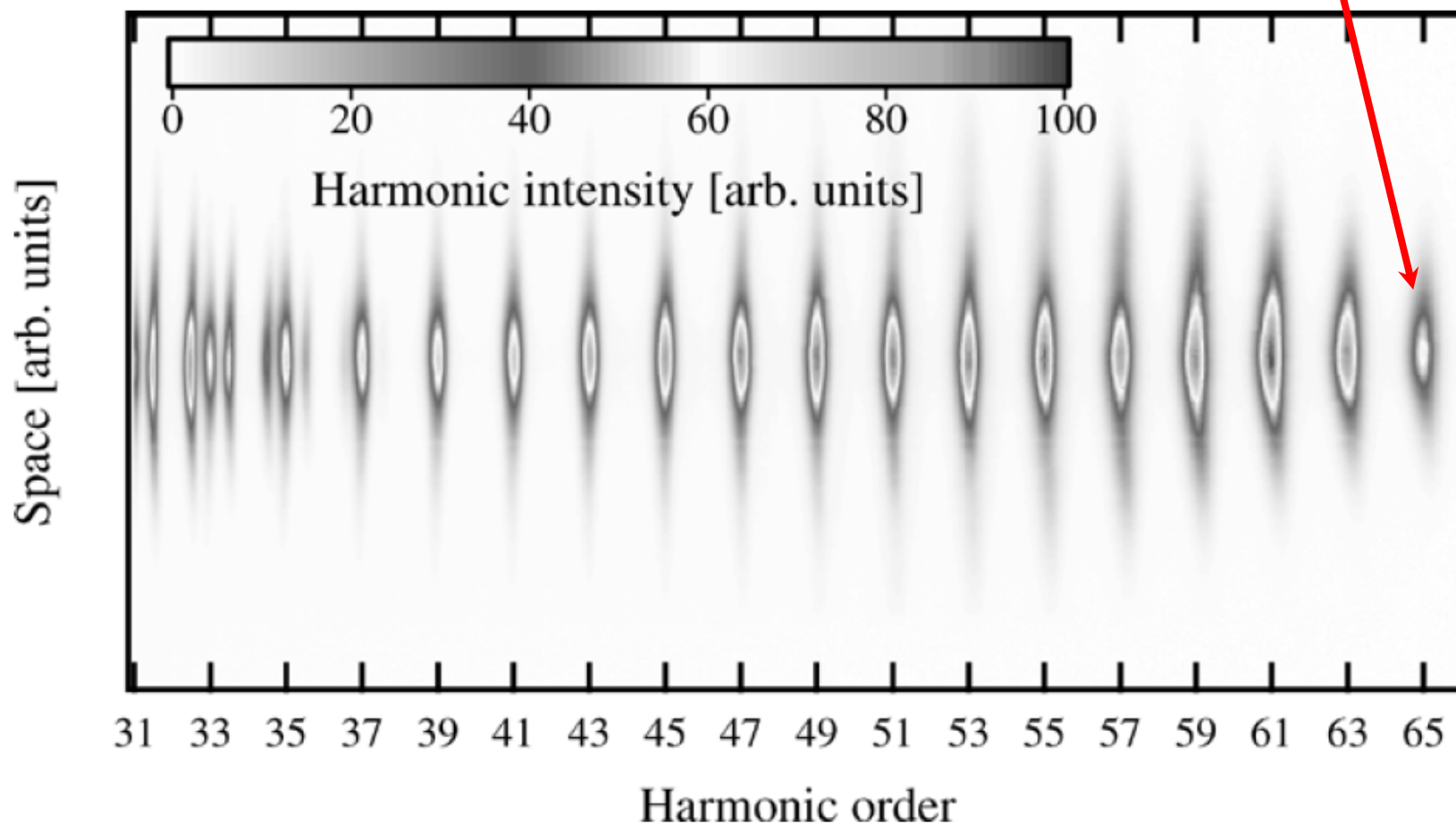
Not so important if 1) applies as need only consider one resonant harmonic => no short pulse structure

Can neglect all non-resonant harmonics? - Yes



HHG Spectral Intensity*

$E_{ph} \sim 100$ eV

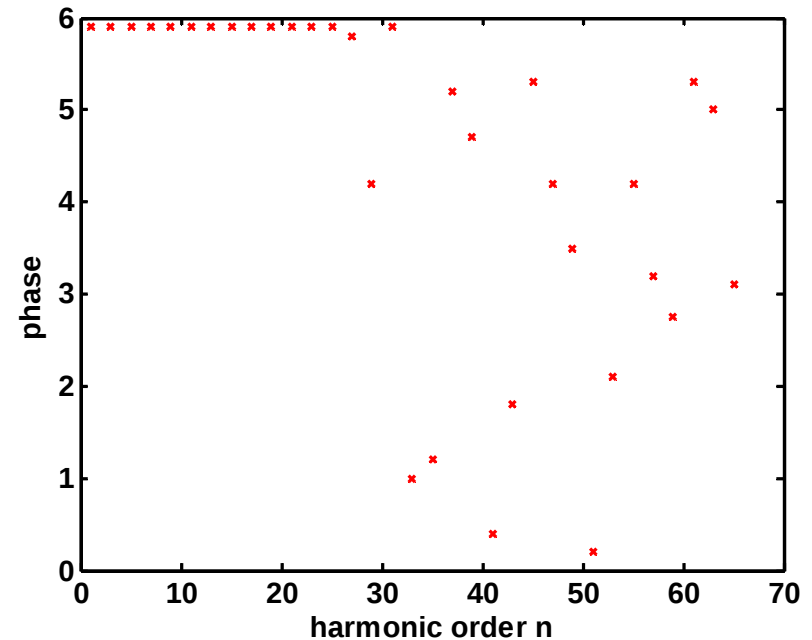
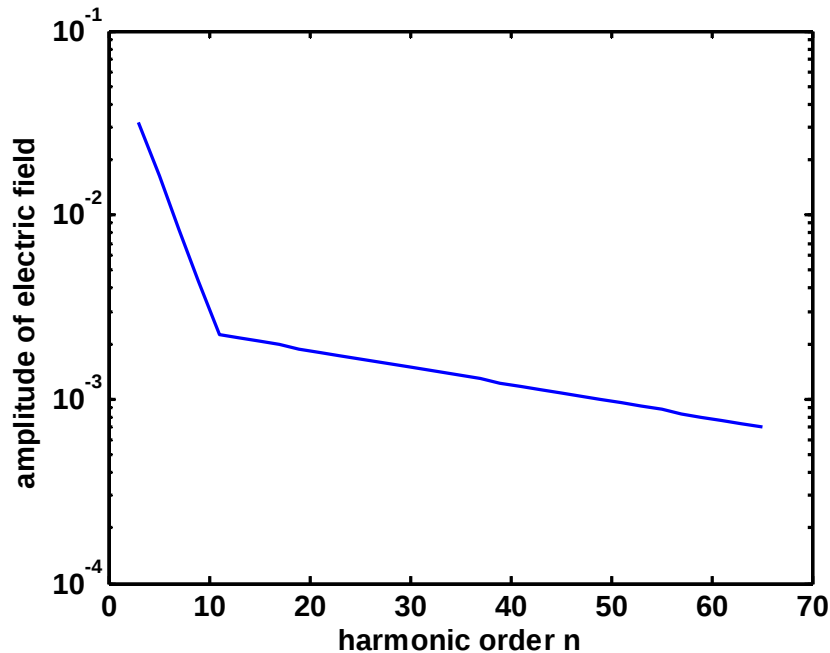


*Eiji J. Takahashi et al., IEEE Journal of Selected Topics in Quantum Electronics, **10**, 1315 (2004)



Simulated electric field

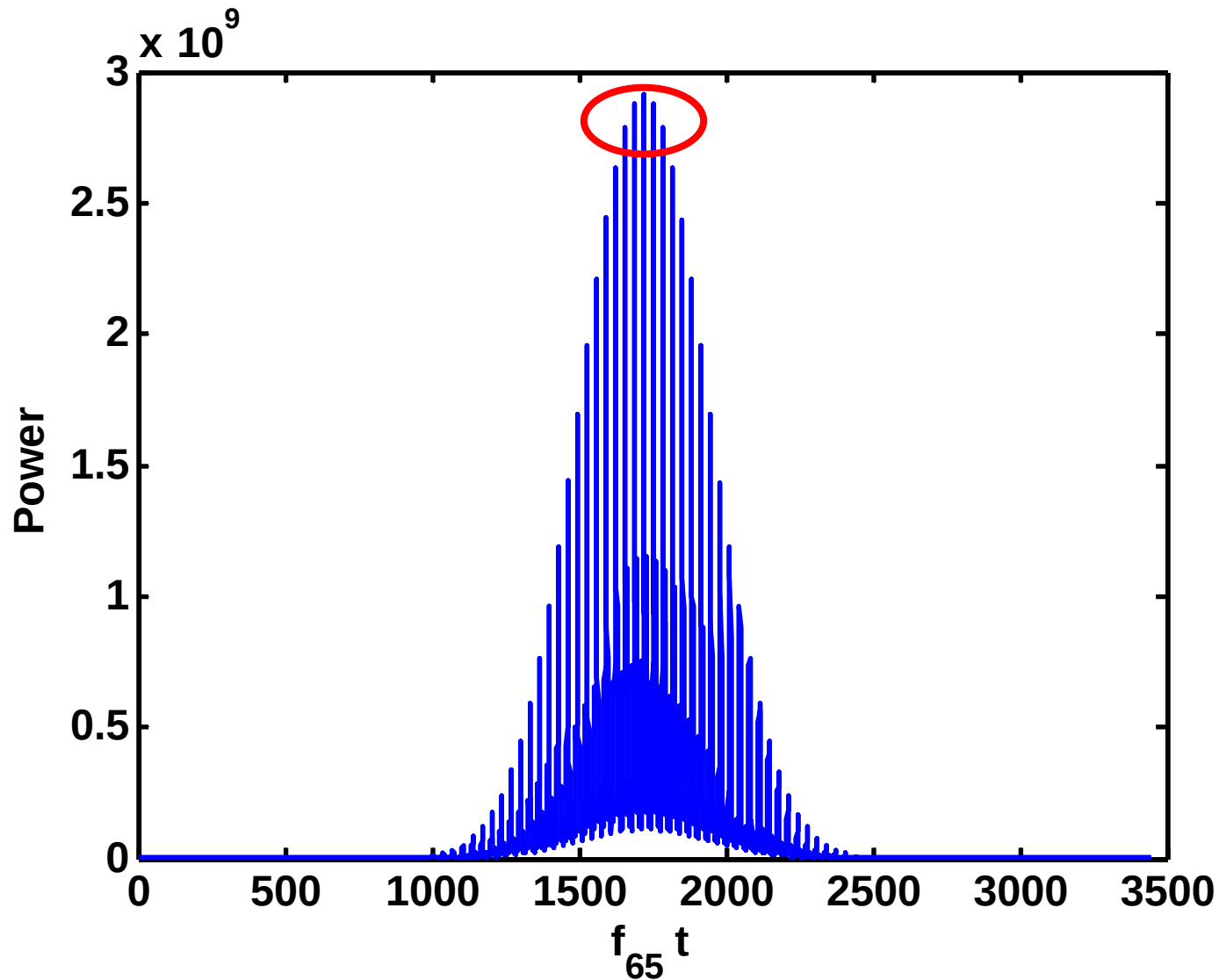
$$E = \sum_{\substack{n=3 \\ \text{odd}}}^{65} E_n e^{-i\omega_n t} e^{i\phi_n}$$



