RECONSTRUCTION OF CONTROLLED OCTUPOLAR ERRORS IN CERN-SPS

Angelina Parfenova,

G. Franchetti, GSI, Darmstadt R. Tomas, G. Vanbavinckhove, CERN, Geneva, Switzerland

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METHODS OF RECONSTRUCTION OF NONLINEAR COMPONENTS

O. Boine-Frankenheim, these proceedings



SIS100 tune diagram

KNOWLEDGE OF NONLINEAR ERRORS ALLOWS TO CONTROL AND COMPENSATE RESONANCE DRIVEN BEAM LOSS

NONLINEAR CHROMATICITY

G. Arduini et. al., PAC, Portland, USA, p. 2240 (2003)

RESONANCE DRIVING TERM

R. Bartolini et. al., ICAP, Chamonix, France, (2006) R. Tomas et. al, Phys. Rev. ST 8, 024001 (2005)

ORBIT DEFORMATION via LOCAL BUMP

F. Pilat et al., PAC, Knoxville, USA, p. 601 (2005)

GLOBAL ORBIT DISTORTION

NONLINEAR TUNE RESPONSE MATRIX (NTRM) METHOD

G. Franchetti, A. Parfenova and I. Hofmann, Phys. Rev. ST 11, 094001 (2008)

NTRM-MODELLING LINEARIZED THEORY

IF THE CLOSED ORBIT IS DEFORMED BY N_t of steering angles θ_X and θ_Y

X-/Y- CLOSED ORBIT (CO) AT NONLINEAR ERROR

$$\textbf{x, y}_{COI} = \sum_{t=1}^{N_t} M_{It}^{x,y} \boldsymbol{\theta}_{x,yt}$$

EQUATIONS OF MOTION FOR A PARTICLE WITH A SMALL AMPLITUDE IN AN ACCELERATOR WITH A DEFORMED CO



CONTRIBUTION TO THE MACHINE TUNES WITH RESPECT TO THE DEFORMED CLOSED ORBIT

RECONSTRUCTION OF CONTROLLED SEXTUPOLAR ERRORS IN GSI-SIS18

THE FIRST ORDER CONTRIBUTION ON THE MACHINE TUNES



LINEAR SYSTEM (2 X 2) FOR 2 ERRORS



Normal		ΔK_2	Calc.	Exp.	Rel. Err.,
errors			×10 ⁻² [m ⁻²]		%
S 1	1	-2	-1.99	-1.79	10.5
	2	1	1.00	1.02	1.8
S 2	1	-4	-3.99	-4.13	3.3
	2	2	2.00	1.55	22.7
S 3	1	-8	-7.99	-7.61	4.9
	2	4	4.00	3.90	2.5
S 4	1	5	5.00	4.97	0.6
	2	-3	-2.99	-2.74	8.7

EXPERIMENTAL BENCHMARKING IN CERN-SPS



SET 2 PROBING NORMAL OCTUPOLAR ERRORS AND USE 2 HORIZONTAL STEERERS

CHANGE THE CLOSED ORBIT AND MEASURE THE TUNE RESPONSE

RECONSTRUCT THE 2 OCTUPOLAR ERRORS

THE EFFECT OF THE OCTUPOLES

THE SECOND ORDER CONTRIBUTION TO THE MACHINE TUNES WITH RESPECT TO THE DISTORTED CO (CONSIDERING SEVERAL OCTUPOLAR ERRORS)

$$\begin{cases} \sum_{x} Q_{tt}^{xx} = \frac{1}{2} \cdot \frac{1}{4\pi} \sum_{l=1}^{N_l} \beta_{xl} K_{3l} (M_{lt}^x)^2 + O(2) \\ \Delta Q_x = \sum_{t=1}^{N_t} \sum_{x} Q_{tt}^{xx} (\theta_{xt})^2 \end{cases}$$

/ IS INDEX FOR OCTUPOLES ($N_{t}=2$), t IS INDEX FOR STEERERS ($N_{t}=2$)

A SECOND ORDER CONTRIBUTION FROM SEXTUPOLES **O(2)** EXISTS *Ref. G. Franchetti et. al., Phys. Rev. ST 11, 094001 (2008)*

