

Recent progress in SR interferometer

-for small beam size measurement-

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Agenda

- 1. Brief introduction of beam size measurement through SR interferometry.**
- 2. Theoretical resolution of interferometry**
- 3. Reflective interferometer for measurement of beam size down to $5\mu\text{m}$ range.**
- 4. Imbalanced input method for measurement of very small beam size less than $5\mu\text{m}$**

- 1. A brief introduction to beam size measurement through SR interferometry**

To measure a size of object by means of spatial coherence of light (interferometry) was first proposed by H. Fizeau in 1868!

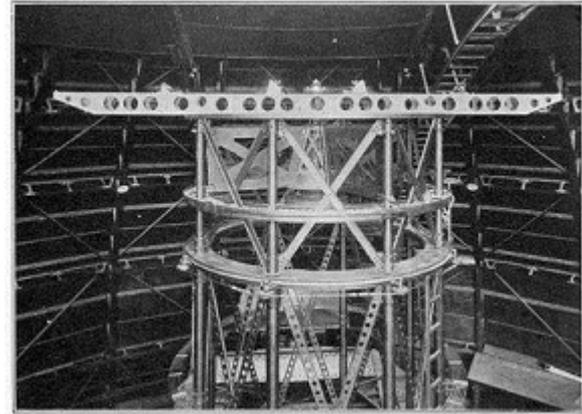
This method was realized by A.A. Michelson as the measurement of apparent diameter of star with his stellar interferometer in 1921.

This principle was now known as “ Van Cittert-Zernike theorem” because of their works;

1934 Van Cittert

1938 Zernike.

Michelson's stellar interferometer Wilson mountain observatory



Spatial coherence and profile of the object

Van Cittert-Zernike theorem

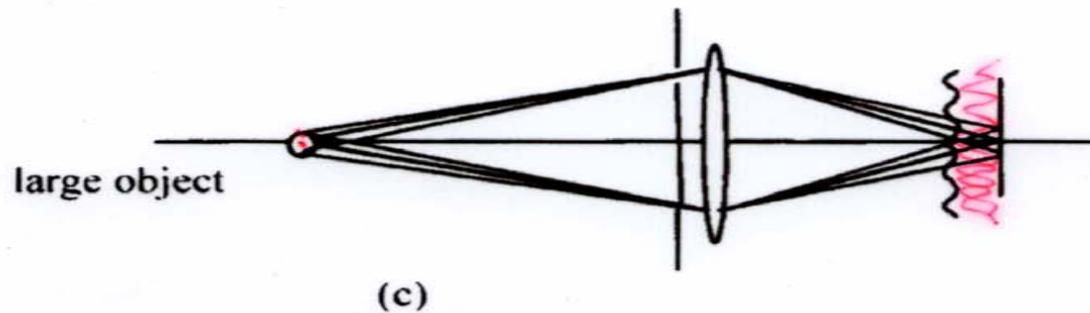
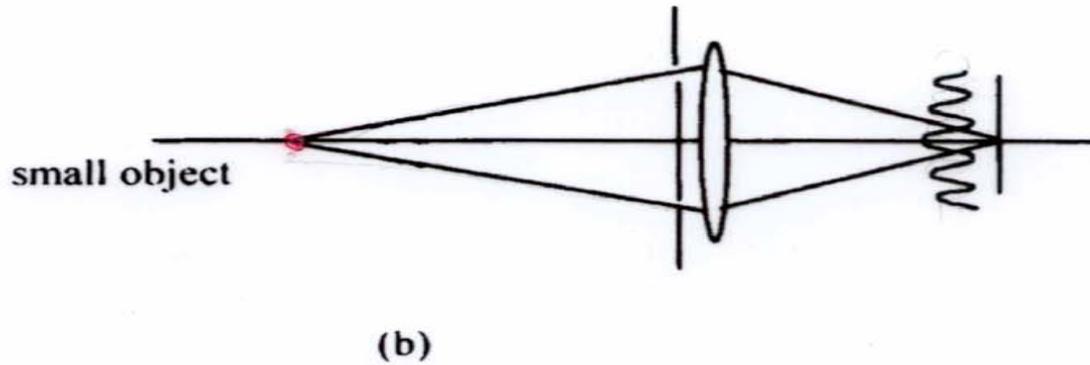
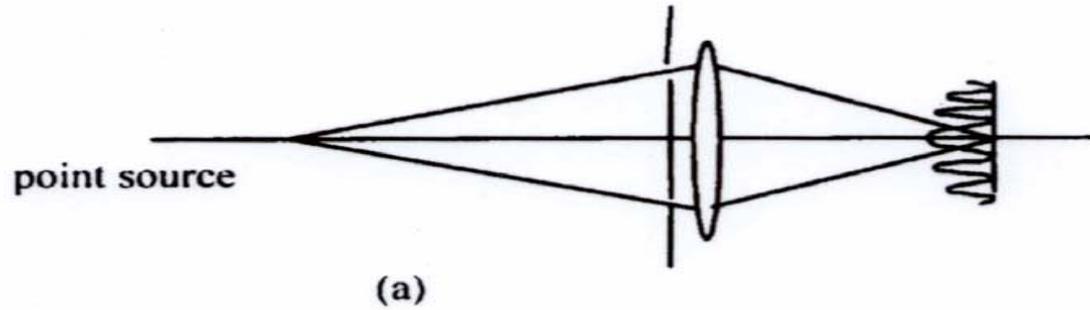
According to van Cittert-Zernike theorem, **with the condition of light is 1st order temporal incoherent (no phase correlation)**, the complex degree of spatial coherence $\gamma(u_x, u_y)$ is given by **the Fourier Transform** of the spatial profile $f(x, y)$ of the object (beam) at longer wavelengths such as visible light.

$$\gamma(u_x, u_y) = \iint f(x, y) \exp \left\{ -i \cdot 2 \cdot \pi (u_x \cdot x + u_y \cdot y) \right\} dx dy$$

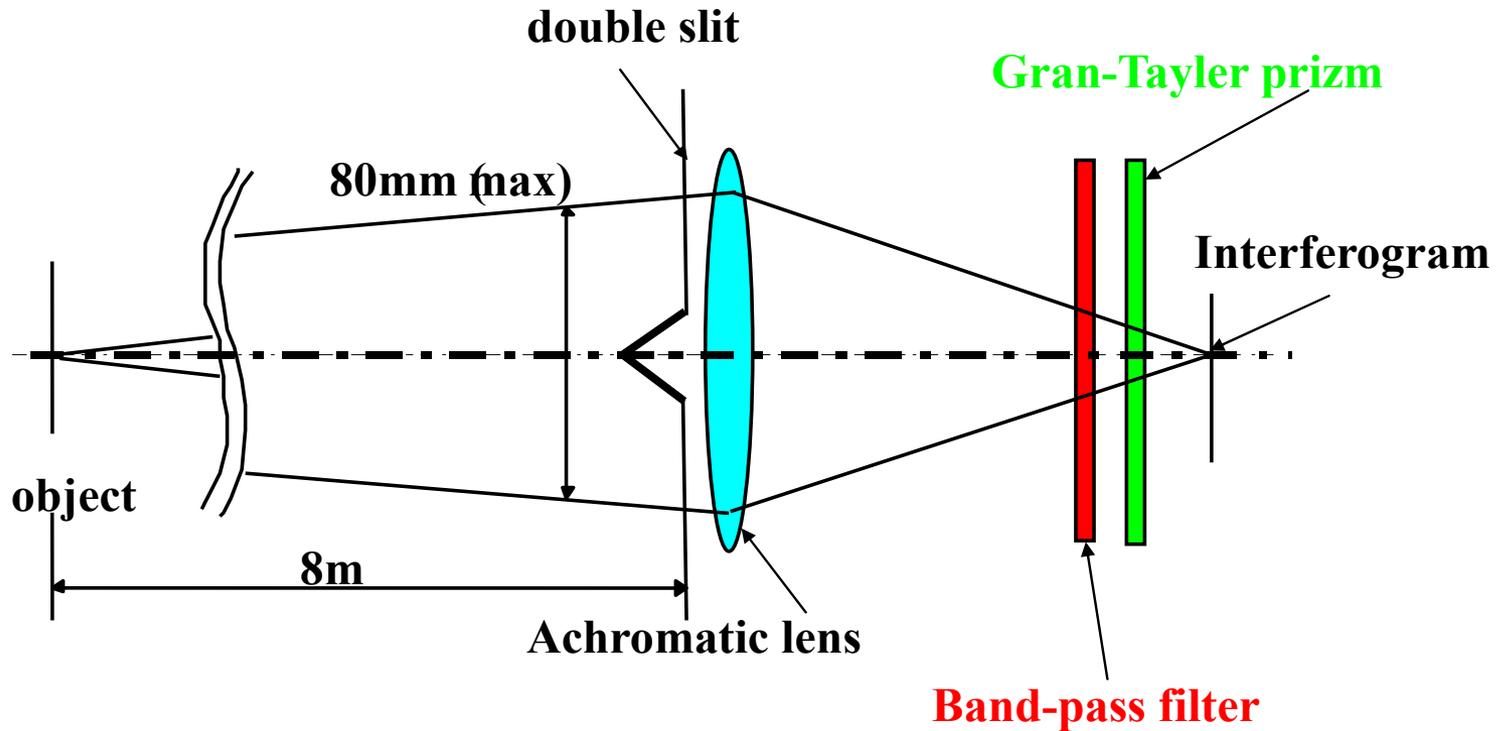
where u_x, u_y are spatial frequencies given by;

$$u_x = \frac{D_x}{\lambda \cdot R_0}, \quad u_y = \frac{D_y}{\lambda \cdot R_0}$$

Simple understanding of van Cittert-Zernike theorem



Typical arrangement for refractive interferometer

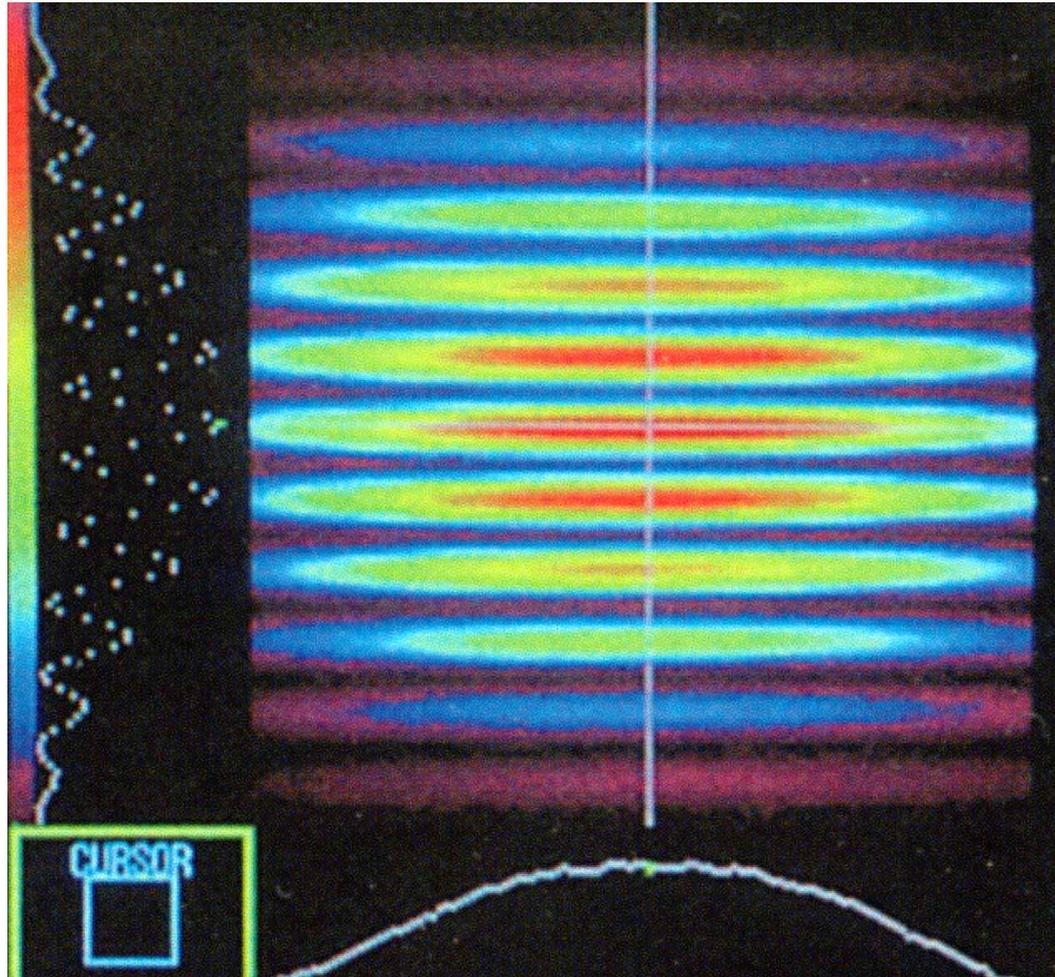


$$I(y, D) = \int (I_1 + I_2) \cdot \left\{ \sin c \left(\frac{\pi \cdot a \cdot y \cdot \chi(D)}{\lambda \cdot f} \right) \right\}^2 \cdot \left\{ 1 + \gamma \cdot \cos \left(k \cdot D \cdot \left(\frac{y}{f} + \psi \right) \right) \right\} d\lambda$$

$$\gamma = \left(\frac{2 \cdot \sqrt{I_1 \cdot I_2}}{I_1 + I_2} \right) \cdot \left(\frac{I_{\max} - I_{\min}}{I_{\max} + I_{\min}} \right), \quad \psi = \tan^{-1} \frac{S(D)}{C(D)}$$

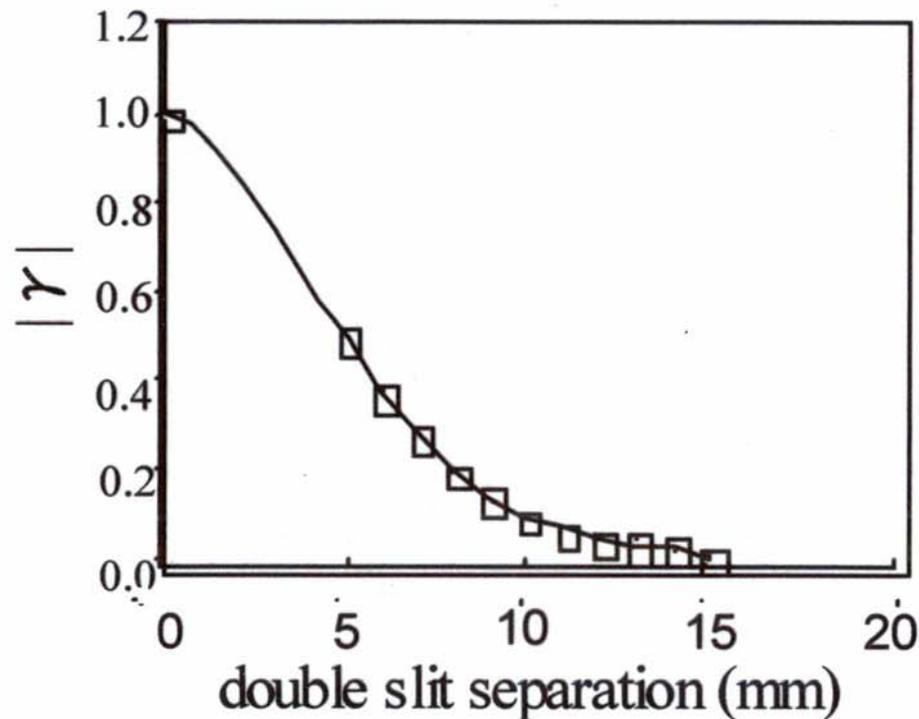
Typical interferogram in vertical direction at the Photon Factory (1994).