P. Antoine et al., Phys. Rev. A **56**, 4960 (1997)

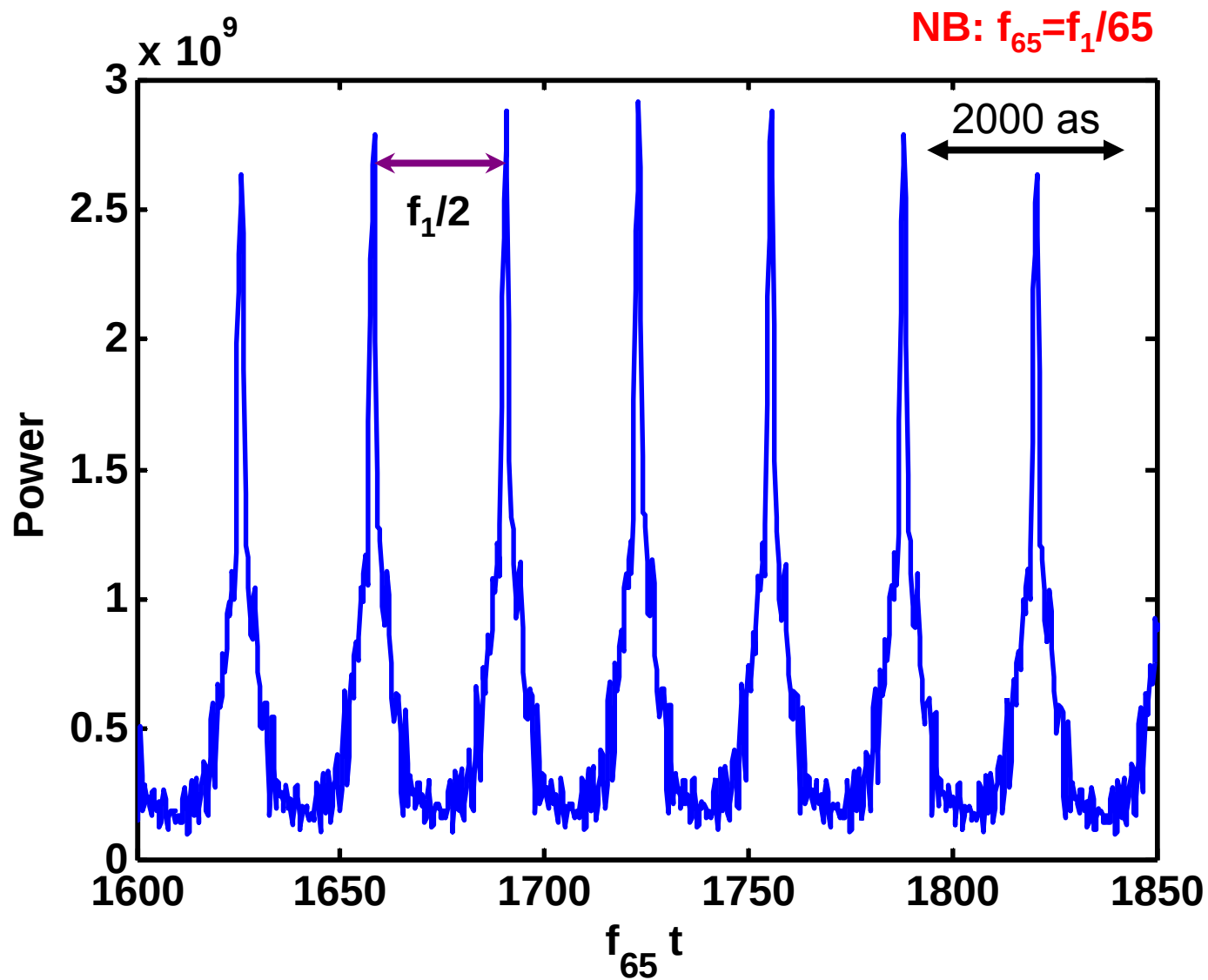


Simulated power



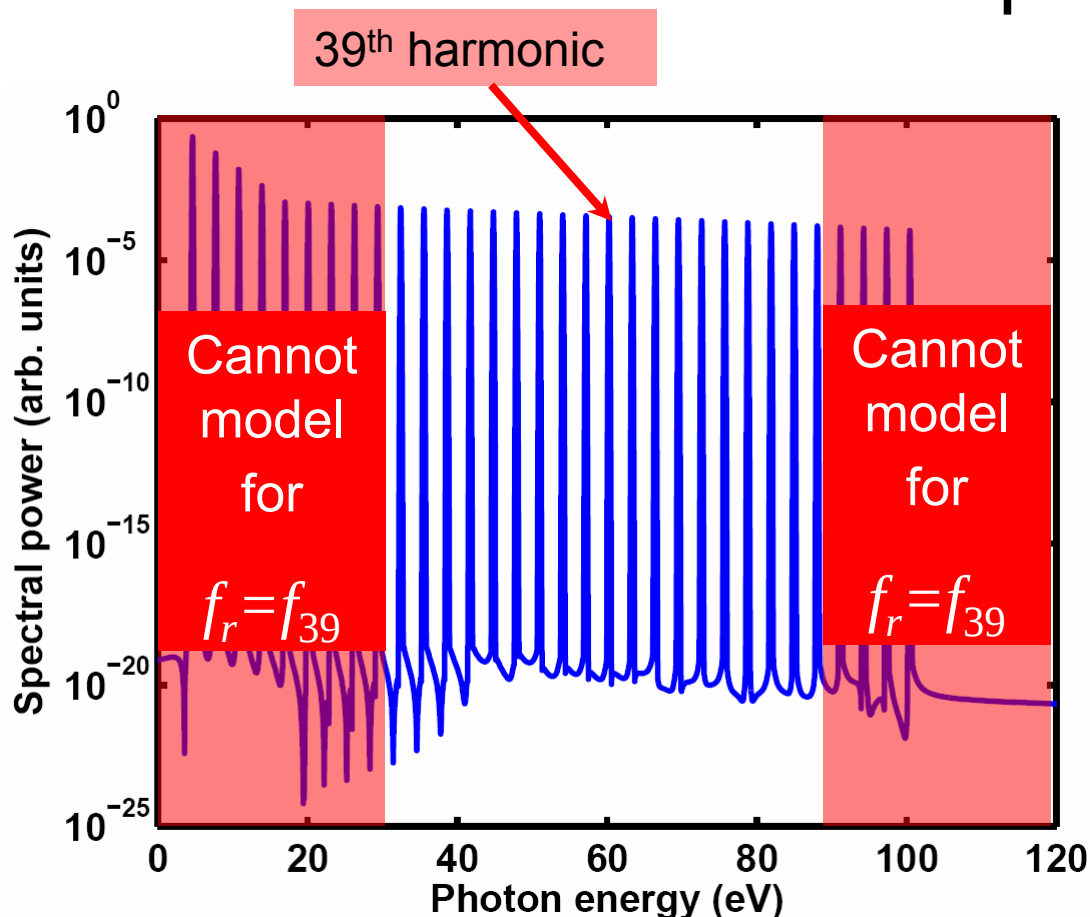


Simulated field detail





Simulated spectrum



For averaged FEL codes the minimum sample rate is:

$$\Delta t_s = f_r^{-1}$$

$$f_{Ny} = \frac{1}{2\Delta t_s} = \frac{1}{2} f_r \quad \text{Nyquist freq.}$$

Freq. range for non-aliasing:

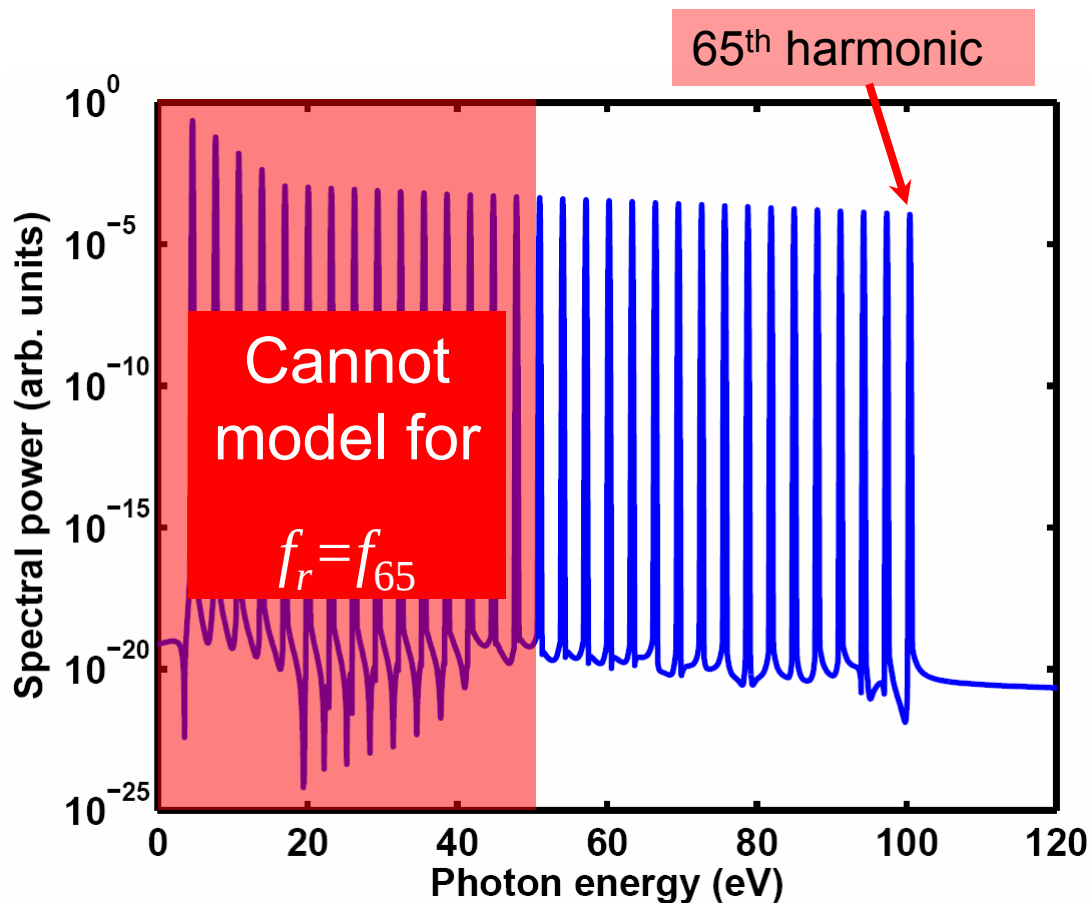
$$(f_r - f_{Ny}) > f_r > (f_r + f_{Ny})$$

=> Freq. range that e.g. Genesis can simulate properly without aliasing is:

$$\frac{f_r}{2} > f_r > \frac{3f_r}{2}$$



Simulated spectrum



For averaged FEL codes the minimum sample rate is:

$$\Delta t_s = f_r^{-1}$$

$$f_{Ny} = \frac{1}{2\Delta t_s} = \frac{1}{2} f_r \quad \text{Nyquist freq.}$$

Freq. range for non-aliasing:

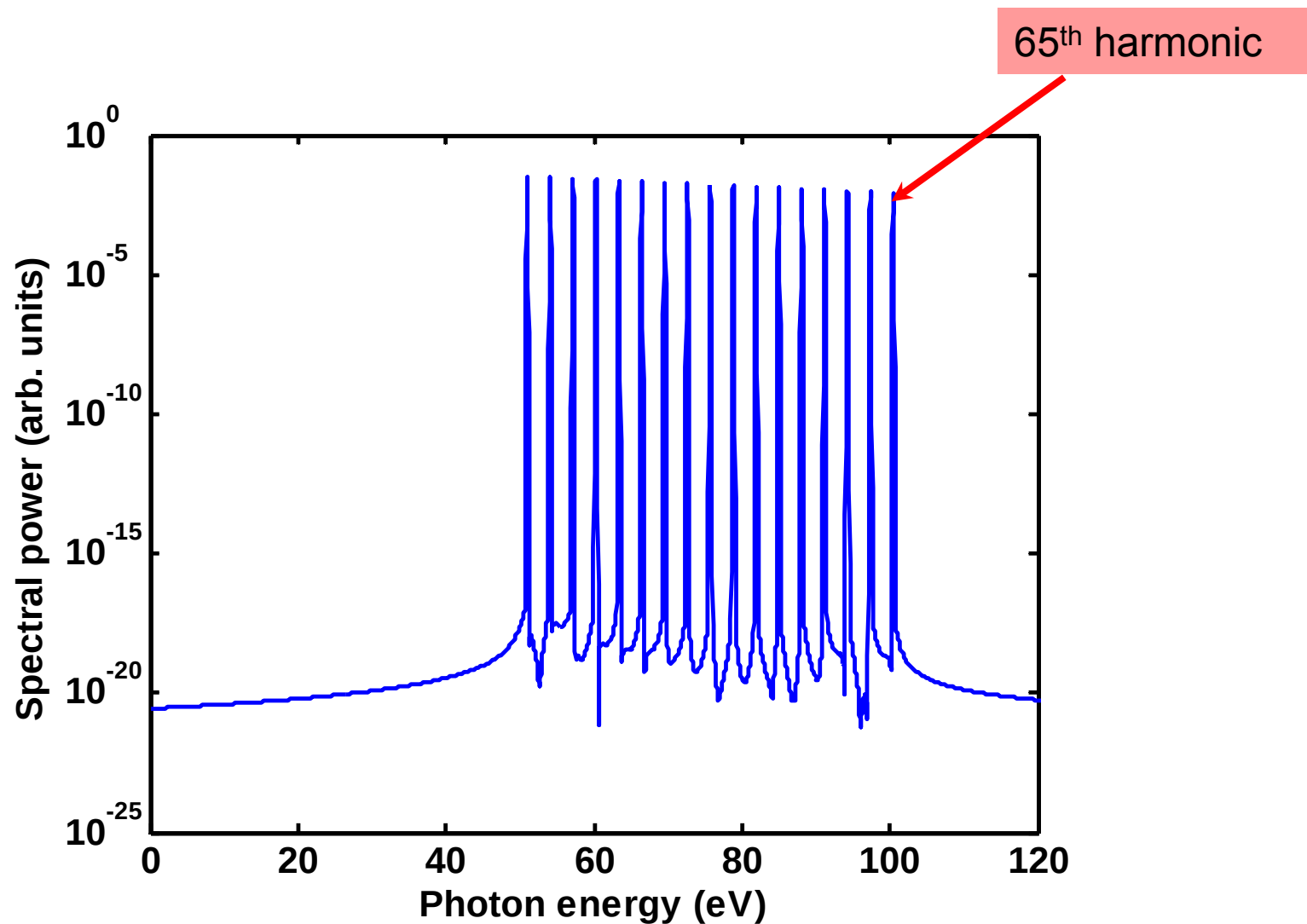
$$(f_r - f_{Ny}) > f_r > (f_r + f_{Ny})$$

=> Freq. range that e.g. Genesis can simulate properly without aliasing is:

$$\frac{f_r}{2} > f_r > \frac{3f_r}{2}$$



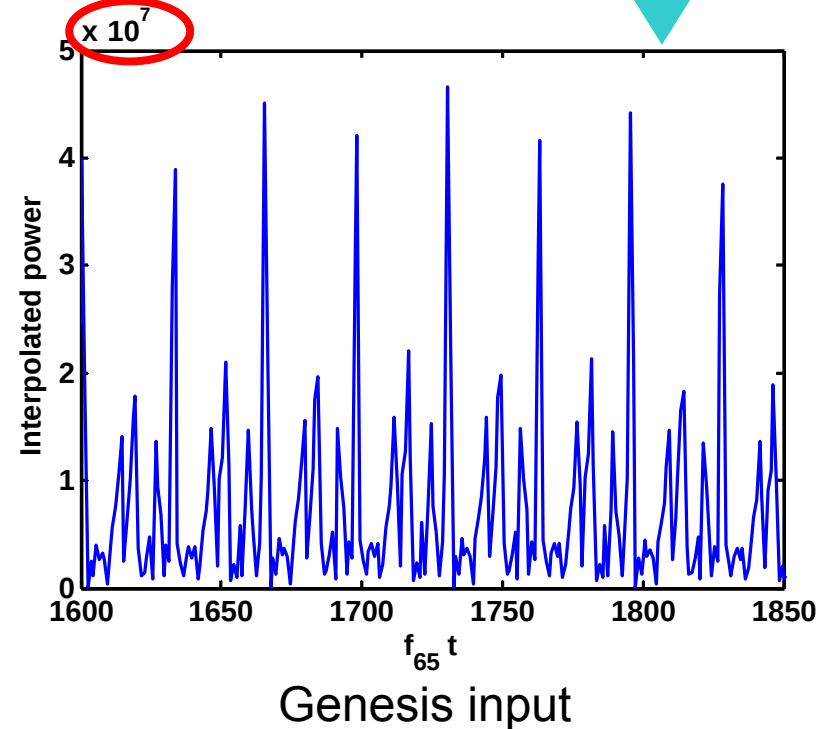
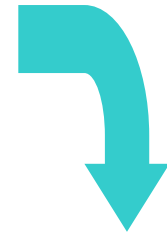
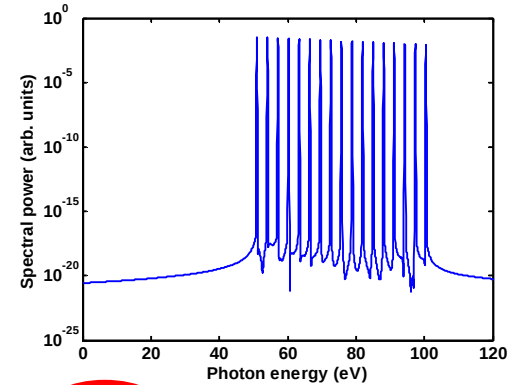
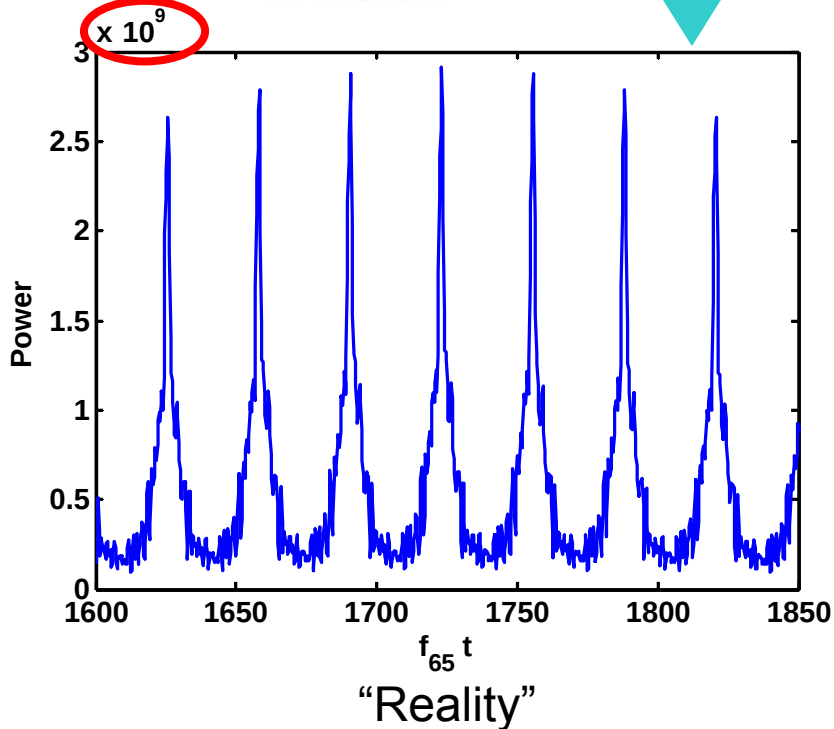
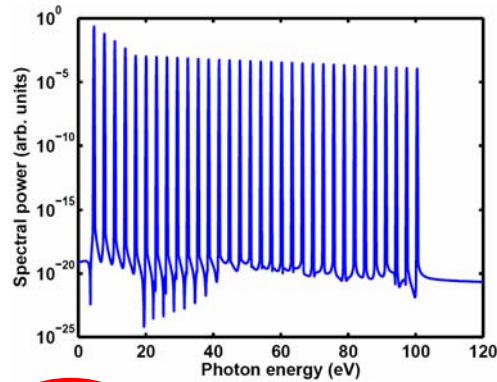
Simulated spectrum