WHEN THE NORMAL OCTUPOLAR ERRORS ARE EXCITED, ONLY HORIZONTAL ORBIT DEFORMATION CAN REVEAL THEM



DIFFICULTY DUE TO SYMMETRIC ARRANGMENT OF OCTUPOLES AND STEERERS IN CERN-SPS

TWO SERIES OF OCTUPOLES **LOF** AND **LOD** ARE ARRANGED WITH THE SAME PERIODICITY (6)

ALL THE STEERERS ARE SYMMETRICALLY PLACED IN THE RING



BREAKING THE SPS SYMMETRY

$${}_{x}Q_{11}^{xx} = \{\beta_{xLOD}K_{3LOD}(M_{LOD\,1}^{x})^{2} + \beta_{xLOF}K_{3LOF}(M_{LOF\,1}^{x})^{2}\}/2/4\pi$$
$${}_{x}Q_{22}^{xx} = \{\beta_{xLOD}K_{3LOD}(M_{LOD\,2}^{x})^{2} + \beta_{xLOF}K_{3LOF}(M_{LOF\,2}^{x})^{2}\}/2/4\pi$$

FOR EACH STEERER S1, S2 OR S3, ... ETC WITH $\theta_1 x Q_{22}^{xx} =_x Q_{11}^{xx}$

COMBINING STEERERS GENERATES A NEW OPTICAL STRUCTURE!!!

FOR A 'VIRTUAL' STEERER S+ WITH θ + (COMBINATION OF S1 AND S2) $\theta_1 = \theta_2 = \theta_+$

$$S^{+} = (S1, S2) = (\theta_{1}, \theta_{2}) = \theta_{+}$$

$${}_{x}Q_{+}^{xx} = 2({}_{x}Q_{11}^{xx} + {}_{x}Q_{12}^{xx}) \neq_{x} Q_{11}^{xx}$$



THE EXPERIMENTAL CONDITIONS

1. ENERGY **26** GeV, INTENSITY $\sim 10^{11}$

2. HORIZONTAL STEERERS:

MDH10207 (AS S) MDH10207 AND MDH20407 (AS S+)

3. A FAST TRANSVERSE KICK WAS GENERATED ON BOTH –X AND –Y PLANES

4. QX, QY WERE MEASURED FOR 1024 TURNS

5. THE STEERING RANGE [-150 µRAD; 150 µRAD]

6. CHROMATICITY SEXTUPOLES WERE SWITCHED ON

7. Two octupole families were excited with the strengths LOF=2 M^{-3} and LOD=4 M^{-3}



TUNE RESPONSE vs. CLOSED ORBIT DEFORMATION (COD)

MEASUREMENT SIMULATION **Chromaticity sextupoles** Chromaticity sextupoles & & octupoles Octupoles 0.134 Fractional Qx 0.134 Chromaticity sextupoles **Chromaticity** sextupoles Ô Fractional 0.132 0.132 0130 0.130 100 -150 -100 -50 50 150 100 n -150 -100 -50 50 150 Steering angle S, mrad Steering angle S, mrad **Chromaticity sextupoles** & octupoles 0.134 0.135 Fractional Qx ð Chromaticity sextupoles 0.132 0.130 Fraglic Chromaticity sextupoles 0.125 ——Chromaticity sextupoles & 100 150 -100 -50 100 150 -150 -100 -50 -150 0 50 50 0 Steering angle S⁺. mrad Steering angle S+, mrad

DIFFERENTIAL TUNE RESPONSE: CONTRIBUTION OF OCTUPOLES

MEASUREMENT SIMULATION S xQ_{11}^{XX} S⁺ 0.004 0.004 $_{x}Q_{+}^{xx}$ xQ_{11}^{XX} $_{x}Q_{+}^{xx}$ Š 0.003 0.003 Š 0.002 0.002 0.001 150 -100 0 50 100 -150 -100 -50 50 100 150 -150 -50 0 Steering angle S and S⁺, mrad Steering angle S and S⁺, mrad N I

$$_{x}Q_{tt}^{xx} = \frac{1}{2} \cdot \frac{1}{4\pi} \sum_{l=1}^{N_{l}} \beta_{xl} K_{3l} (M_{lt}^{x})^{2} + O(2)$$

NOTE THAT:

1. LARGE TUNE FLUCTUATIONS WERE OBSERVED IN THE MEASUREMENT

2. STEERERS S AND S+ WERE VARIED ALMOST TO THE MAXIMUM STEERING RANGE [-150 μ RAD; 150 μ RAD]

3. OCTUPOLES WERE EXCITED STRONG LOF=2 M⁻³ AND LOD=4 M⁻³

MEASUREMENT OF INITIAL CLOSED ORBIT (CO)

PRECONDITION FOR NTRM IS INITIALLY CORRECTED CO

THE METHOD HAD TO BE APPLIED WITH A CERTAIN INITIAL CO DISTORTION



MEASURED AND SIMULATED CLOSED ORBIT DEFORMATION (COD)

ARE WE IN THE REGIME OF LINEAR COD?