$D=10\text{mm}$



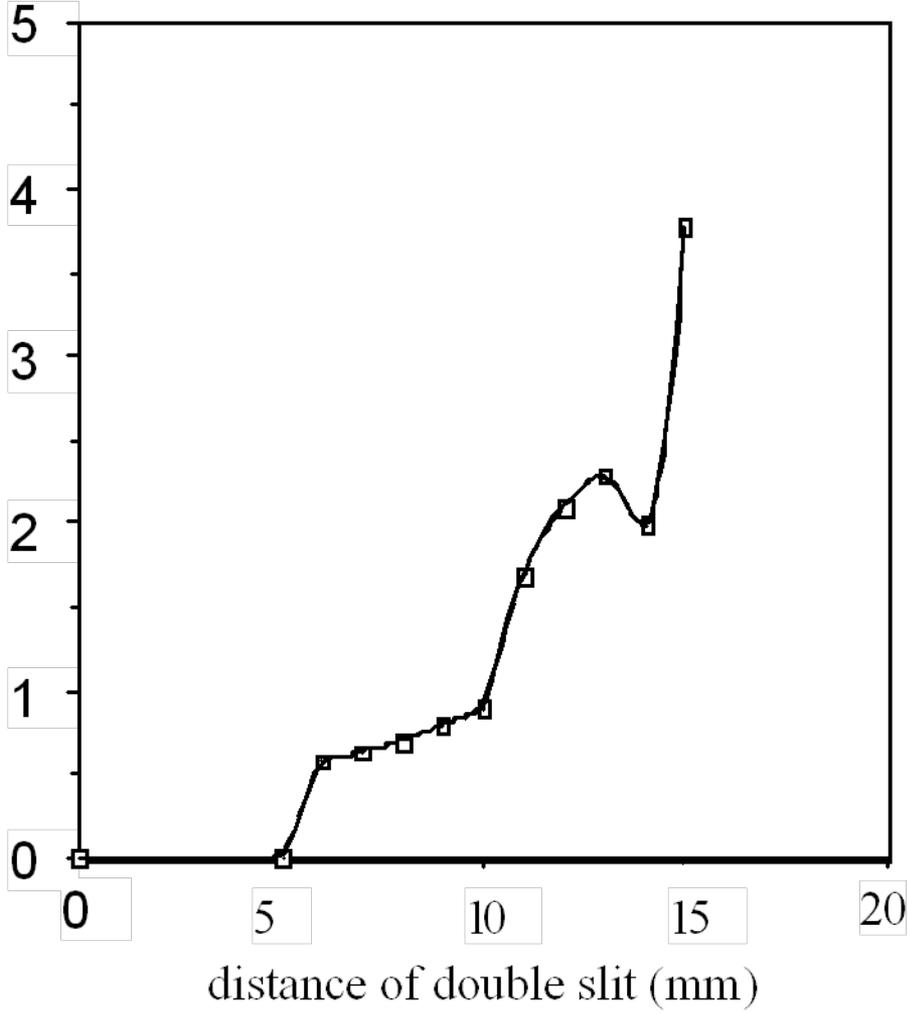
Result of spatial coherence measurement

Absolute value of the complex degree of spatial coherence measured at the Photon Factory KEK. (1994)

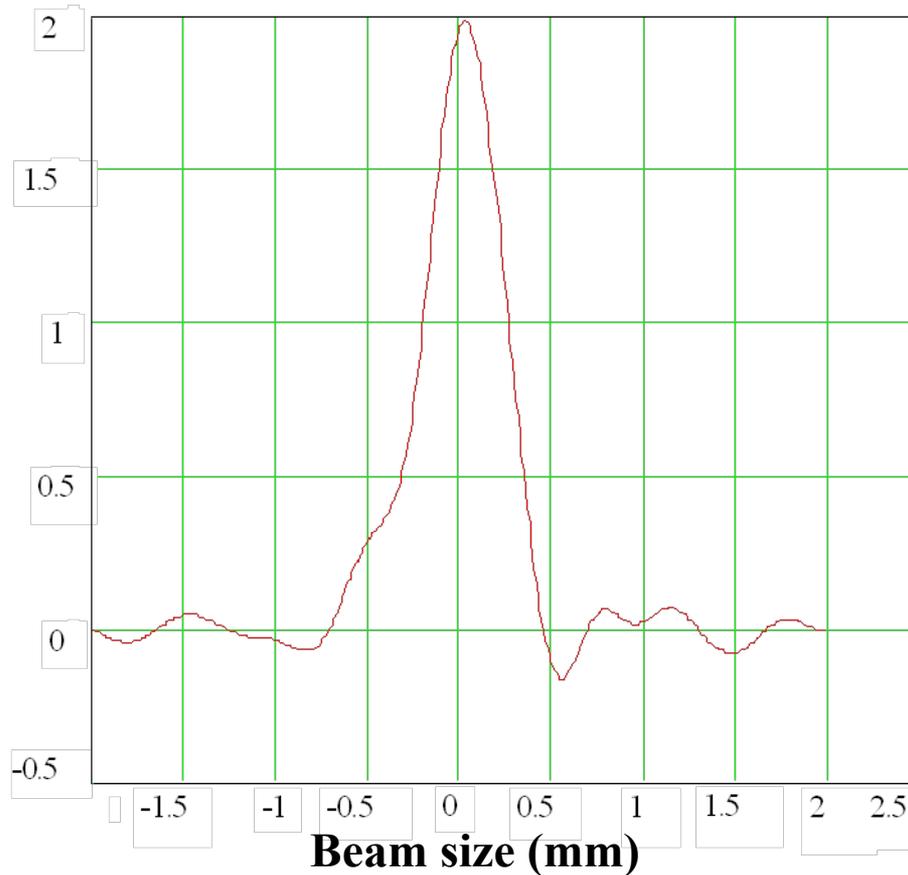


Phase of the complex degree of spatial coherence

vertical axis is phase in radian



Reconstruction of beam profile by Fourier transform

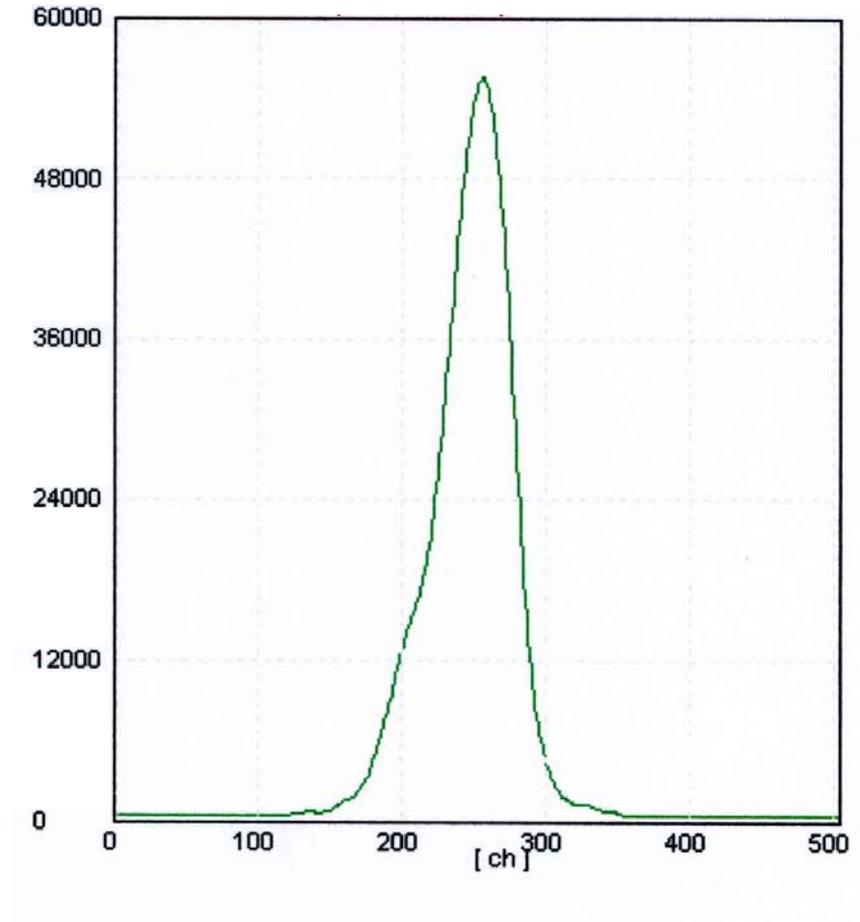


Vertical beam profile obtained by a Fourier transform of the complex degree of coherence.

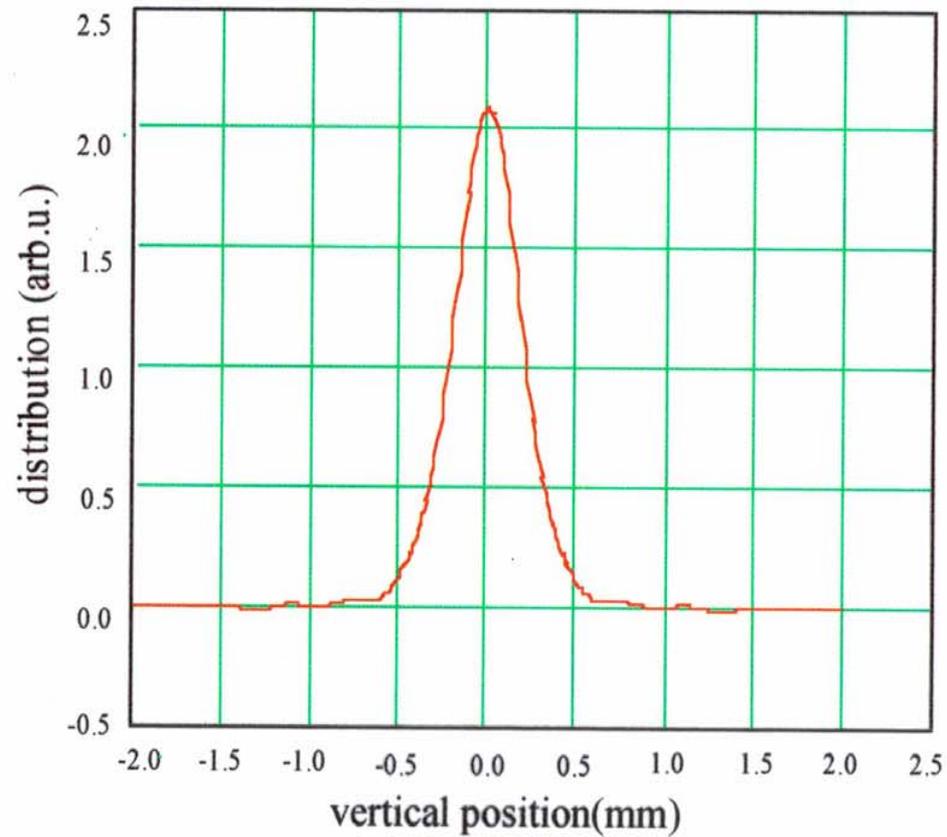
Comparison between image



Beam profile taken with
an imaging system



Vertical beam profile obtained by Fourier Cosine transform



SMALL BEAM SIZE MEASUREMENT

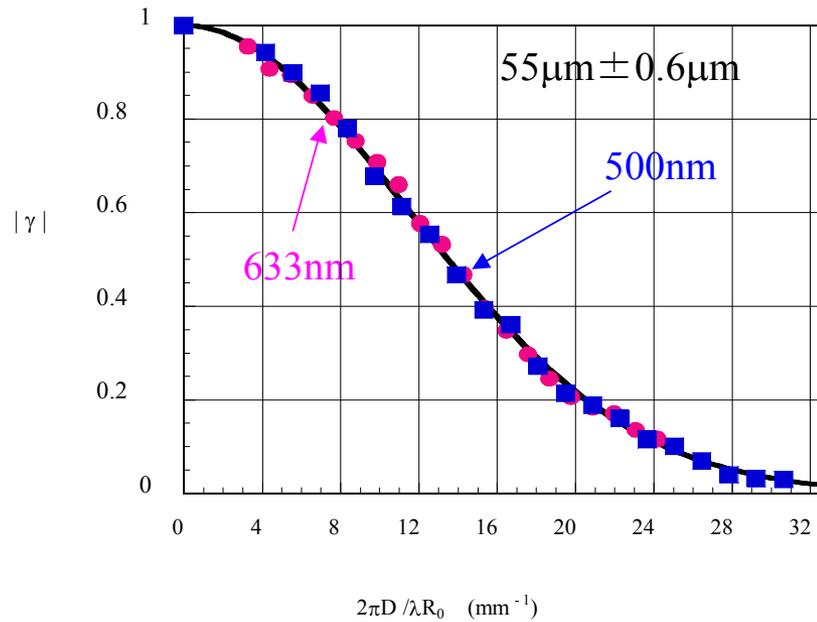
We often approximate the beam profile with a Gaussian shape. A spatial coherence is also given by a Gauss function. We can evaluate a RMS width of spatial coherence by using q least-squares analysis. The RMS beam size σ_{beam} is given by the RMS width of the spatial coherence curve σ_γ as follows:

$$\sigma_{beam} = \frac{\lambda \cdot R}{2 \cdot \pi \cdot \sigma_\gamma}$$

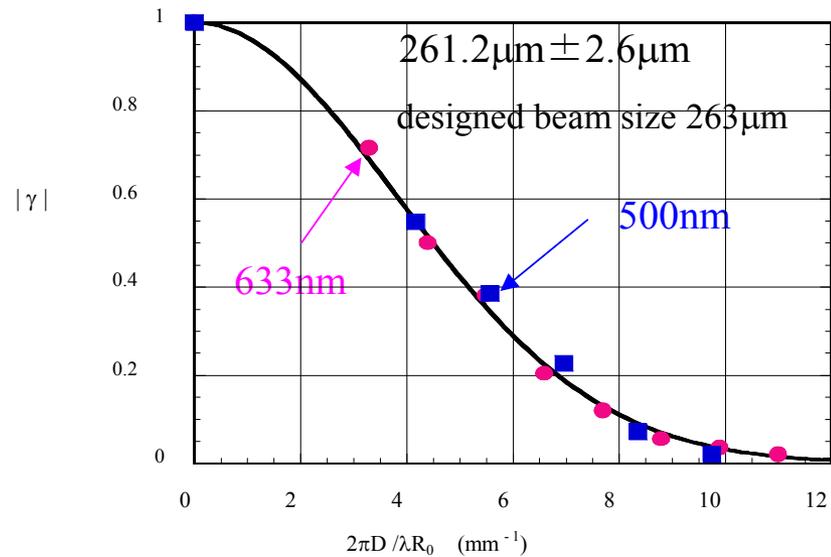
where R : distance between the beam and the double slit.

λ : wave length

Vertical and horizontal beam size at the Photon Factory



(a) vertical



(b) horizontal

We can also evaluate the RMS. beam size from one data of visibility, which is measured at a fixed separation of double slit. The RMS beam size σ_{beam} is given by ,

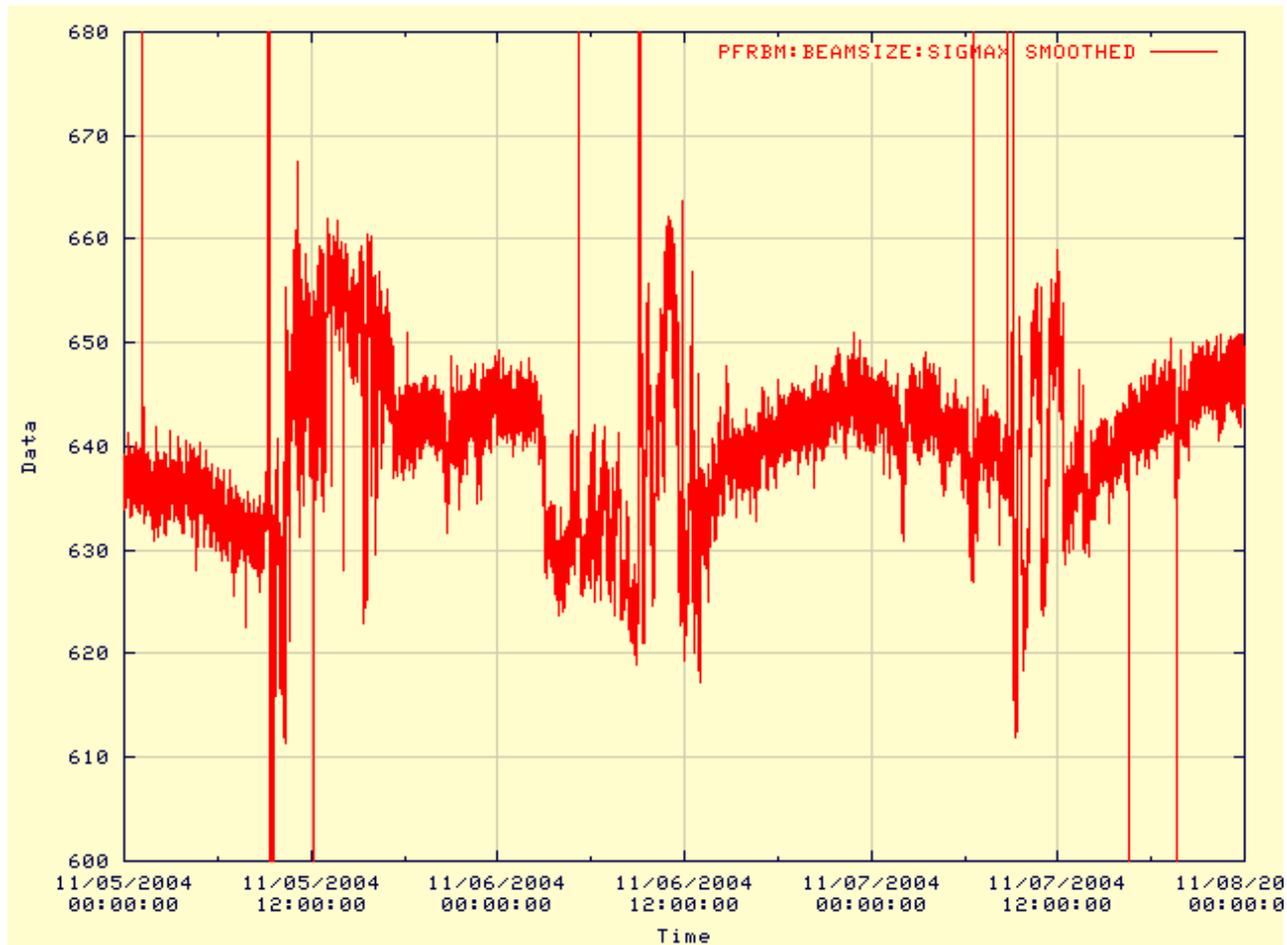
$$\sigma_{\text{beam}} = \frac{\lambda \cdot R_0}{\pi \cdot D} \cdot \sqrt{\frac{1}{2} \cdot \ln\left(\frac{1}{\gamma}\right)}$$

where γ denotes the visibility, which is measured at a double slit separation of D.