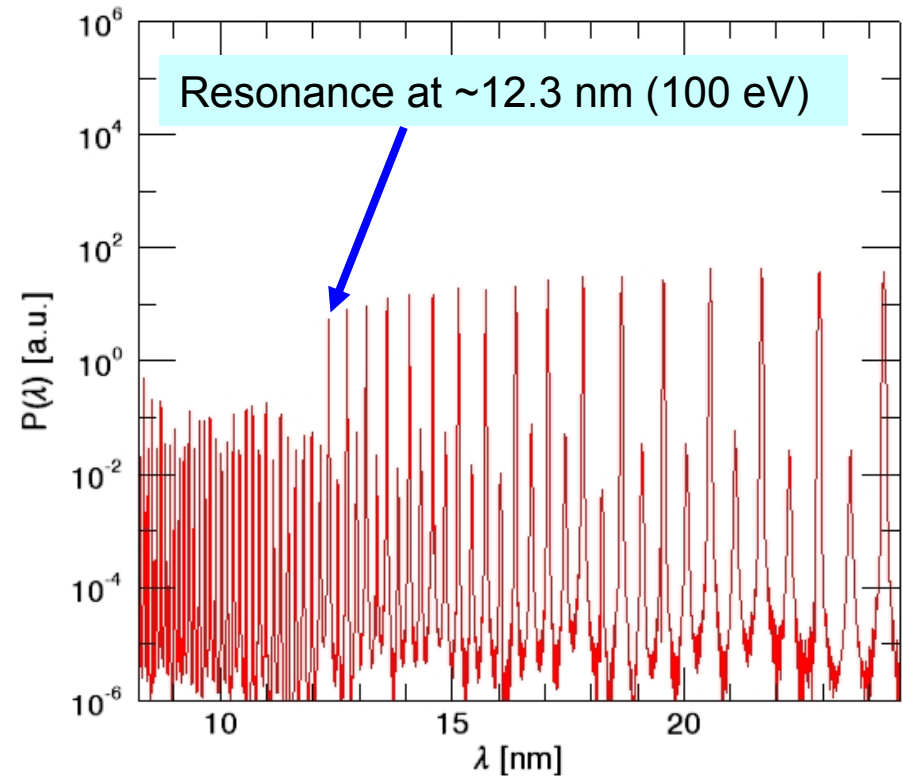
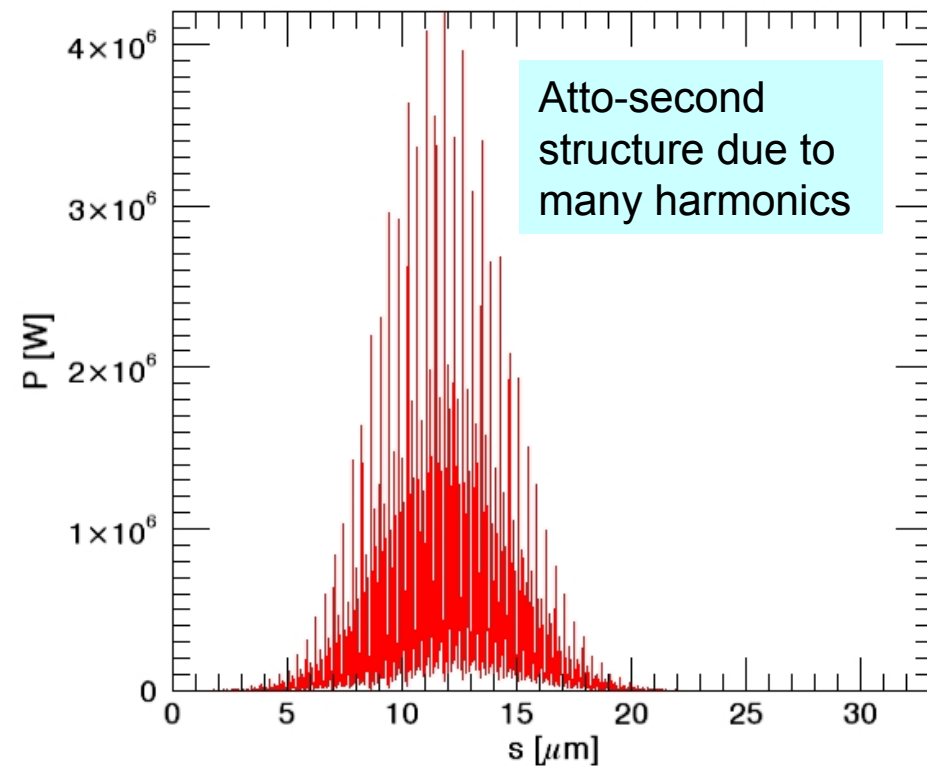
University of
Strathclyde

Reality check





Genesis simulation

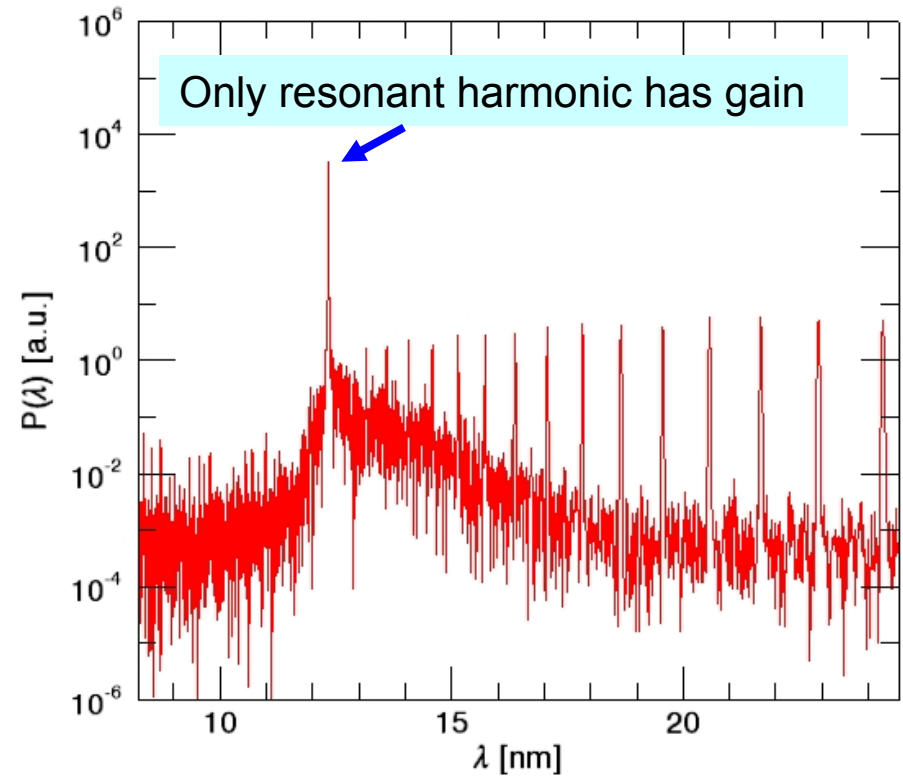
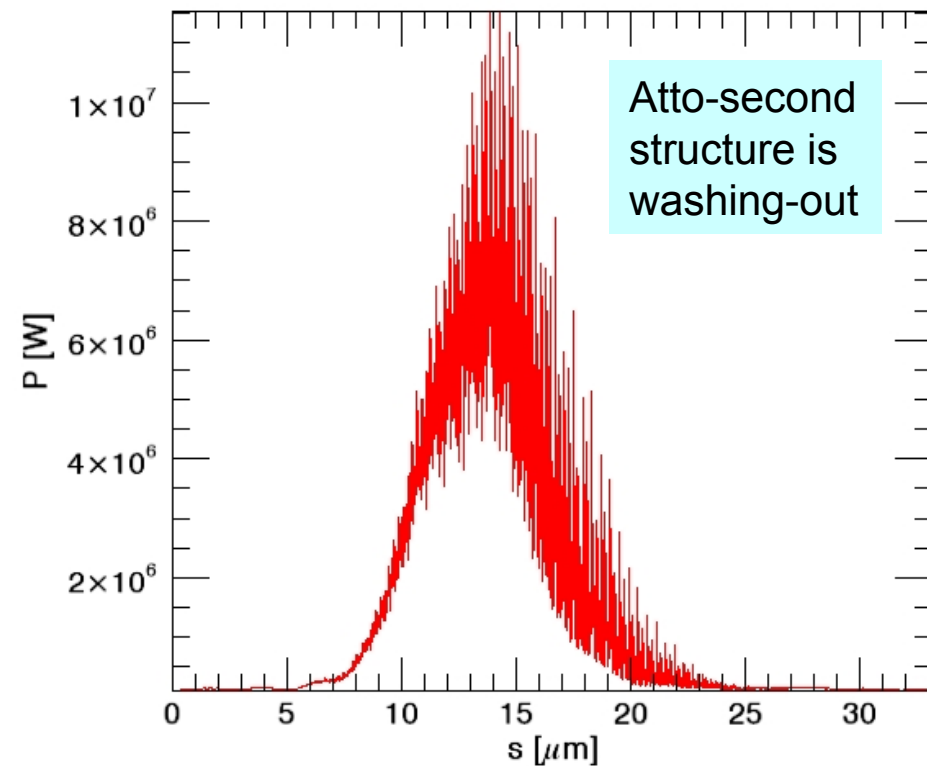