To consider that in the case to make an image, the resolution is limited by diffraction which is a Fourier transform using a given region of spatial frequency space (measurement in the real space).

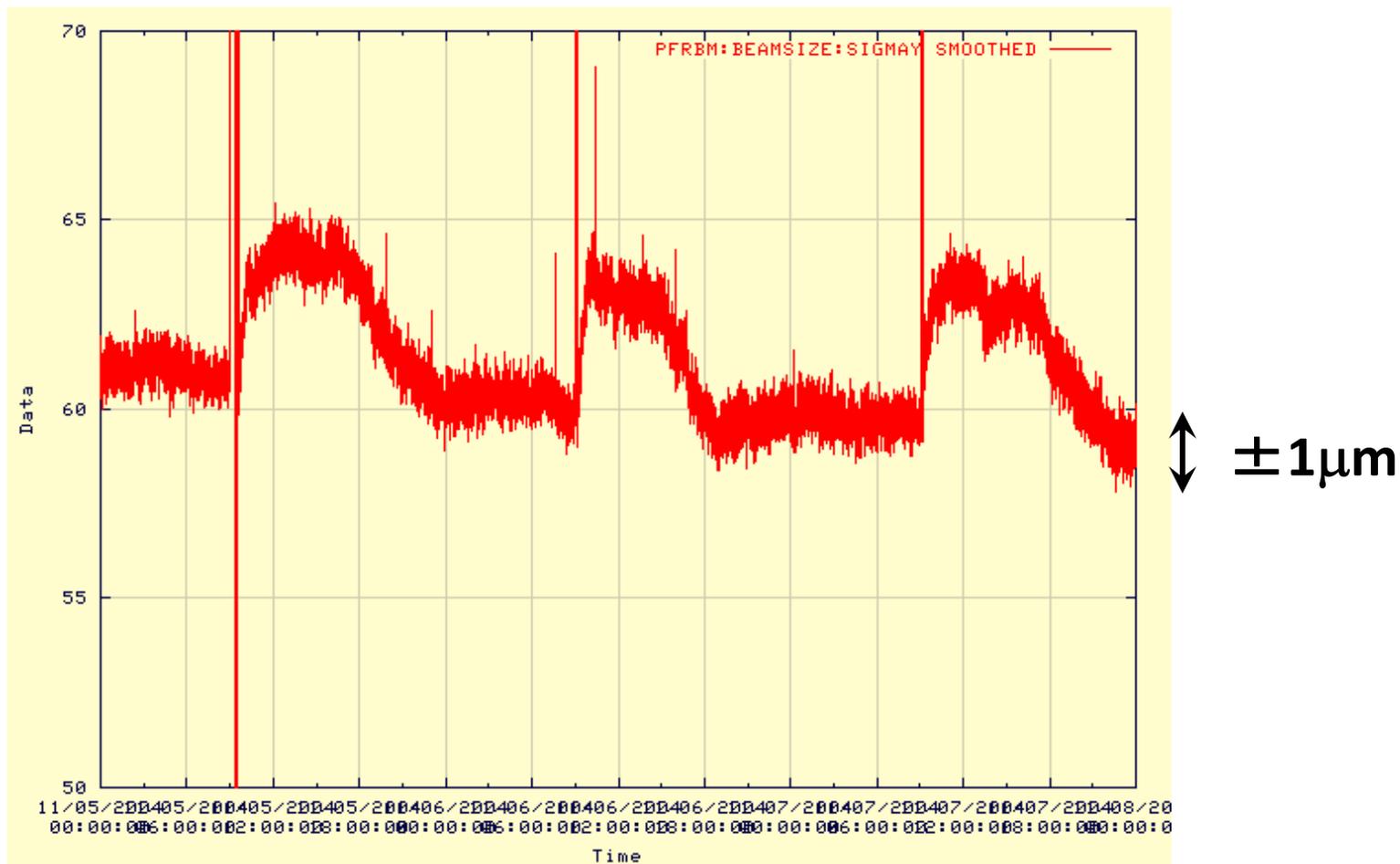
In the case of interferometry, we can measure a small beam size with limited region of spatial frequency space by means of these two methods (measurement in the inverse space).

Horizontal beam size measurement



↕ ±3μm

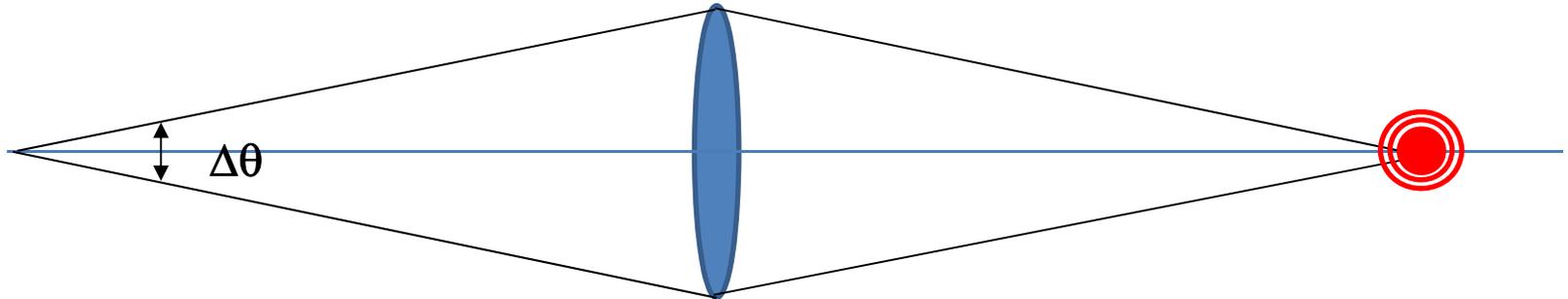
Vertical beam size measurement



2.Theoretical resolution of interferometry

**Uncertainty principle
in phase of light**

Uncertainty principal in imaging.

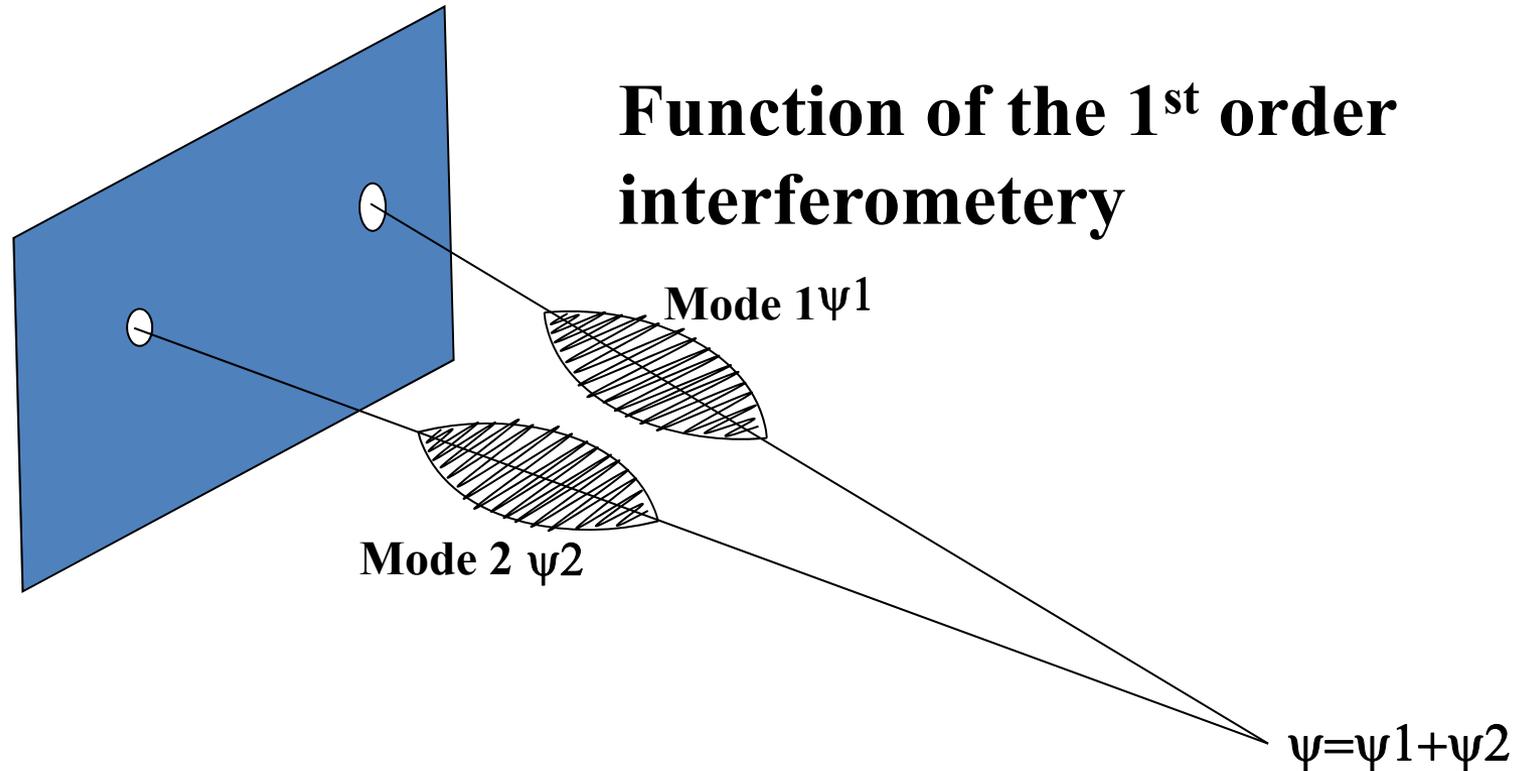


$$\Delta\theta/\lambda \cdot \Delta x \geq 1,$$

**So, large opening of light will
necessary to obtain a good spatial
resolution.**

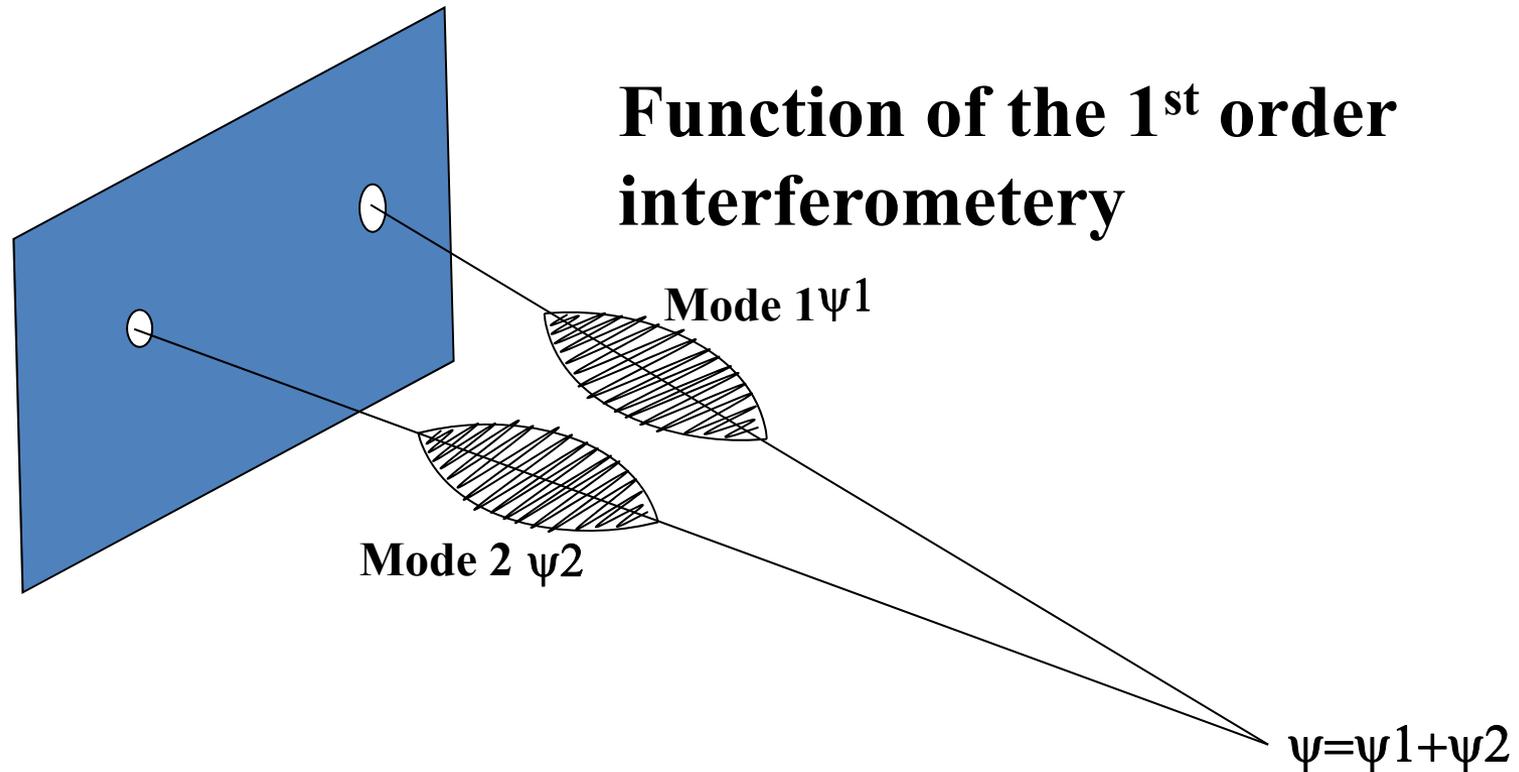
Uncertainty principal in interferometry ?

Uncertainty principal in interferometry



Measure the correlation of light phase in two modes

Uncertainty principal in interferometry



Measure the correlation of light phase in two modes



Uncertainty in Phase $\Delta\phi$

The interference fringe will be smeared by the uncertainty of phase.

$$I(y, D) = \int_{\Delta\phi} (I_1 + I_2) \cdot \left\{ \text{sinc} \left(\frac{\pi \cdot a \cdot y}{\lambda \cdot f} \right) \right\}^2 \left\{ 1 + \cos \left(k \cdot D \frac{y}{f} + \phi \right) \right\} d\phi$$

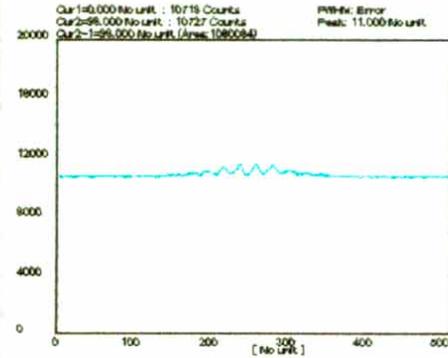
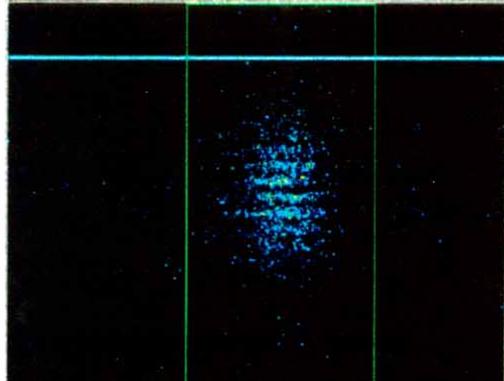
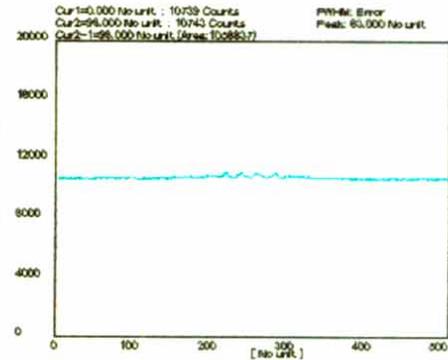
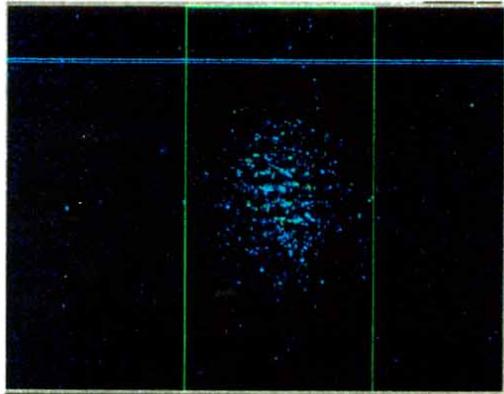
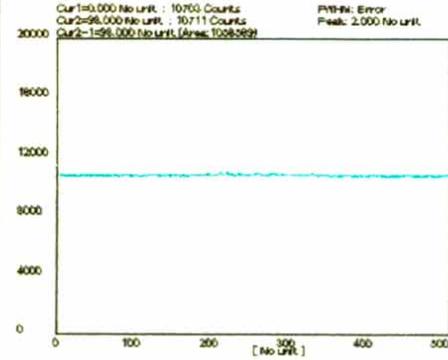
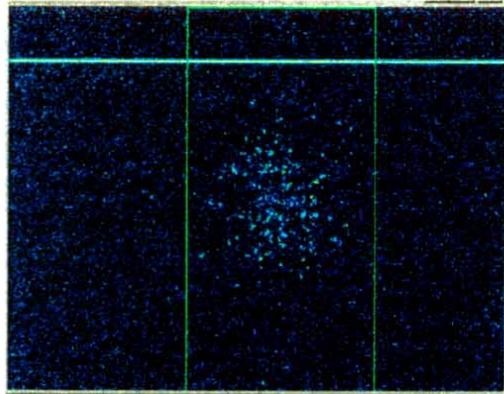
According to quantum optics,