HHG seed field at entrance to FEL

XUV-FEL parameters but with uniform current $I=1.5$ kA



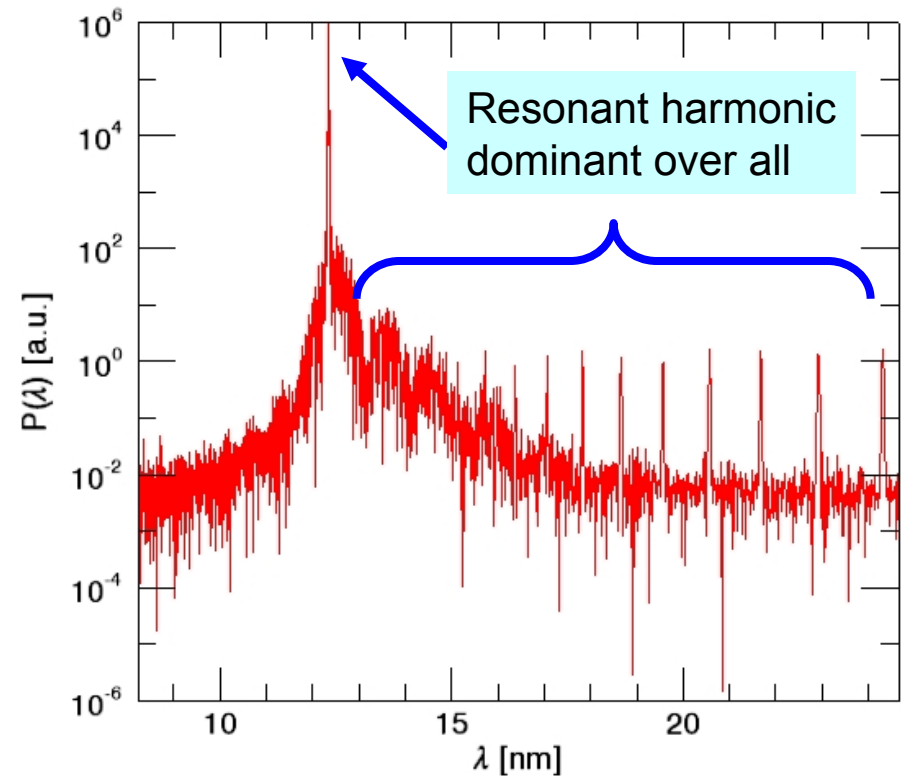
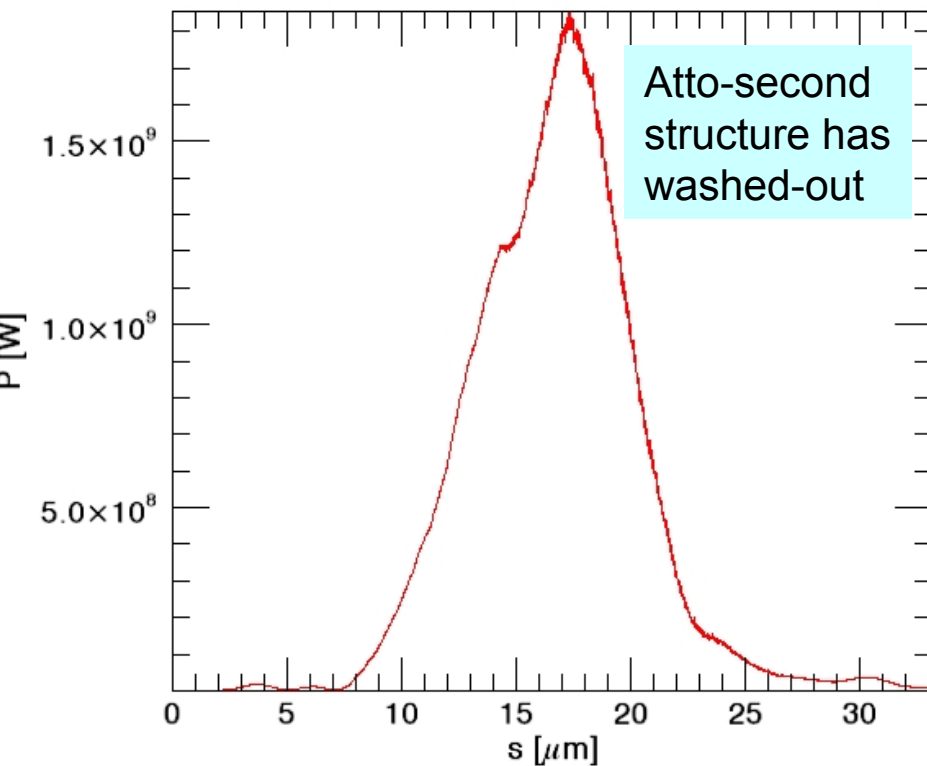
Genesis simulation



HHG seeded field at $z \sim 16$ m



Genesis simulation



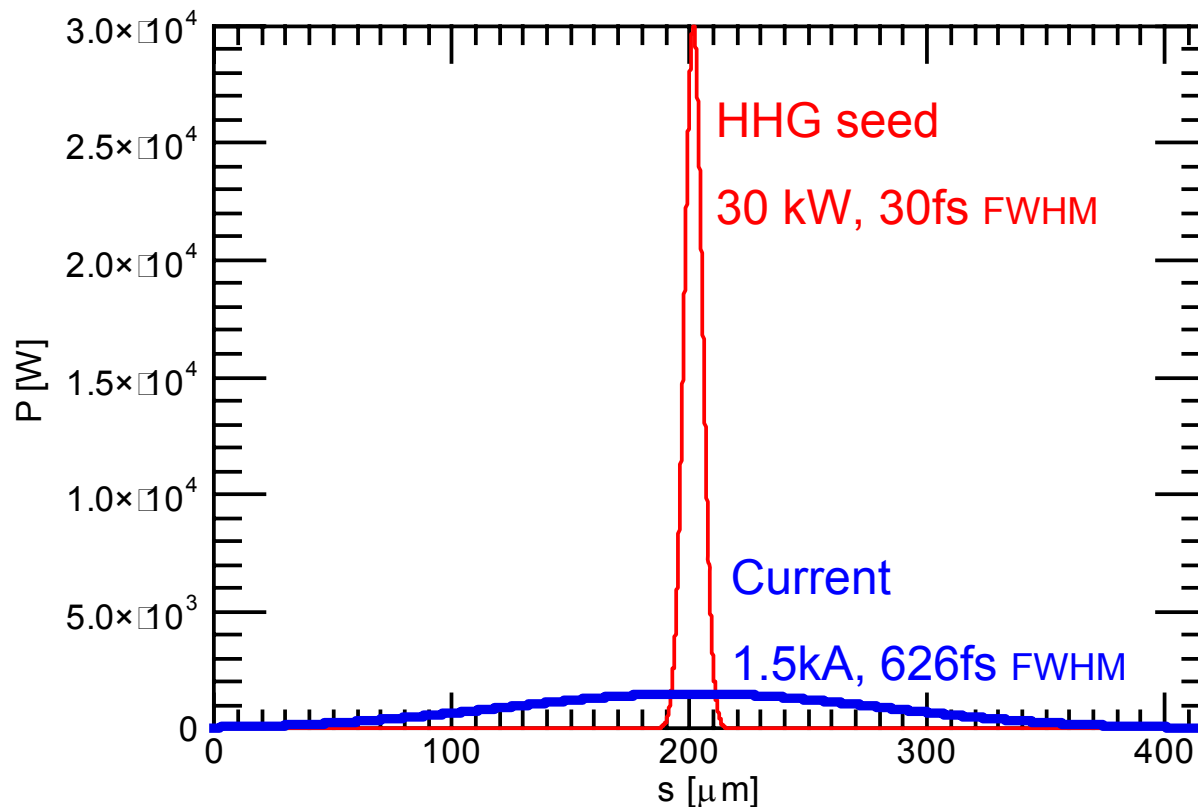
HHG seeded field at saturation $z \sim 32$ m



Simulations of XUV-FEL conceptual design XUV lasing - Pulses @ 100eV



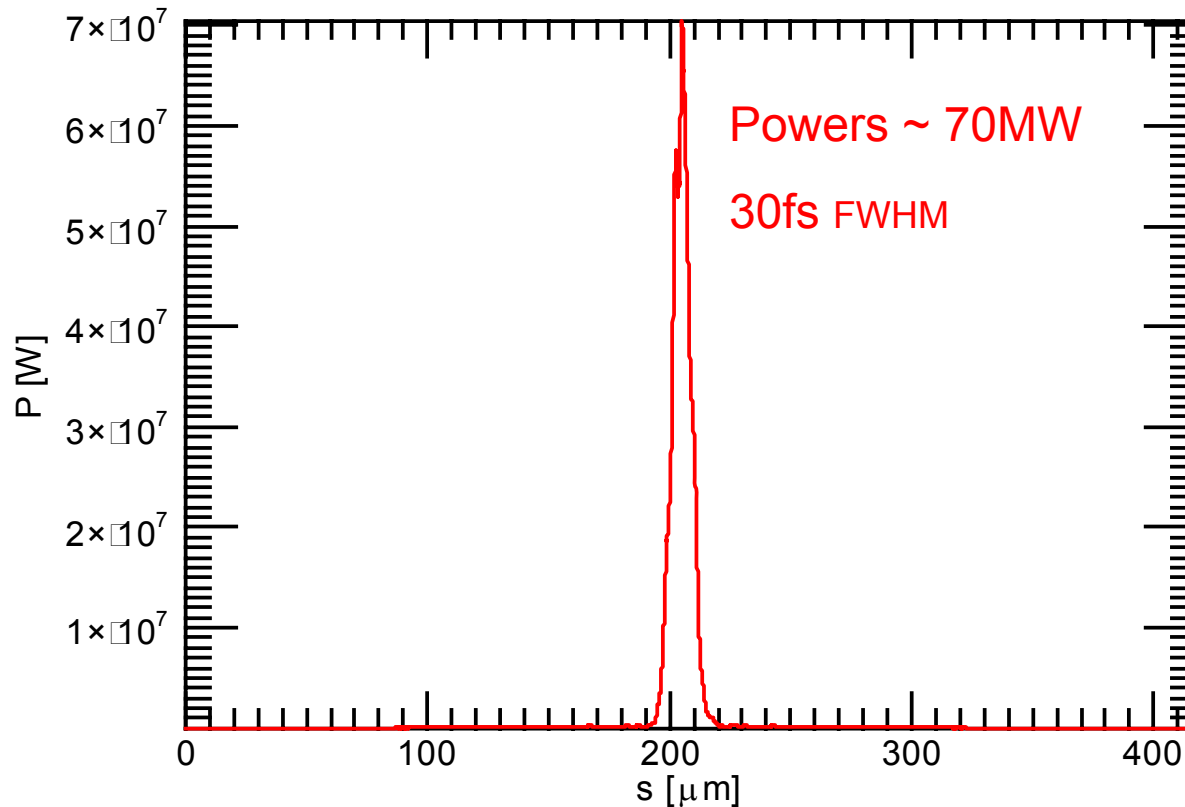
Planar undulator start PU1



VU5 VU4 VU3 VU2 VU1 PU8 PU7 PU6 PU5 PU4 PU3 PU2 PU1



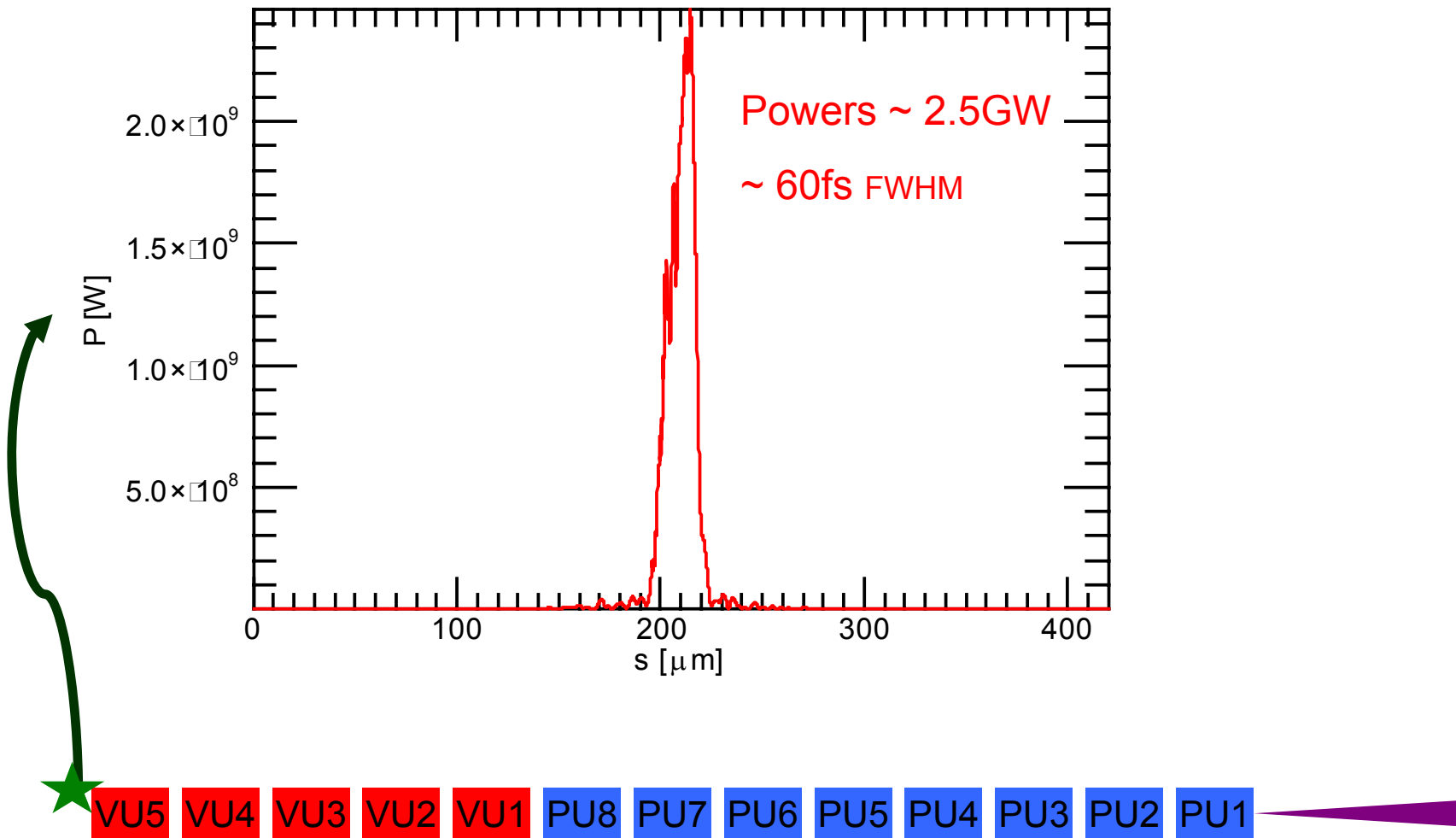
Planar undulator end PU8



VU5 VU4 VU3 VU2 VU1 PU8 PU7 PU6 PU5 PU4 PU3 PU2 PU1

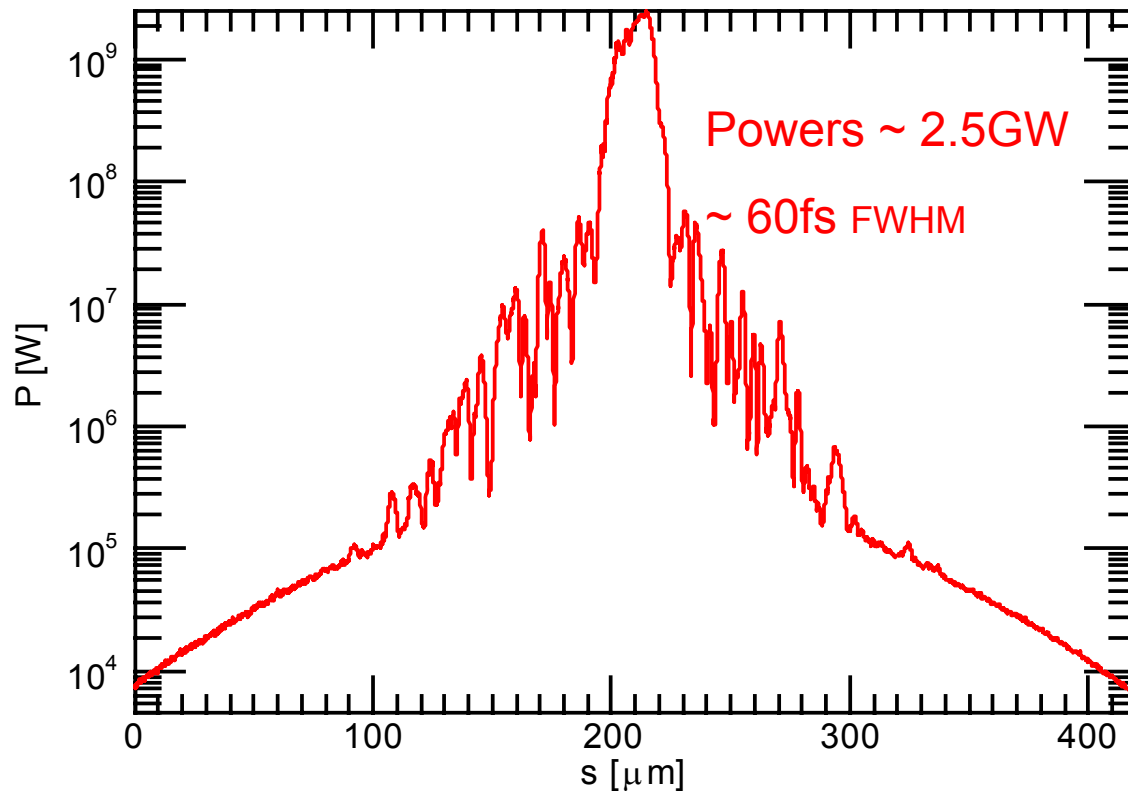


Helical undulator end VU5





Helical undulator end VU5

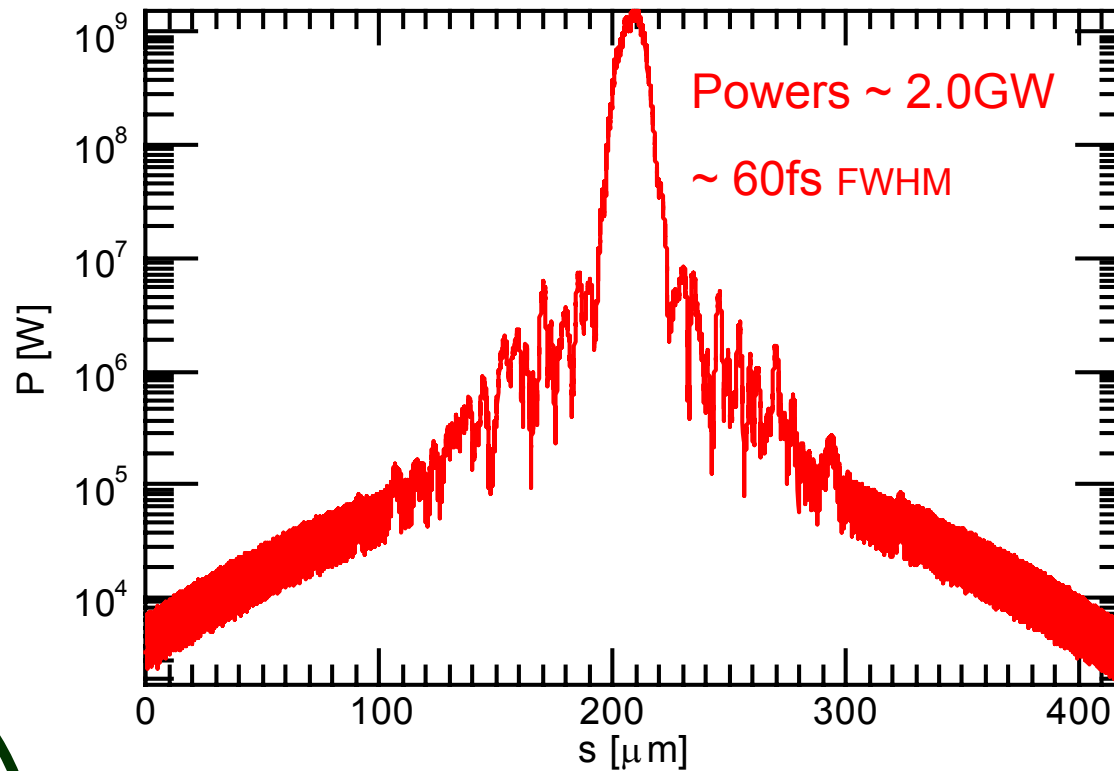


VU5 VU4 VU3 VU2 VU1 PU8 PU7 PU6 PU5 PU4 PU3 PU2 PU1