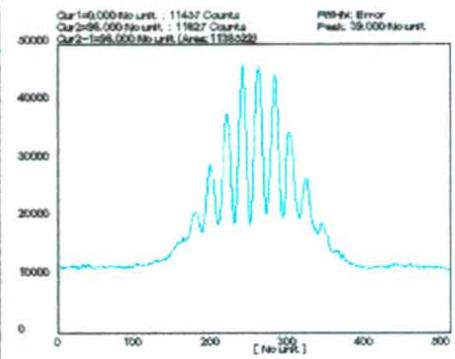
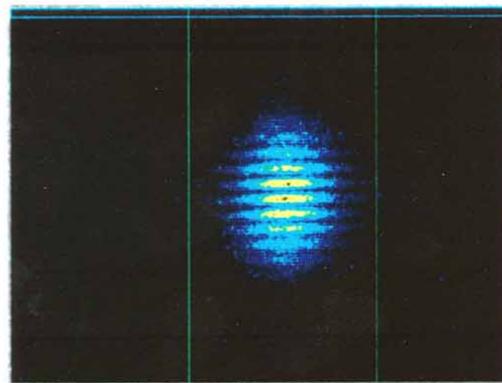
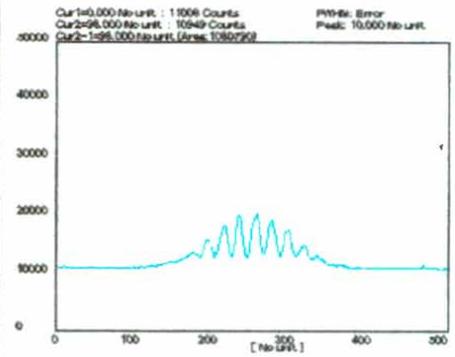
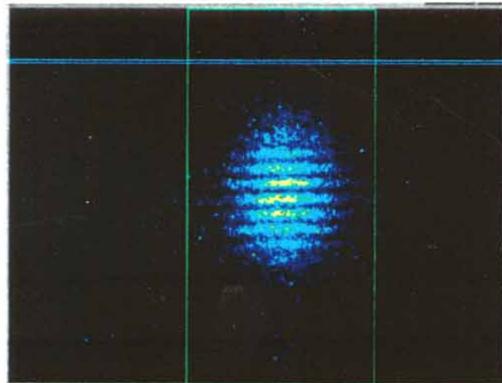
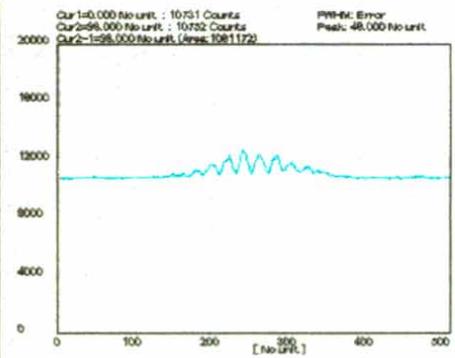
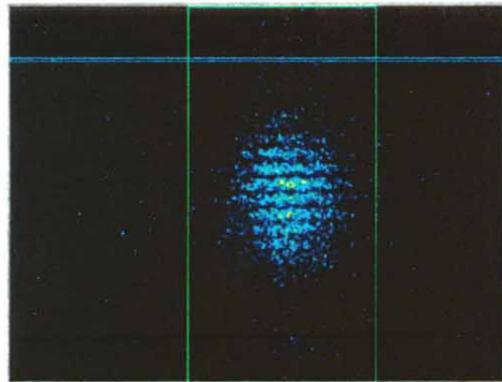
Uncertainty principle concerning to phase is given by

$$\Delta\phi \cdot \Delta N \geq 1/2$$

where ΔN is uncertainty of photon number.

We cannot observe interference fringe with small number of photons!





Actually, different from imaging, we can use large number of photons (intensity), so uncertainty in phase is very small (this is the reason light seems wave)

**A comparison between imaging,
we can use large number of photons
(intensity), so uncertainty in phase is
very small (this is the reason light
seems wave)**

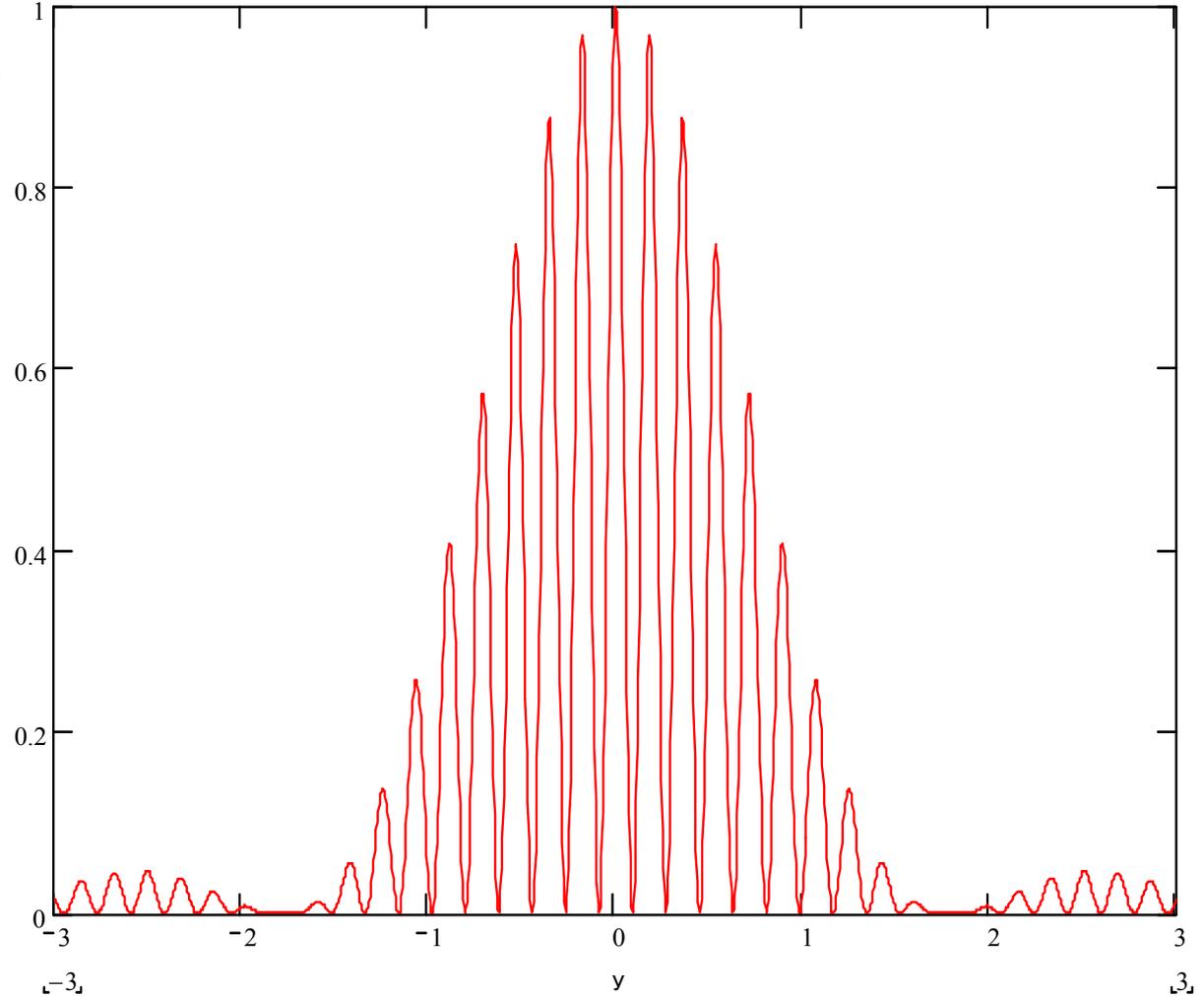


**As a result, theoretical resolution is
very high, and practically resolution
will be limited by measurement error
such as baseline noise in detector.**

Small size of the beam will give a good visibility

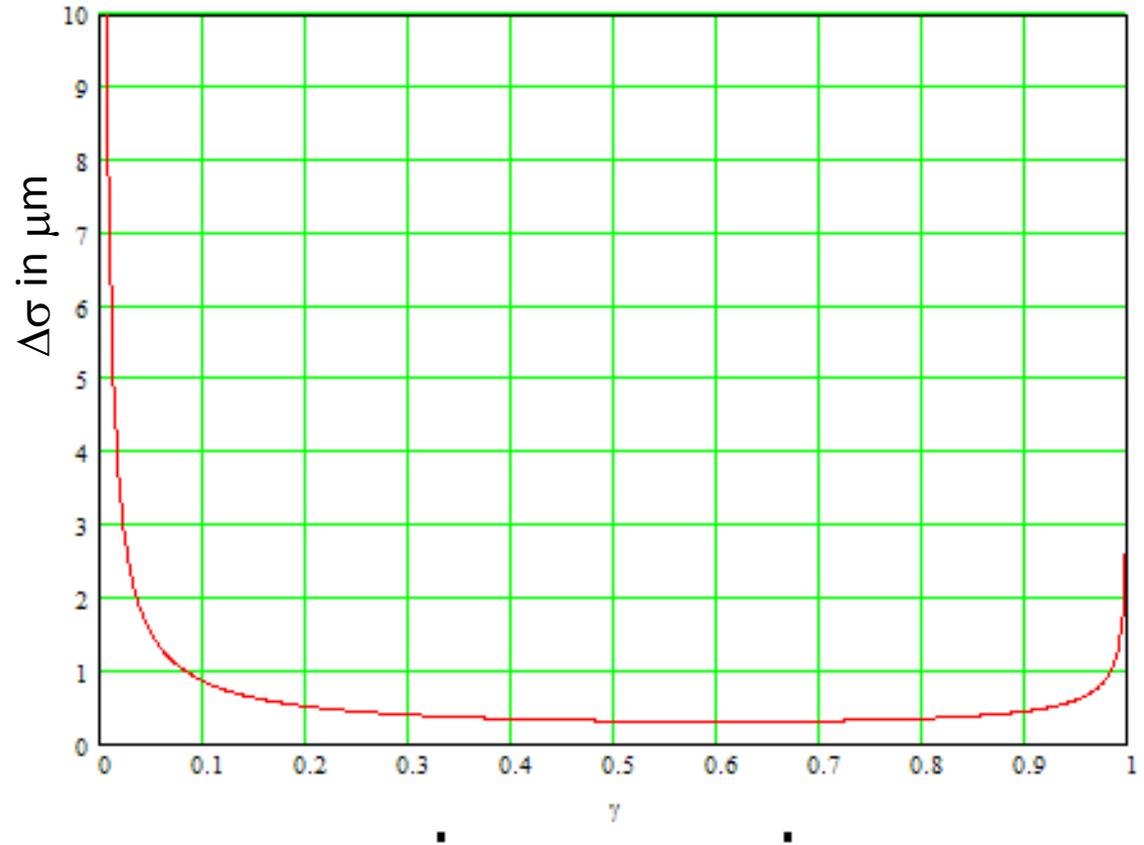


Strongly influenced by baseline noise!



Error transfer
from $\Delta\gamma$ to $\Delta\sigma$
with constant $\Delta\gamma$

$$\Delta\sigma \propto \frac{1}{\gamma \cdot \ln\left(\frac{1}{\gamma}\right)} \Delta\gamma$$



**So, important point in small beam
size measurement**

is

**How to escape from noise in
visibility measurement**

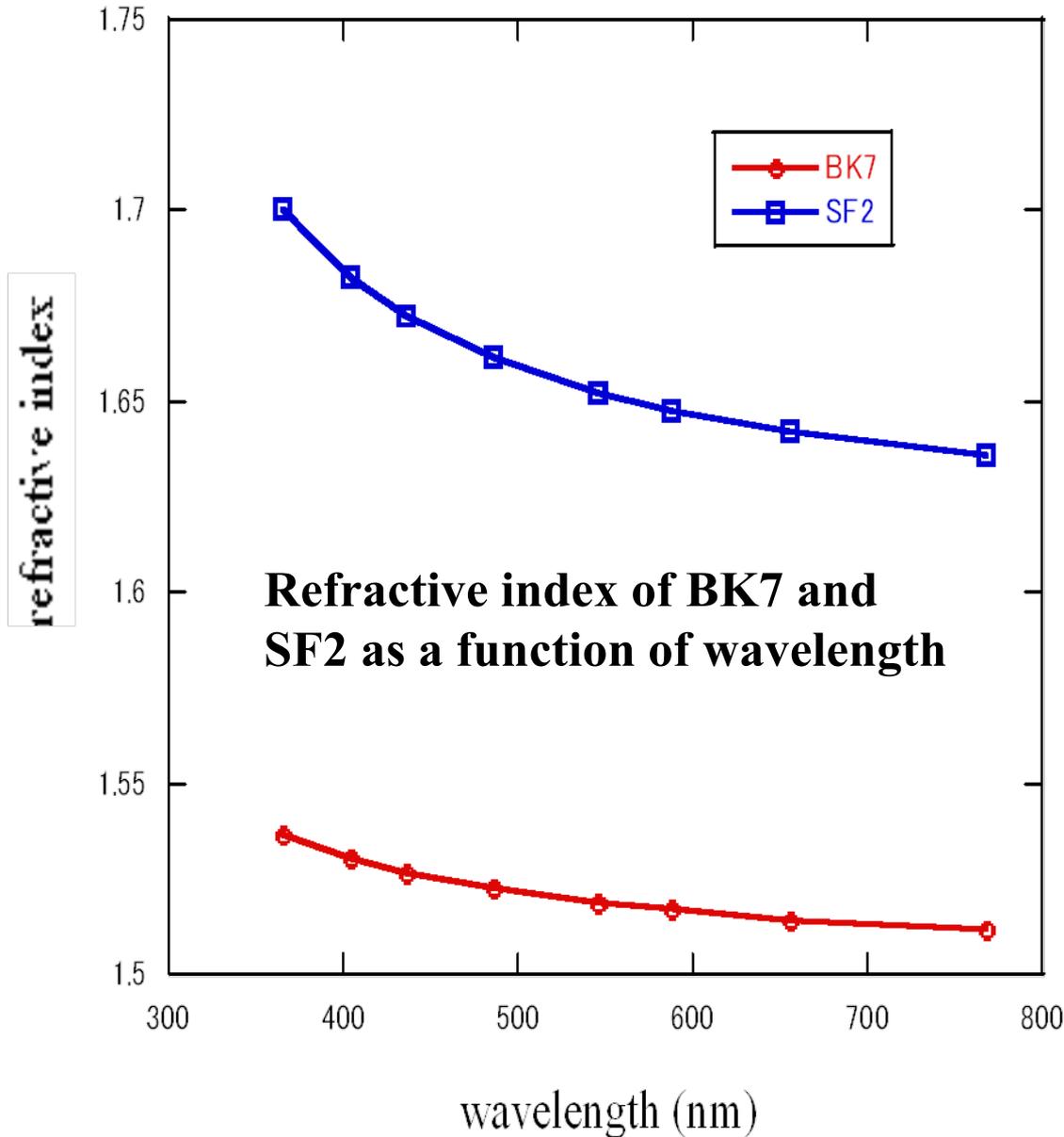
1. Use larger separation of double slit

2. Use shorter wavelength

Both of this will reduce visibility of interferogram

- 1. Use larger separation of double slit
limited by opening angle of SR**
- 2. Use shorter wavelength
mainly limited by chromatic
aberrations in focusing optics.**

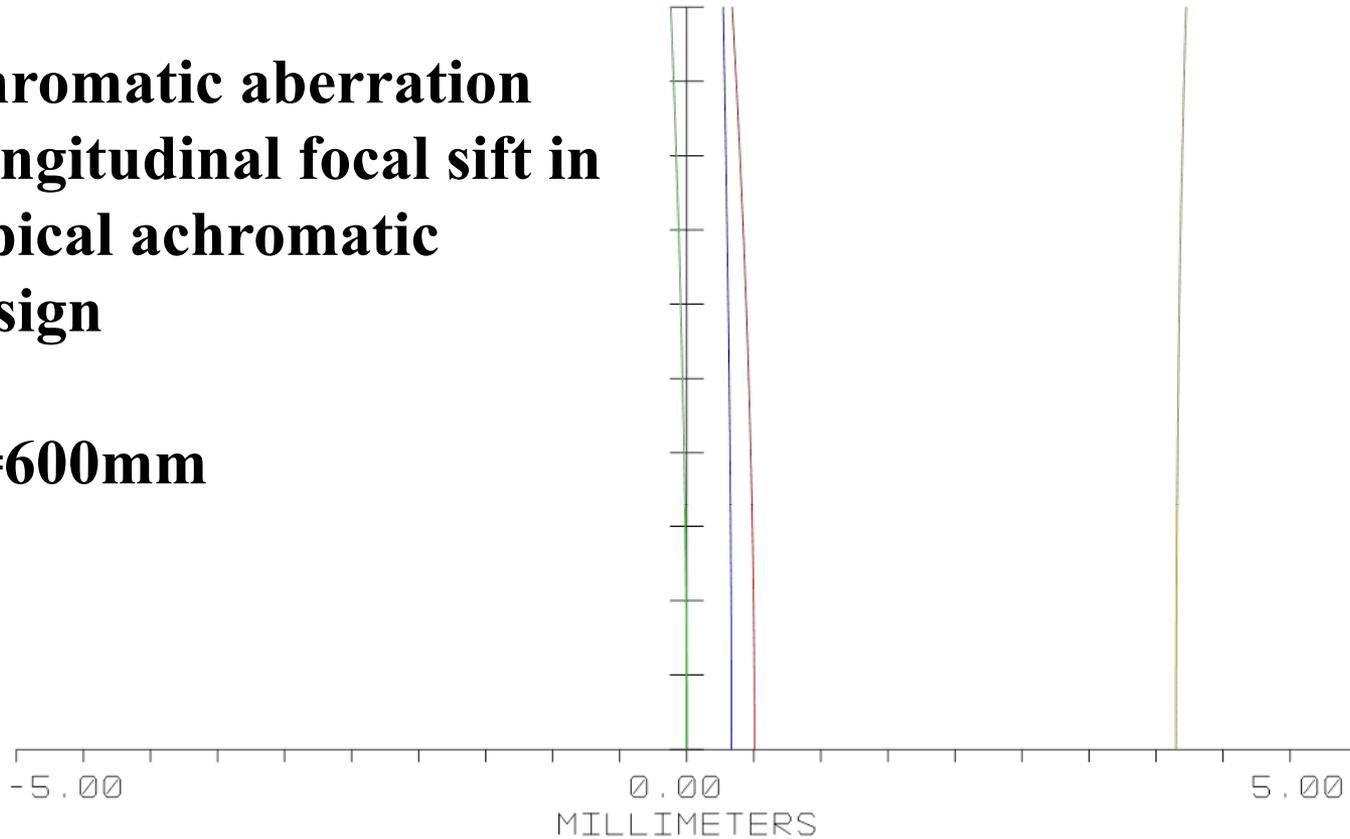
Elimination of chromatic aberration at 400nm is very difficult due to large partial dispersion ratio of glass



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Chromatic aberration (longitudinal focal shift in typical achromatic design)

F=600mm



LONGITUDINAL ABERRATION

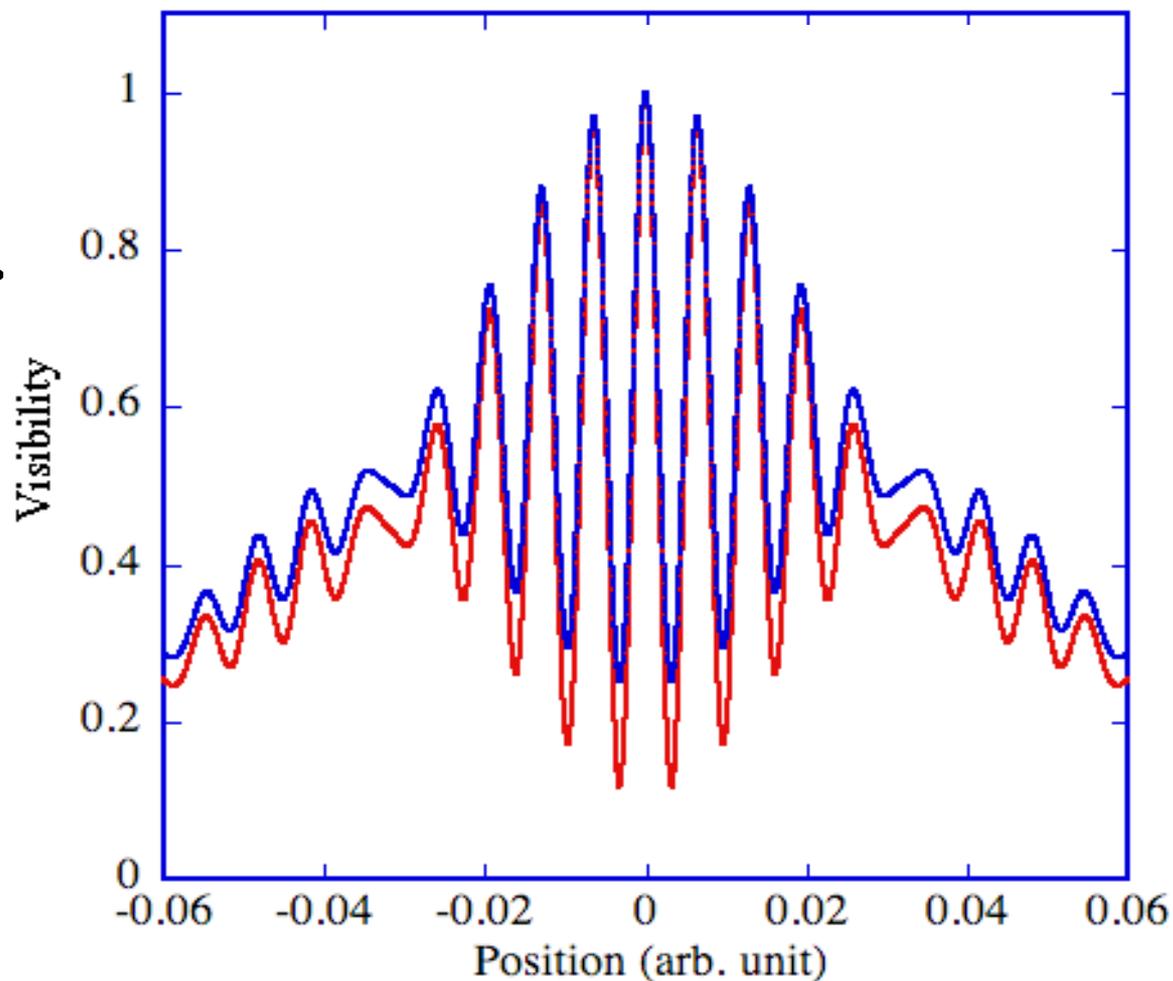
01LA0807 PRECISION OPTIMIZED ACHROMATS

TUE NOV 1 2011

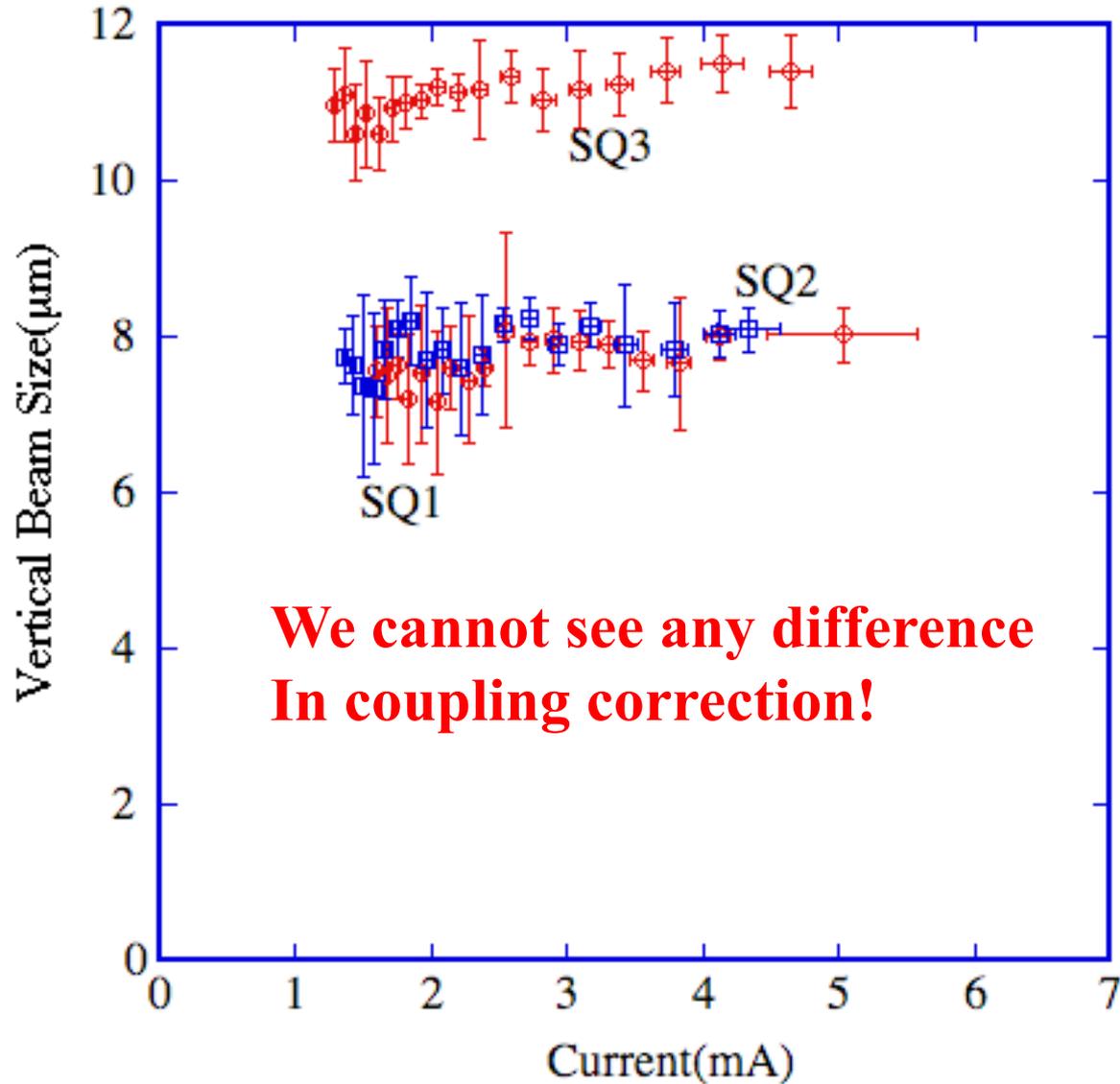
WAVELENGTHS: 0.480 0.546 0.644 0.400

TEMPSTOK.ZMX
CONFIGURATION 1 OF 1

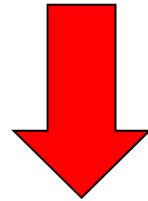
**Interferogram with
chromatic aberration
and without
chromatic aberration.
 $\lambda=400\text{nm}$, $\Delta\lambda=80\text{nm}$
Lens:achromat
 $D=45\text{mm}$ $f=600\text{mm}$**



Results by normal refractive interferometer using $\lambda=400\text{nm}$



If the chromatic aberration at 400nm is measure source of error in 5 μ m range beam size measurement,



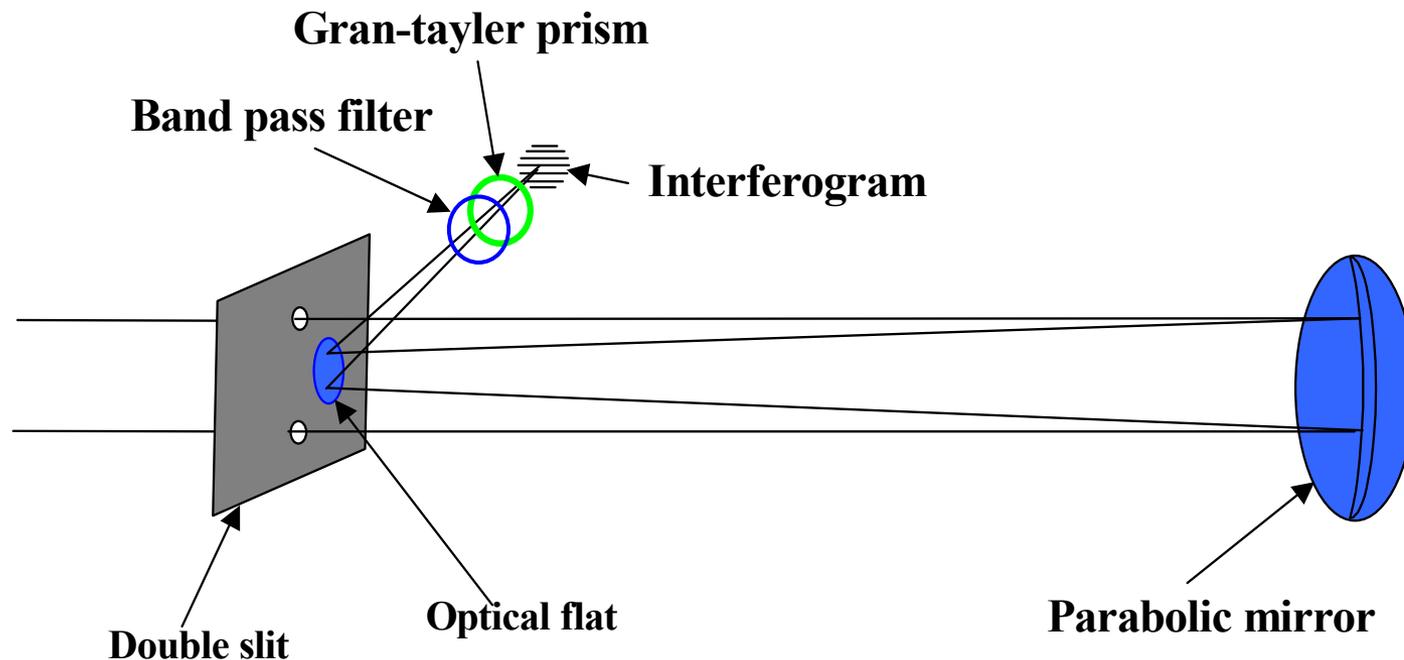
**Use reflective optics!
Reflective system has no chromatic aberration.**

3. Reflective interferometer

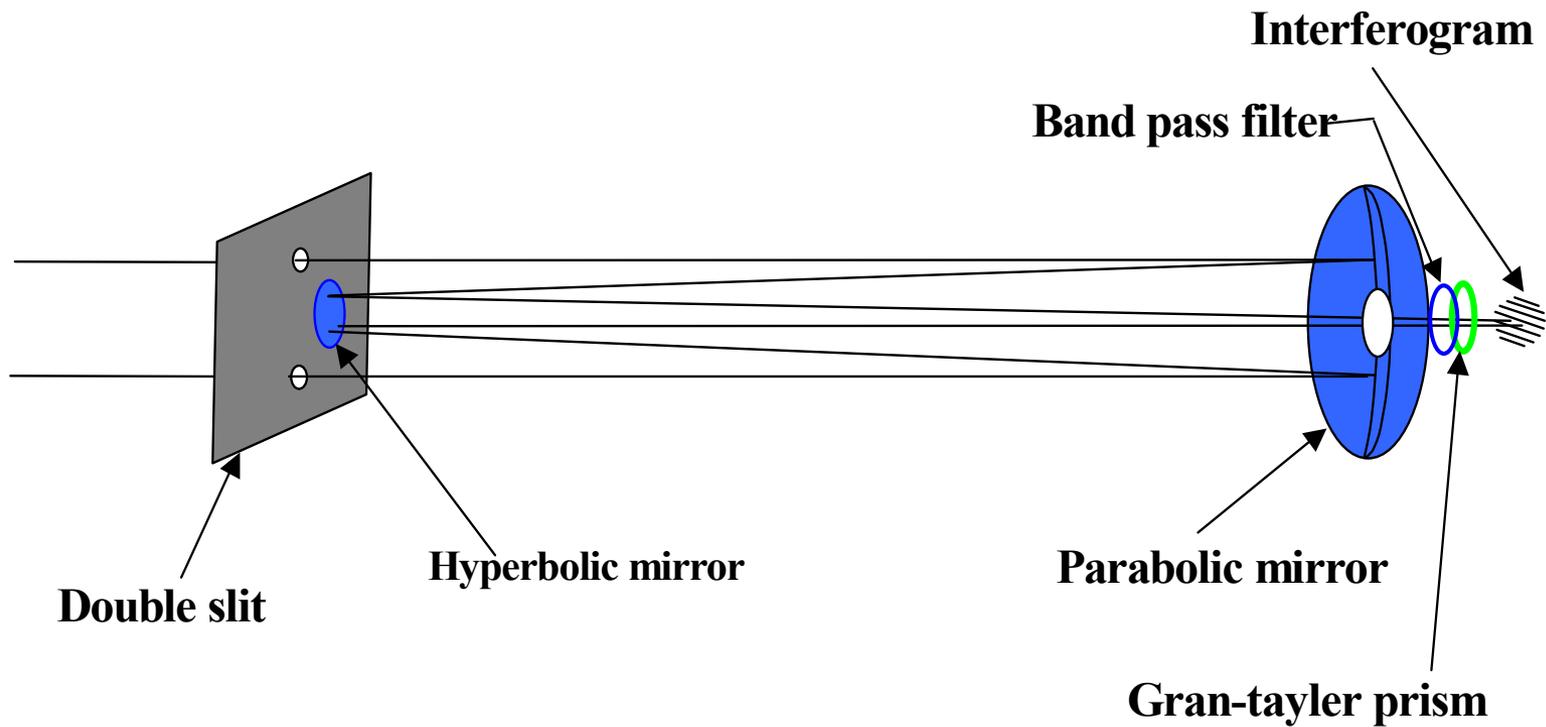
Possible arrangement for reflective optics for interferometer