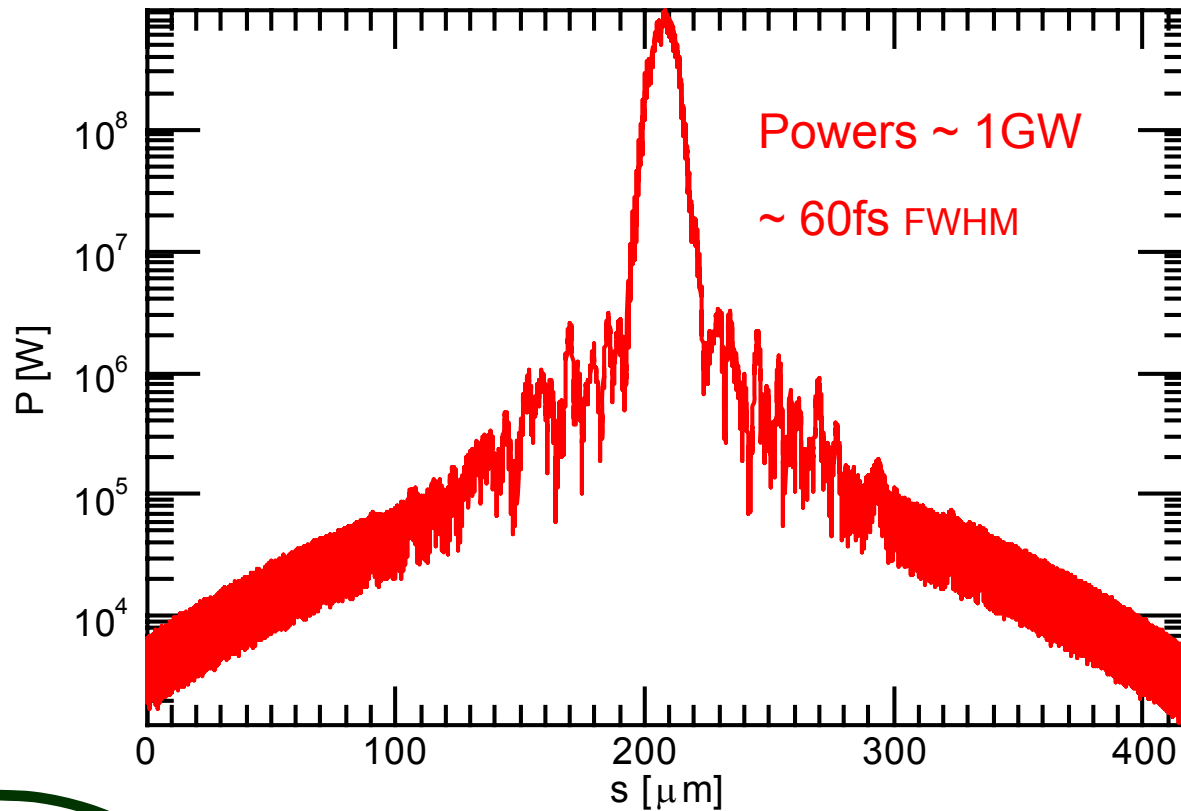


Helical undulator end VU4





Helical undulator end VU3

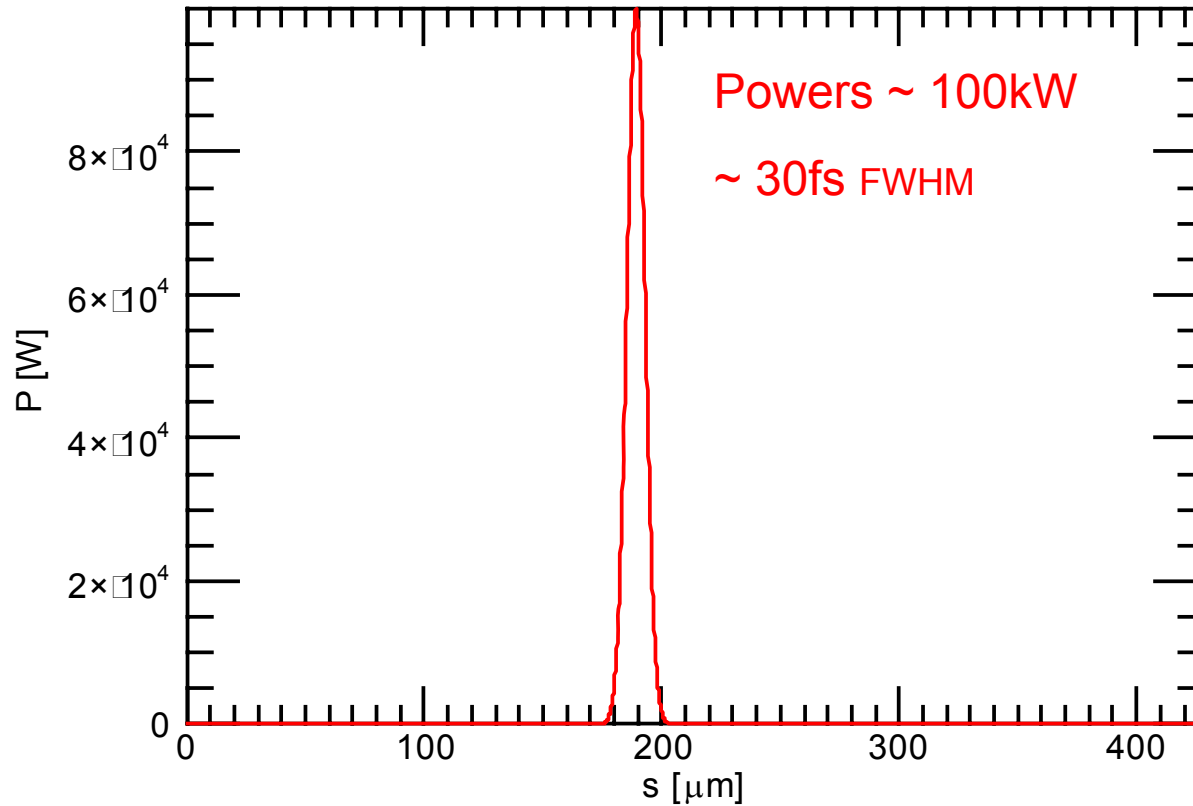




Simulations of XUV-FEL conceptual design XUV lasing - Pulses @ 10eV



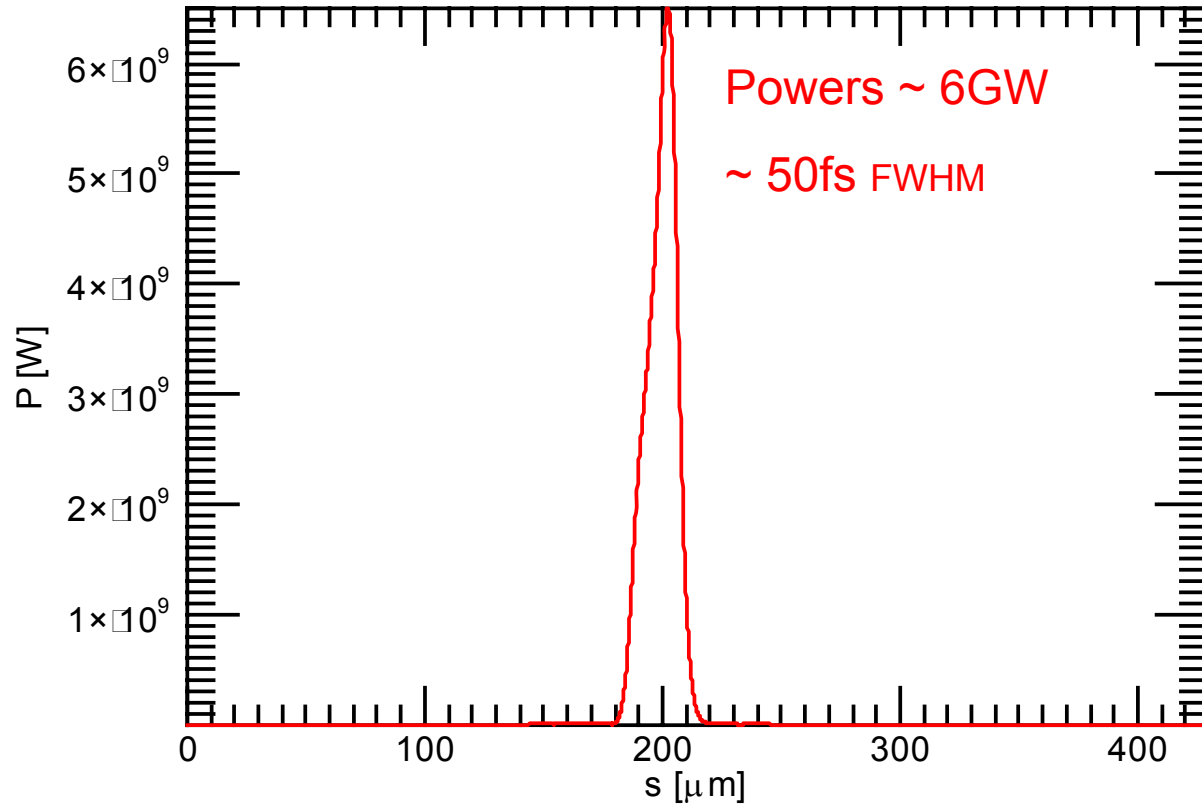
Helical undulator start VU1



VU5 VU4 VU3 VU2 VU1 ~~PU8 PU7 PU6 PU5 PU4 PU3 PU2 PU1~~



Helical undulator end VU5

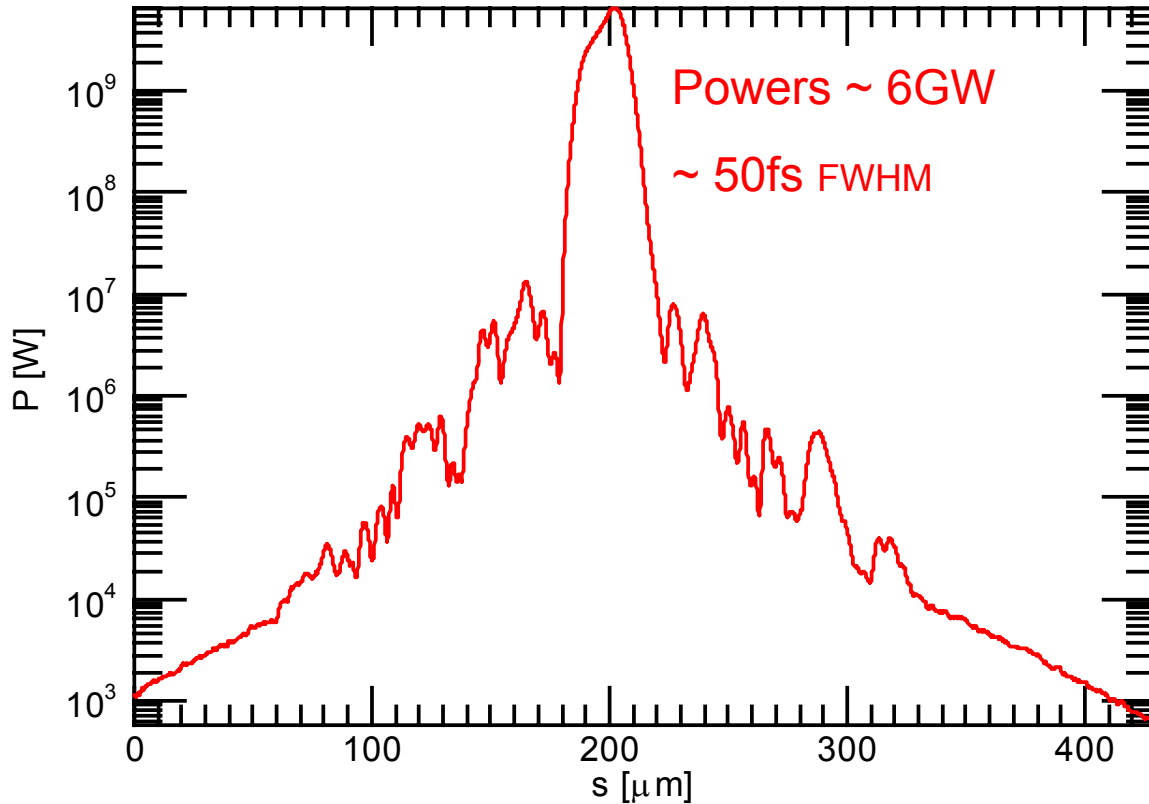


VU5 VU4 VU3 VU2 VU1 PU8 PU7 PU6 PU5 PU4 PU3 PU2 PU1



University of
Strathclyde

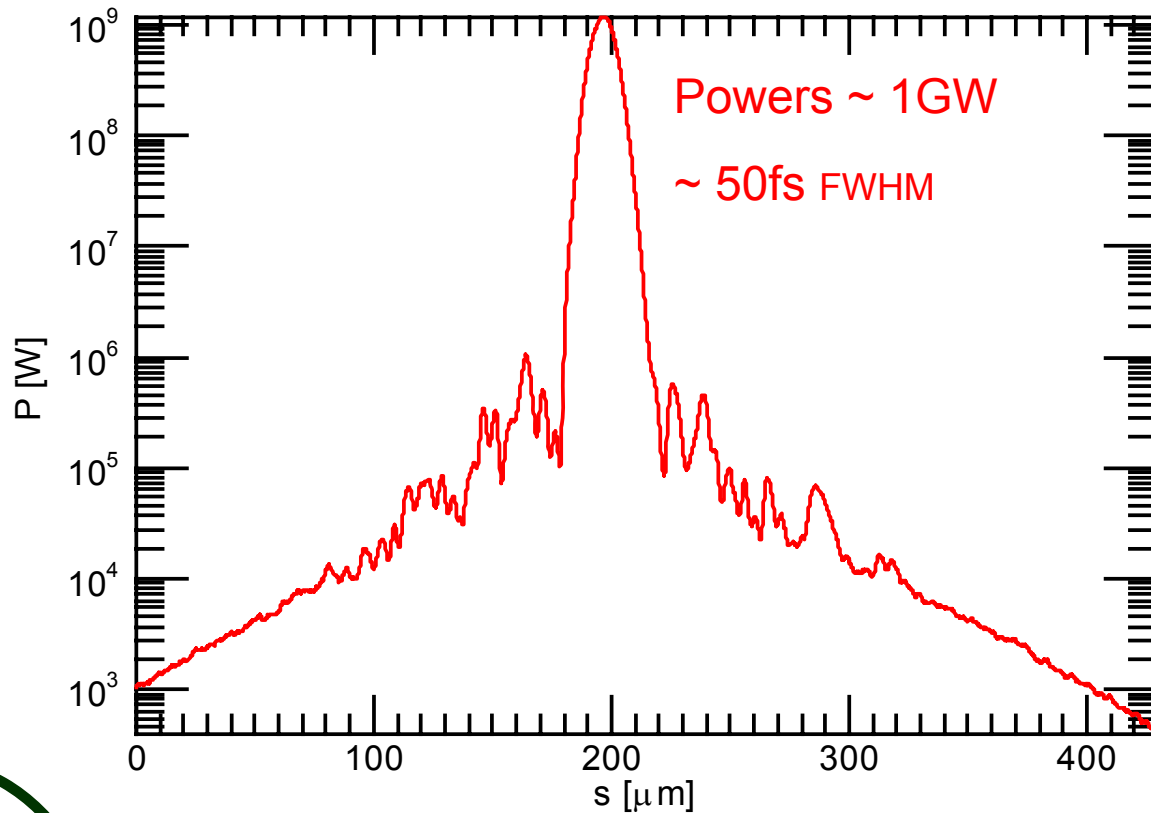
Helical undulator end VU5



VU5 VU4 VU3 VU2 VU1 PU8 PU7 PU6 PU5 PU4 PU3 PU2 PU1



Helical undulator end VU4





Further design work for XUV-FEL

- “Further design” – Technical Design Report ~end 2007
- Keep up-to-date with HHG development – rapidly changing
- Synchronism issues are critical*: keep timing offsets < 100 fs
- Need to simulate with more realistic electron pulses – s2e
- Re-assess and optimise design

*D. Dunning et al., *First Tolerance Studies for the 4GLS FEL Sources*, TUPPH057