1. On axis arrangement

Newtonian arrangement of optics

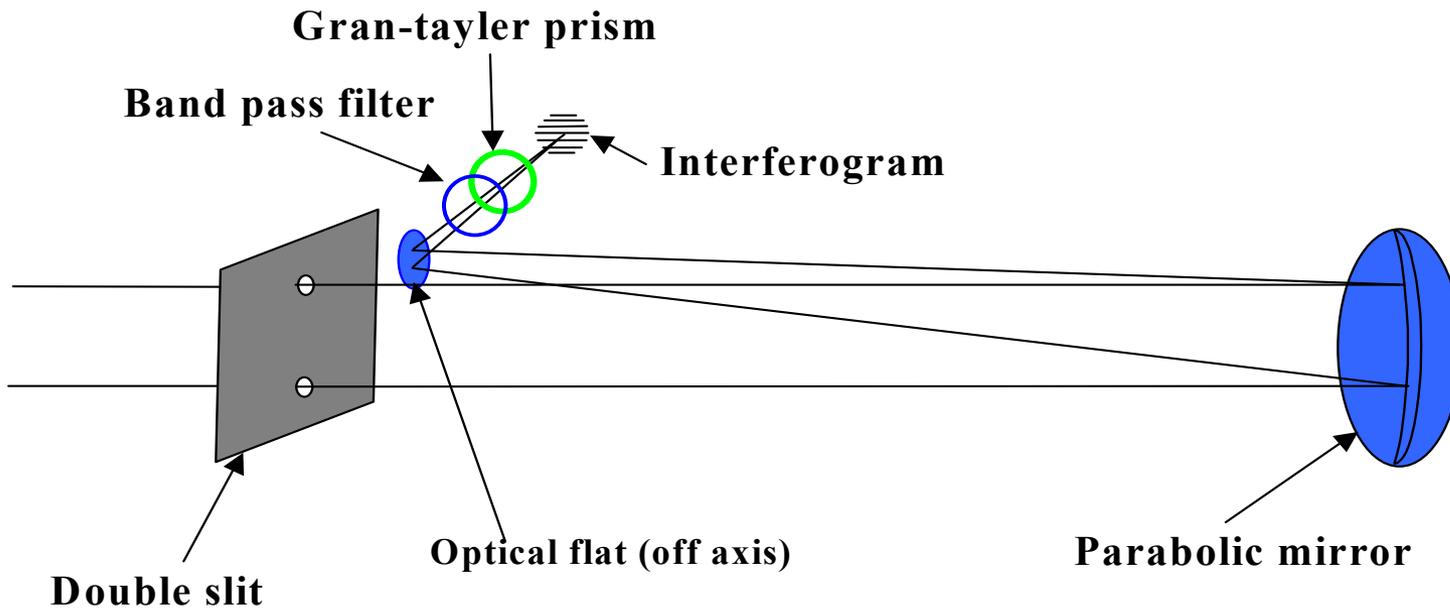


Cassegrainian arrangement of optics



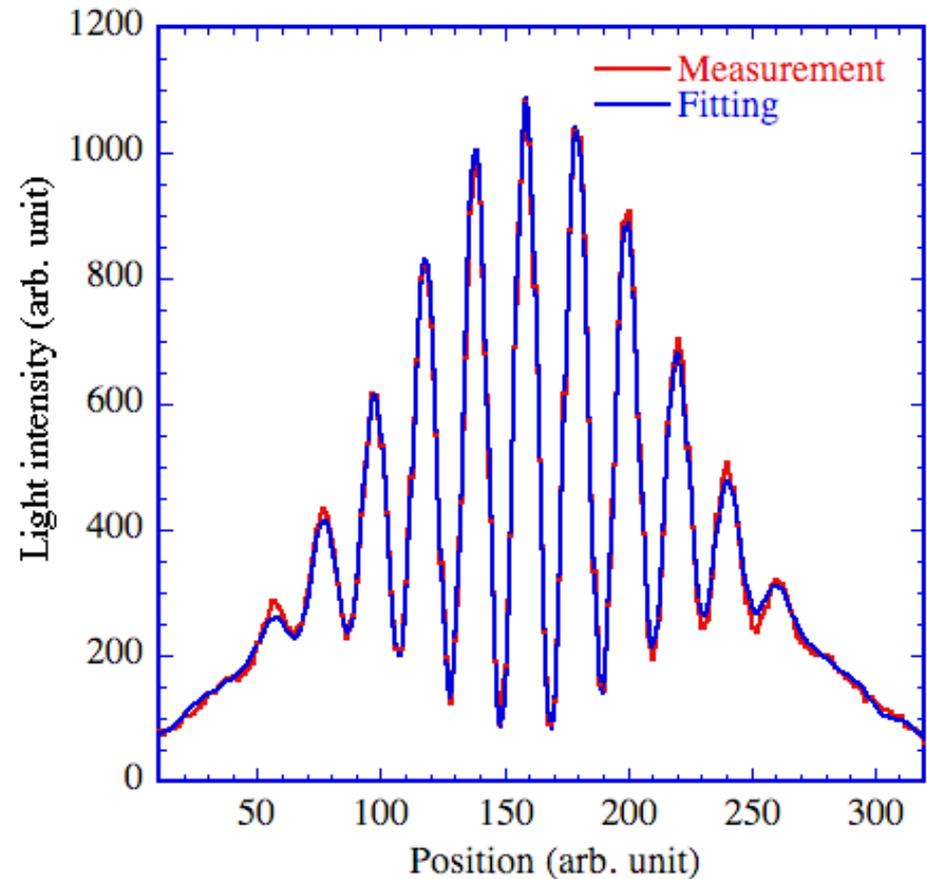
2. Off axis arrangement

Herschelian arrangement of optics

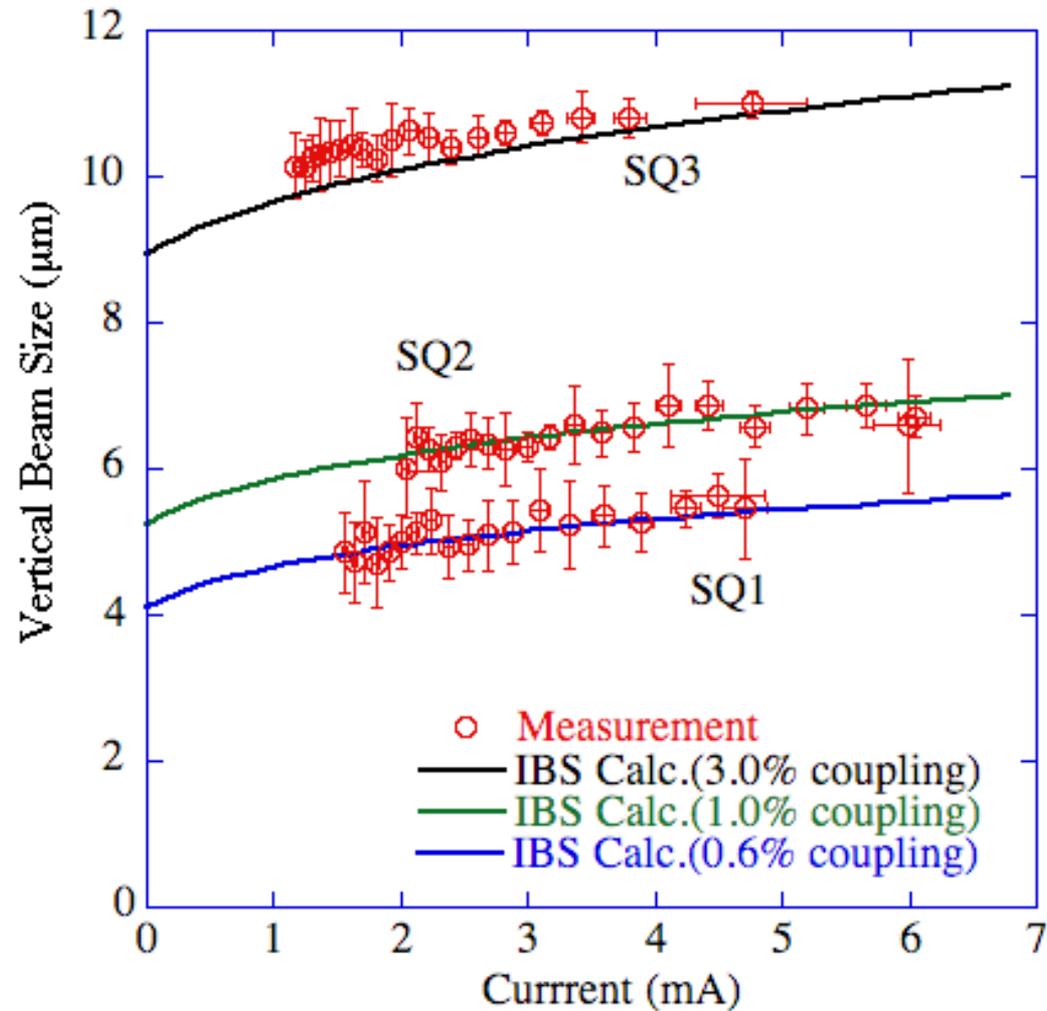


Measured interferogram At ATF, KEK

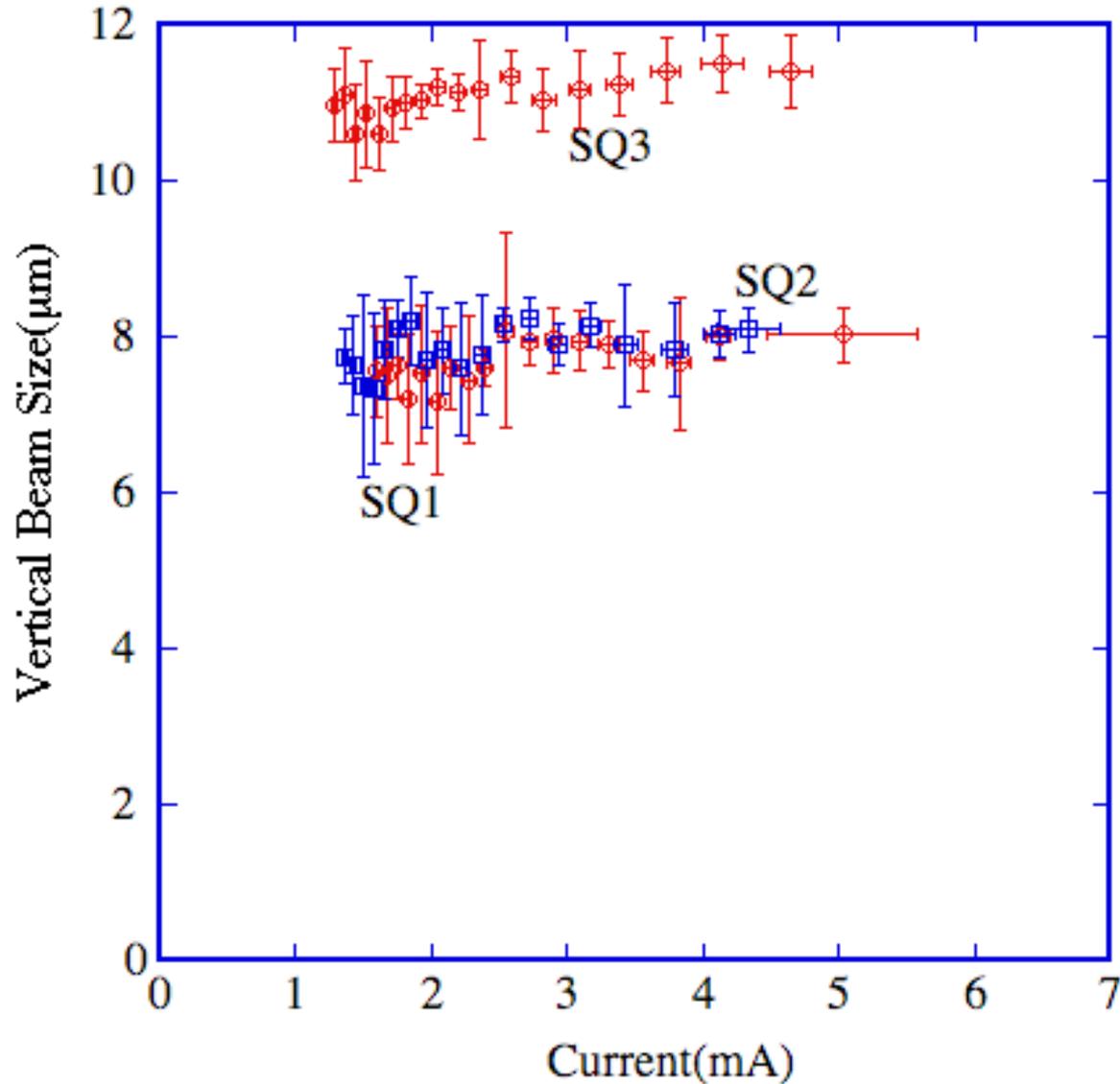
Result of beam size is
 $4.73\mu\text{m} \pm 0.55\mu\text{m}$



The x-y coupling is controlled by the strength of the skew Q at ATF



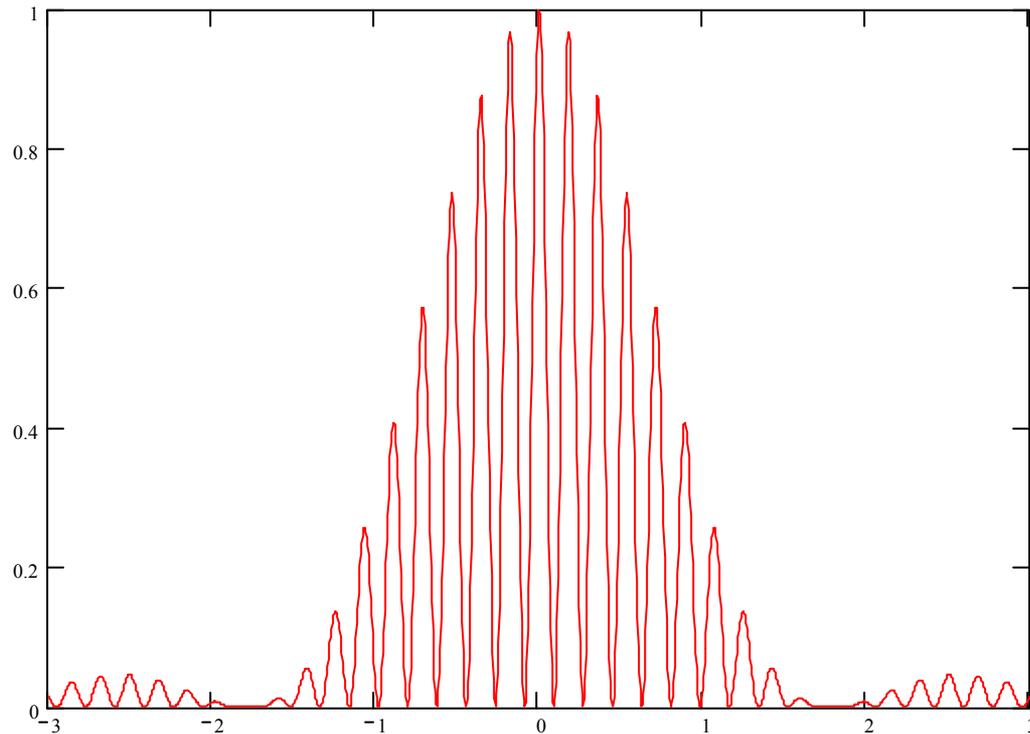
Remember same results by normal refractive interferometer using $\lambda=400\text{nm}$



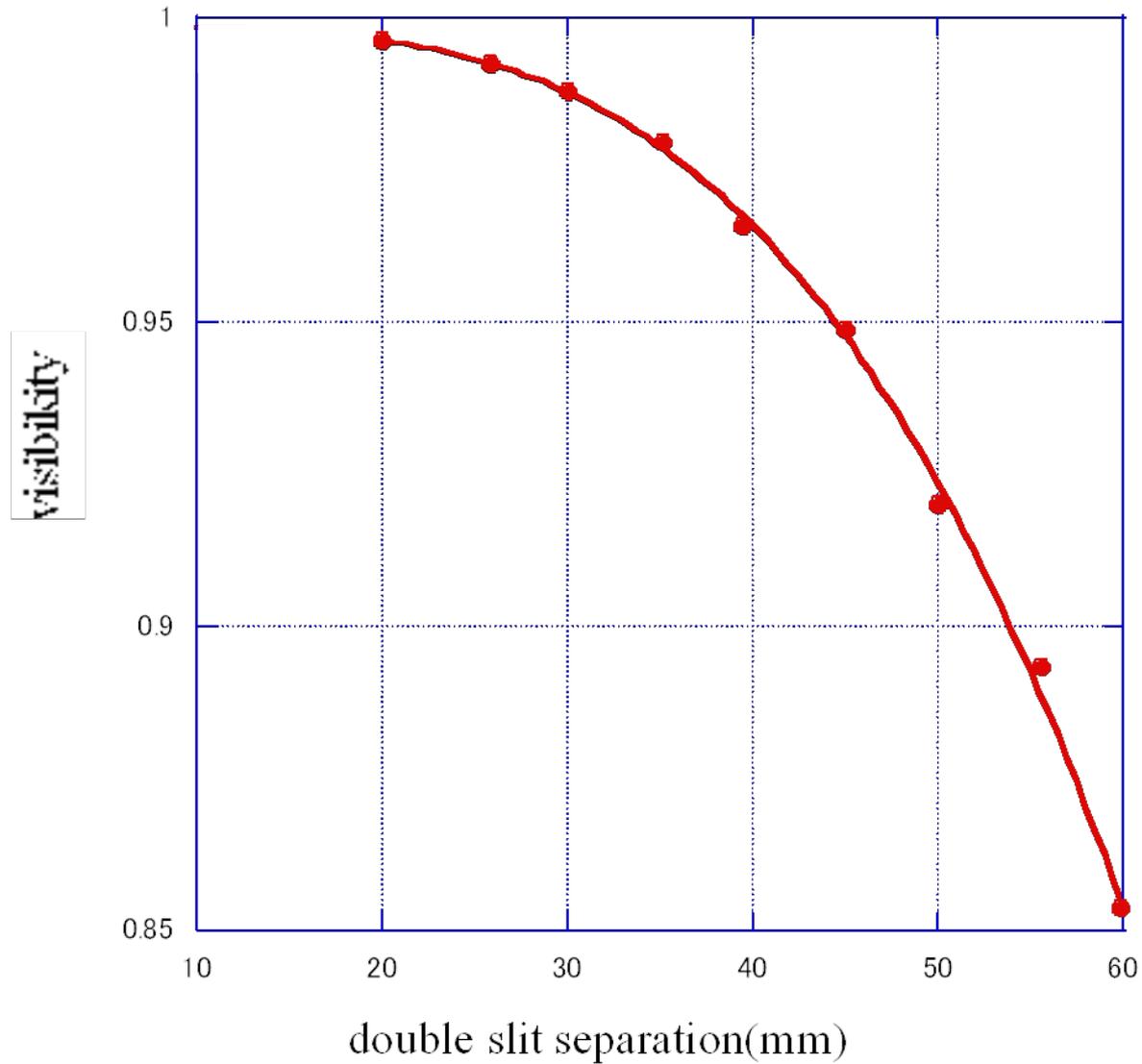
The reflective interferometer is more useful than refractive interferometer especially for shorter wavelength range.

Actually, it is **chromatic aberration-free**, and reflectors are cheaper than lenses in large aperture.

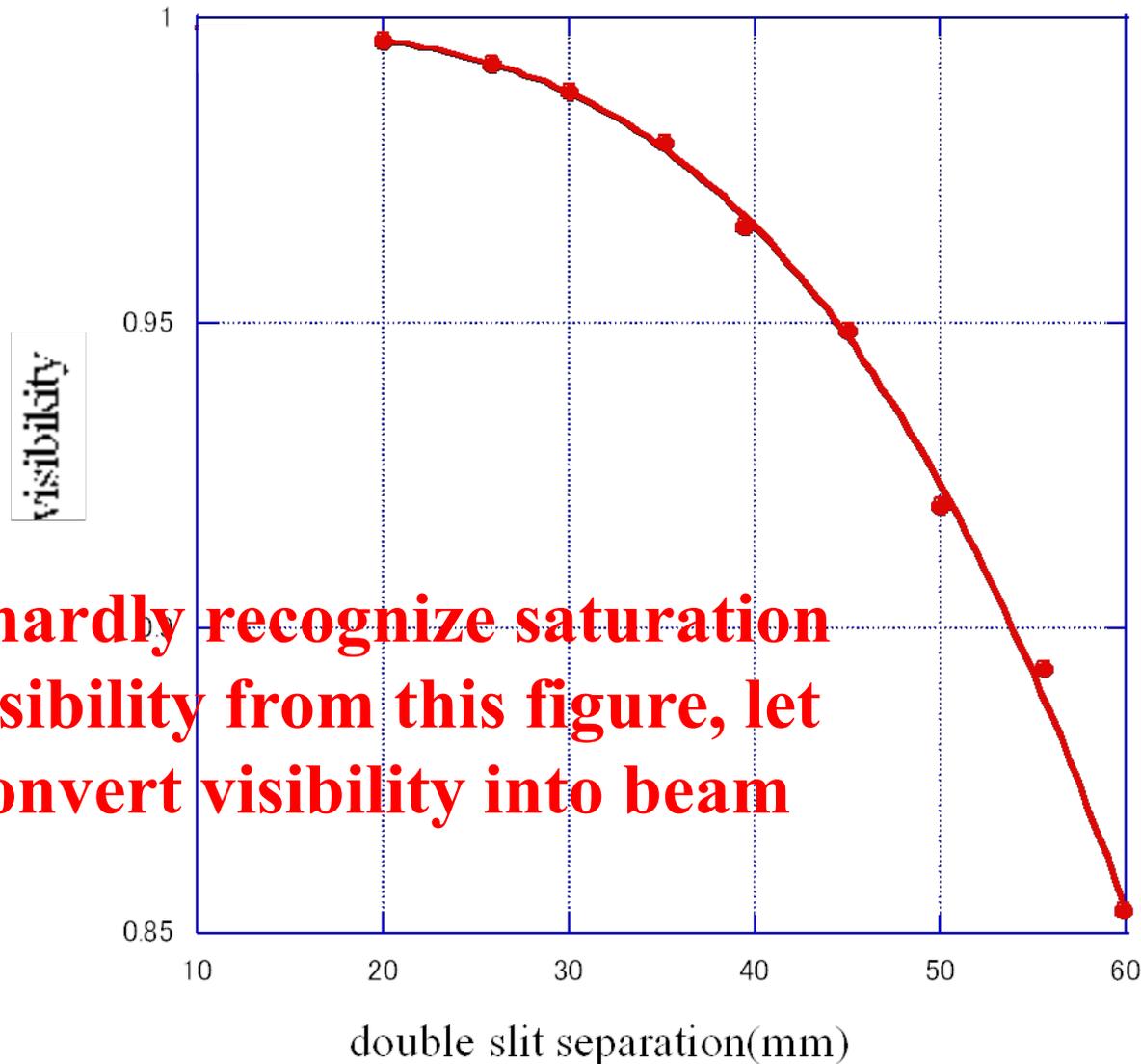
**If we can not use more shorter wave
length,
How we can do for more smaller beam
size measurement?**



Result of visibility for beam size $5.8\mu\text{m}$ ($\lambda=550\text{nm}$) with several separation of double slit.

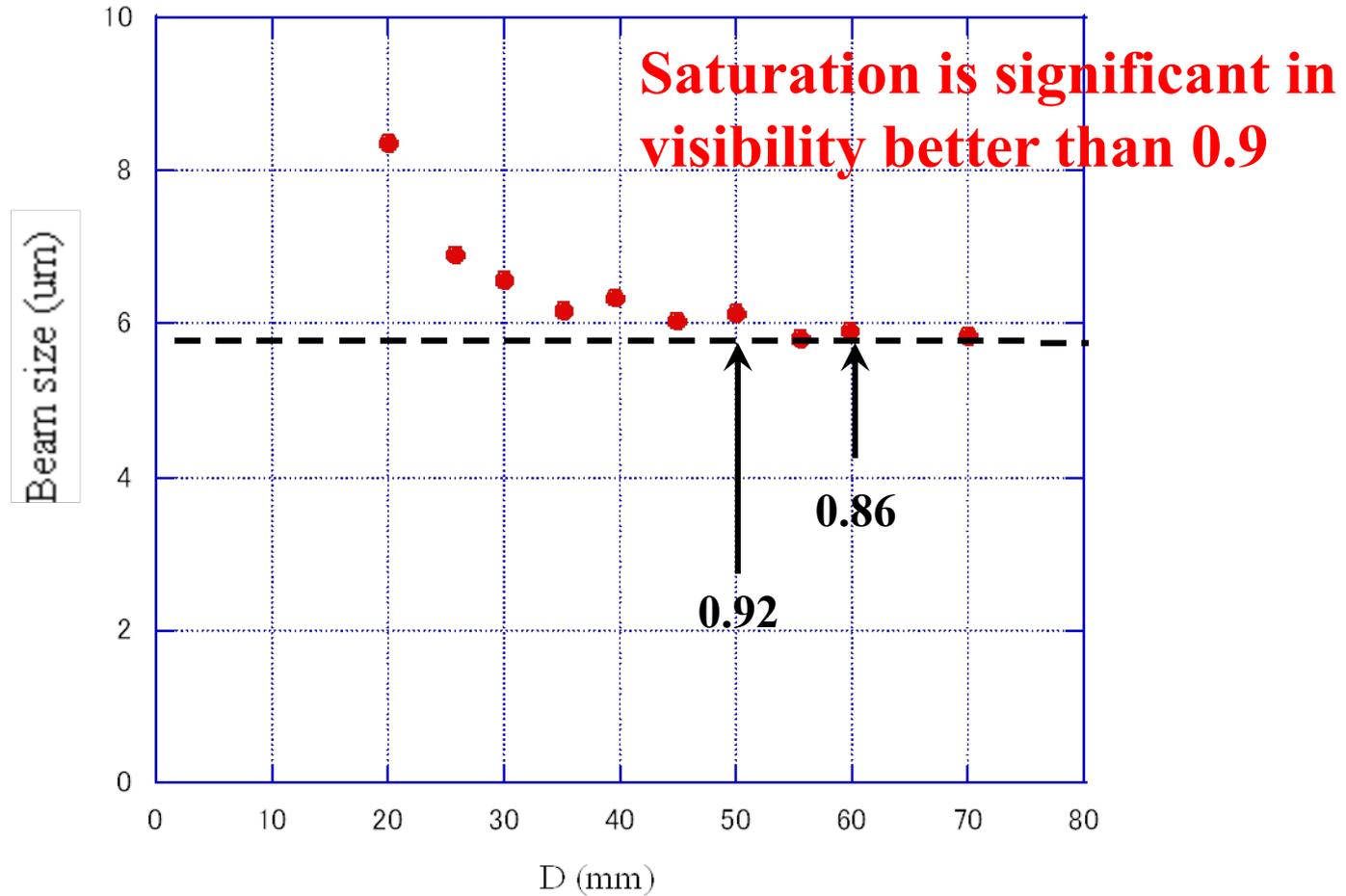


Result of visibility for beam size $5.8\mu\text{m}$ ($\lambda=550\text{nm}$) with several separation of double slit.



We hardly recognize saturation in visibility from this figure, let us convert visibility into beam size!

Convert visibility into beam size. We can see clear saturation in smaller double slit range which has visibility near 1.



4. Imbalanced input method

**Another method to escape from noise
for
more small beamsize measurement**

Let's us consider equation for interferogram.

$$I(y, D) = \int (I_1 + I_2) \cdot \left\{ \sin c \left(\frac{\pi \cdot a \cdot y \cdot \chi(D)}{\lambda \cdot f} \right) \right\}^2 \cdot \left\{ 1 + \gamma \cdot \cos \left(k \cdot D \cdot \left(\frac{y}{f} + \psi \right) \right) \right\} d\lambda$$

$$\gamma = \left(\frac{2 \cdot \sqrt{I_1 \cdot I_2}}{I_1 + I_2} \right) \cdot \left(\frac{I_{\max} - I_{\min}}{I_{\max} + I_{\min}} \right), \quad \psi = \tan^{-1} \frac{S(D)}{C(D)}$$

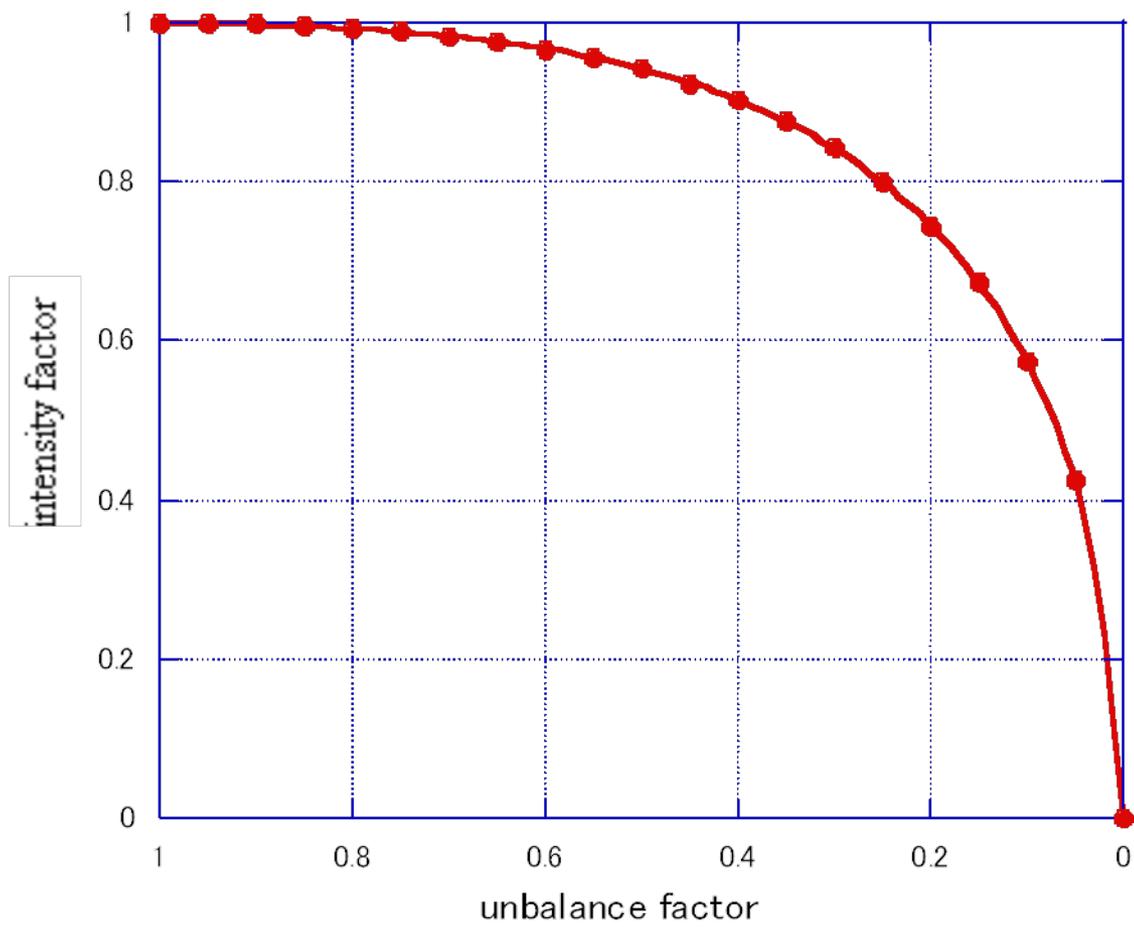
In this equation, the term “ γ ” has not only real part of complex degree of spatial coherence but also intensity factor!

$$\gamma = \left(\frac{2 \cdot \sqrt{I_1 \cdot I_2}}{I_1 + I_2} \right) \cdot \left(\frac{I_{\max} - I_{\min}}{I_{\max} + I_{\min}} \right)$$

If $I_1=I_2$, γ is just equal to real part of complex degree of spatial coherence , but if $I_1 \neq I_2$, we must take into account of intensity factor;

$$\frac{2 \cdot \sqrt{I_1 \cdot I_2}}{I_1 + I_2}$$

This intensity factor is always smaller than 1 for $I_1 \neq I_2$.

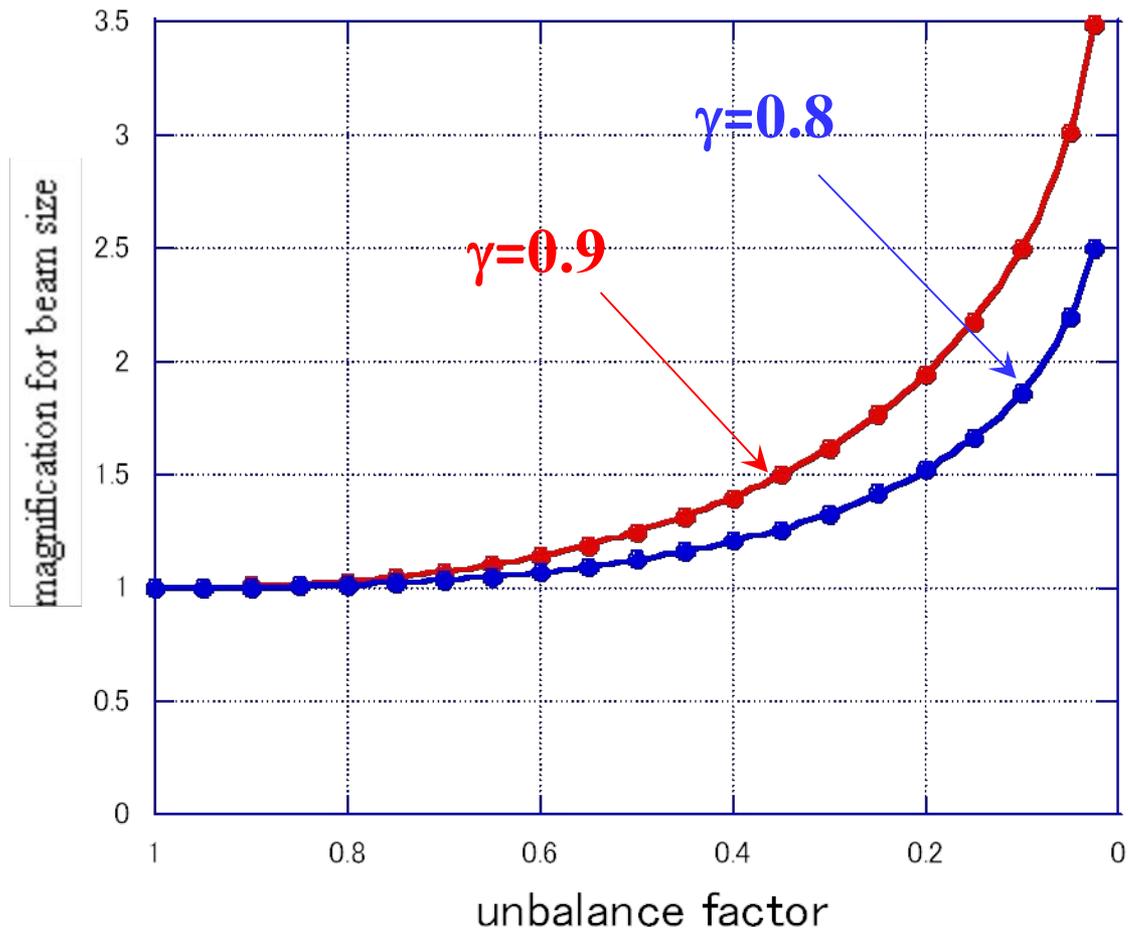


$$\gamma = \left(\frac{2 \cdot \sqrt{I_1 \cdot I_2}}{I_1 + I_2} \right) \cdot \left(\frac{I_{\max} - I_{\min}}{I_{\max} + I_{\min}} \right)$$

Since intensity factor is smaller than 1 for $I_1 \neq I_2$, the “ γ ” will be observed smaller than real part of complex degree of spatial coherence.

This means beam size will be observed larger than primary size and we know ratio between observed size and primary size.

This is magnification!

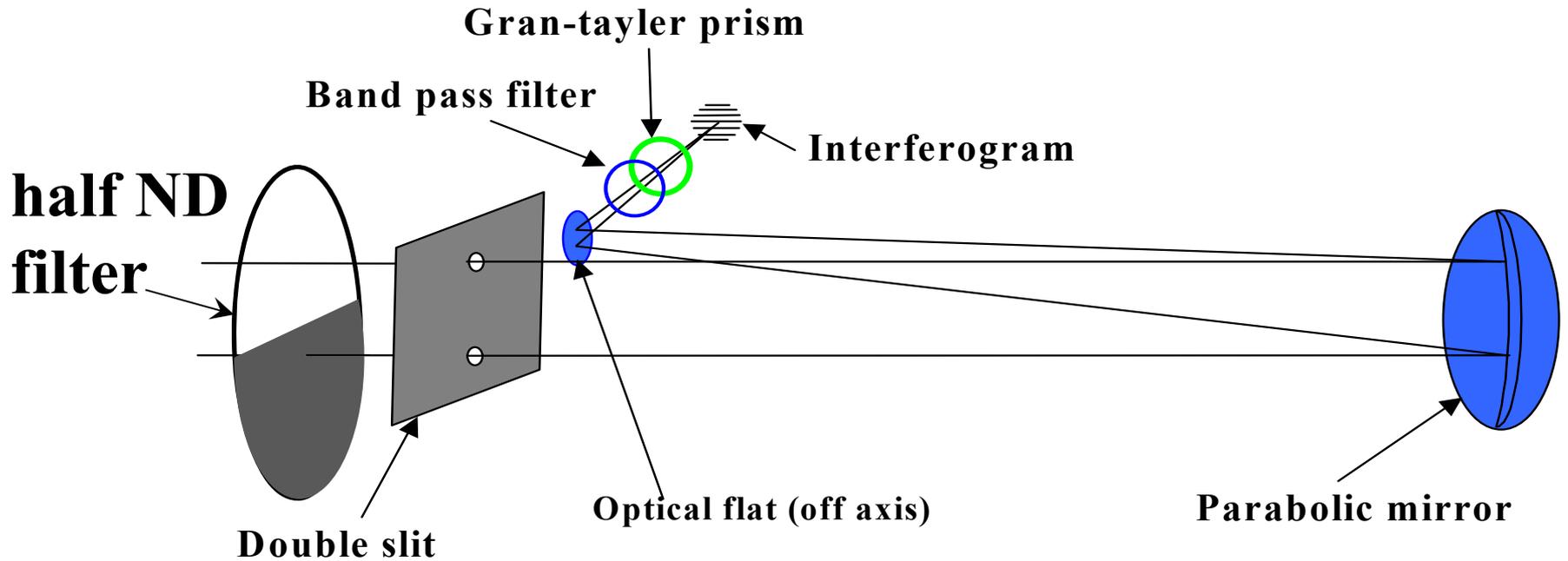


We can use magnification range up to 2 for $I_1 : I_2 = 1 : 0.2$ or 3 for $1 : 0.05$.

**In interferometry, we can
magnify beam size by very
simple way;**

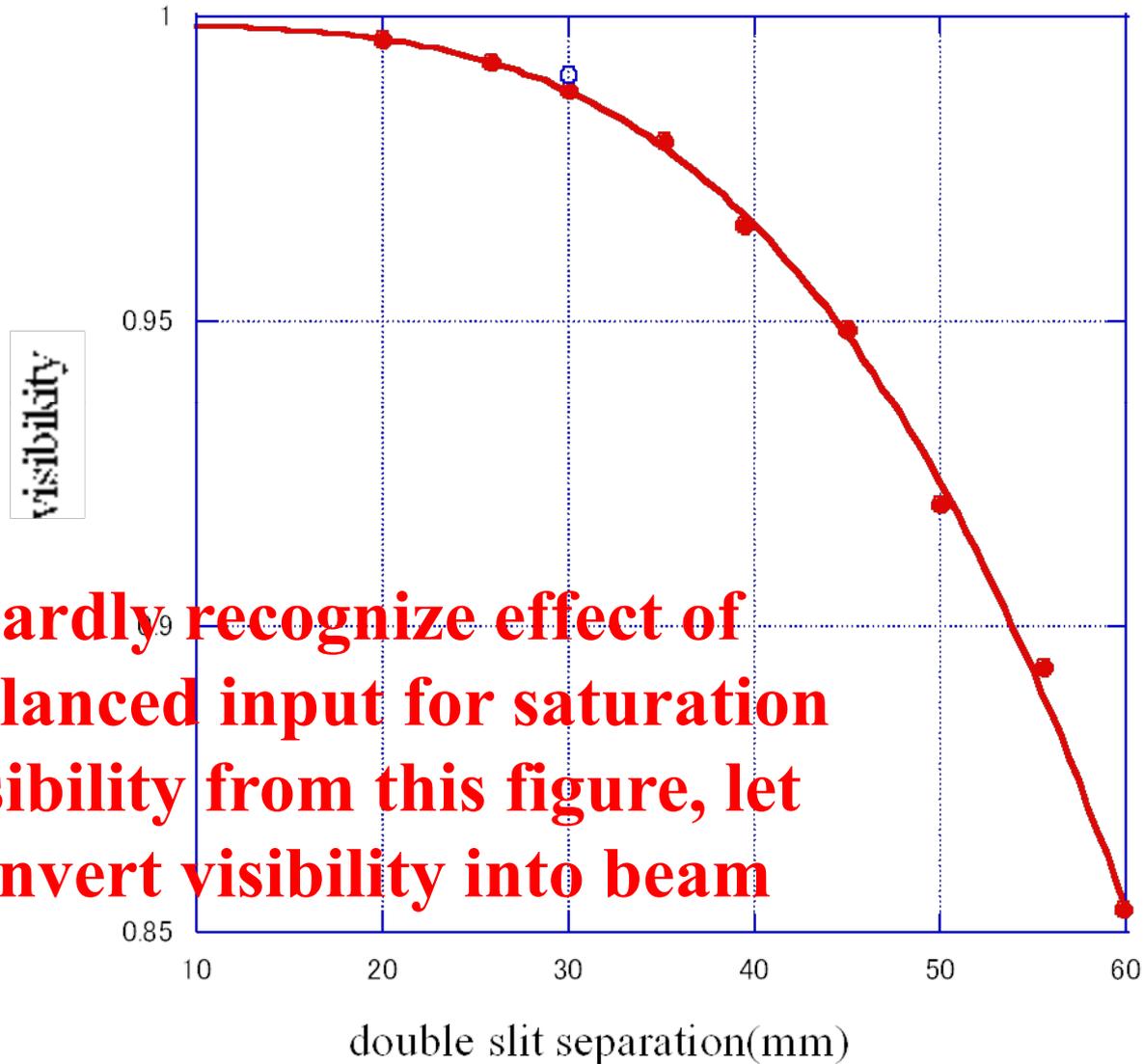
**applying imbalance input for
double slit!**

Setup for imbalanced input by half ND filter

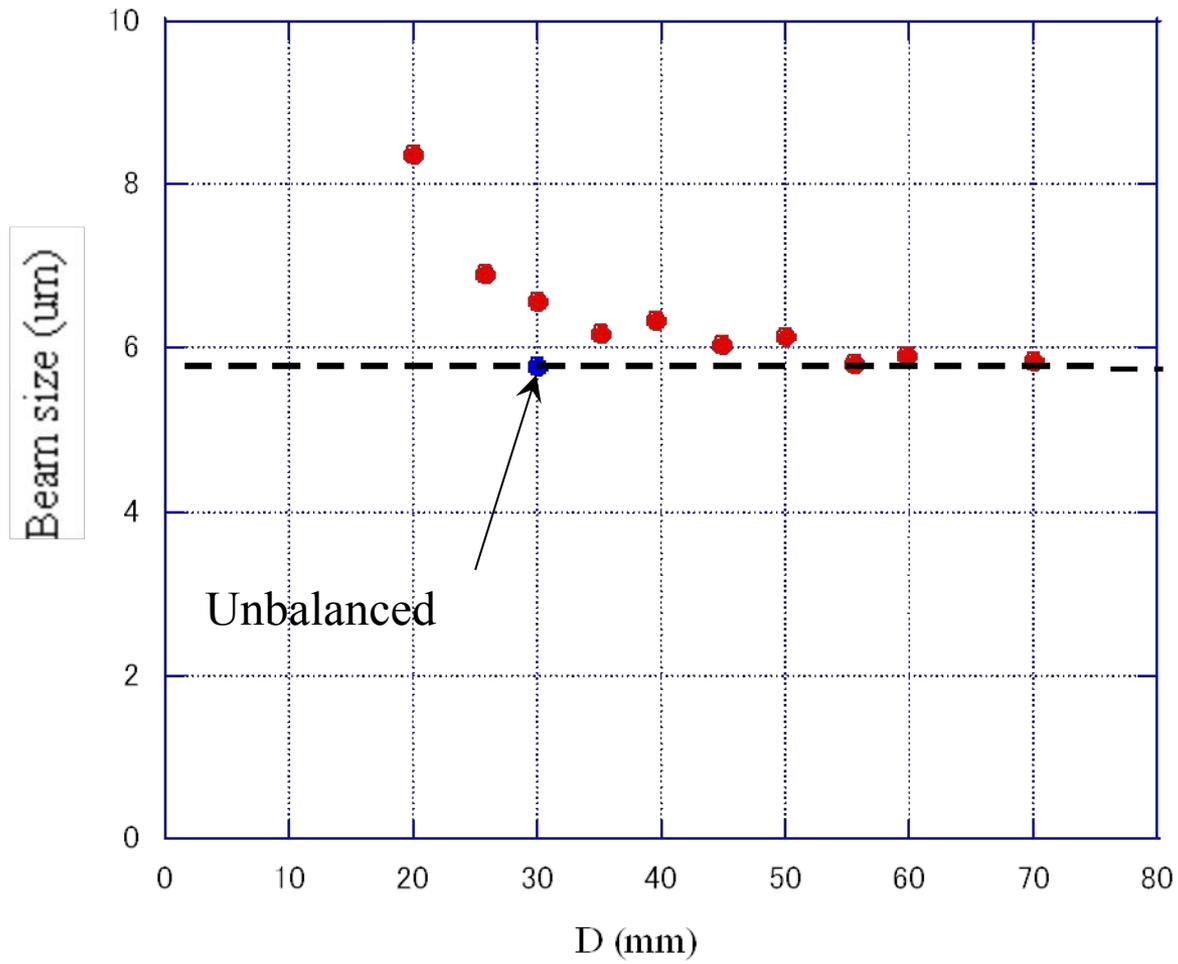


Applying unbalance method for $D=30\text{mm}$.

$I_1 : I_2 = 0.853 : 0.249$



We hardly recognize effect of unbalanced input for saturation in visibility from this figure, let us convert visibility into beam size!



Further result of
unbalanced technique,
please hear presentation
of Dr. Mark Boland

Conclusion

Smallest result of beam size at ATF is $4.7\mu\text{m}$ with reflective SR interferometer using double slit separation of 45-55mm, $\lambda=400\text{nm}$. **This size is almost small limit with equal input method.**

When we will apply imbalanced method;

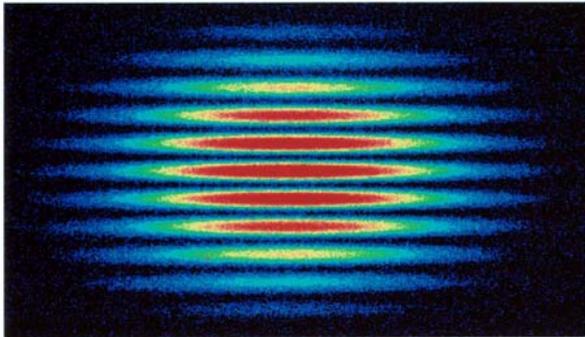
With magnification factor 2 \longrightarrow $2.4\mu\text{m}$

With magnification factor 3 \longrightarrow $1.6\mu\text{m}$

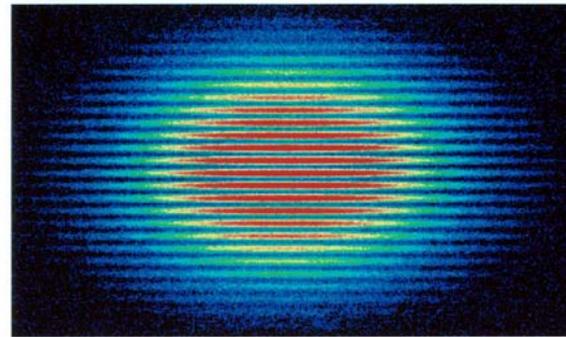
We are waiting beam size in this range!

Thank you very much for your attention.

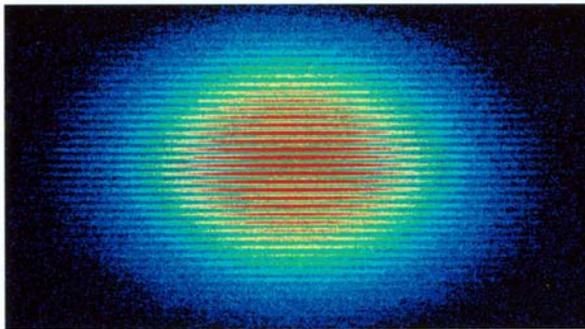
$\lambda = 550\text{nm}$



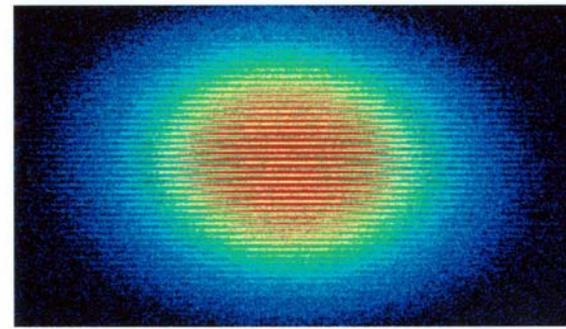
D=6.7mm (1.79mrad)



D=14.7mm (3.92mrad)



D=22.7mm (6.05mrad)



D=28.7mm (7.65mrad)