A CO EXAMPLE FROM A SINGLE BPM (#20608)

MEASUREMENT

SIMULATION



steerer S+ is a combination of steerers MDH10207 and MDH20407

MEASUREMENT AND SIMULATION: DIFFERENTIAL CO RESPONSE

SUBTRACTING CO MODULATION WITH- (LOF=2 M⁻³ AND LOD=4 M⁻³) AND WITHOUT OCTUPOLES FOR EACH STEERER S AND S+ RESPECTIVELY



THE MEASURED CO DATA DO NOT ALLOW TO DETERMINE THE RANGE OF THE LINEAR CO REGIME!

APPLICABILITY OF THE NTRM THEORY

SIMULATION



NUMERICALLY RECONSTRUCTED STRENGTHS FOR THE TWO OCTUPOLAR SERIES:

1. CO IN LINEAR REGIME

LOF=1.99 M⁻³ AND LOD=4.01 M⁻³

2. CO TAKEN OVER THE FULL RANGE [-150 μRAD; 150 μRAD]

LOF=0.86 M⁻³ AND LOD=18.24 M⁻³

FOR A WEAKER OCTUPOLAR EXCITATION THE LINEAR REGIME OF THE CLOSED ORBIT IS LARGER



RESULTS AND DISCUSSION

SMALLER RANGE



THE RECONSTRUCTION DEPENDS ON THE STEERER VARIATION RANGE!!!

RESULTS FOR ALL RANGES ARE SHOWN

LOD IS NOT RECONSTRUCTED

TO THE SET VALUE

LOF=2.35 +/- 0.32 M⁻³



CONCLUSION

RECONSTRUCTION OF SEXTUPOLES IN GSI-SIS18:

THE NEW NTRM METHOD WAS SUCCESSFULLY VALIDATED WITH TWO CONTROLLED SEXTUPOLAR ERRORS IN GSI-SIS18. THE BENCHMARKING OF NTRM CONTINUES WITH EXPERIMENTS FOR SIX TO TWELVE CONTROLLED SEXTUPOLAR ERRORS

RECONSTRUCTION OF OCTUPOLES IN CERN-SPS:

A FIRST ATTEMPT TO RECONSTRUCT EXPERIMENTALLY TWO CONTROLLED OCTUPOLAR ERRORS USING THE NTRM WAS CARRIED OUT IN CERN-SPS,

LARGE TUNE FLUCTUATIONS HAVE REQUIRED:

1. STRONG SET VALUES FOR OCTUPOLES **&**

2. A LARGE RANGE OF **COD**, WHICH ARE NOT OPTIMAL FOR THE **NTRM** METHOD.

3. THE SYMMETRIC ARRANGEMENT OF STEERERS **&** OCTUPOLES HAS CREATED ADDITIONAL DIFFICULTIES.

4. THE NTRM HAD TO BE APPLIED WITH A CERTAIN INITIAL CO DISTORTION.

AS A RESULT ONLY ONE OF THE TWO SET OCTUPOLES COULD BE RECONSTRUCTED

LOF=2.35 +/- 0.32 M⁻³



CONTENT

1. INTRODUCTION

2. NTRM and EXPERIMENT in CERN-SPS

- a. CONSEQUENCES OF THE SPS SYMMETRY
- **b.** The experimental procedure
- **c.** NUMERICAL EXAMPLES

3. EXPERIMENTAL RESULTS

- a. MEASUREMENT PARAMETERS
- **b.** LIMITS OF **NTRM** APPLICATION

4. CONCLUSION



MEASURED AND SIMULATED CO

Another BPM

A CO example from a single BPM (#60608) measured and simulated.



STATISTICAL APPROACH FOR SIMULATION DATA







INITIAL CO DEFORMATION FOR THE SIMULATION



WITH AN INITIAL COD OF ~ 2MM

MEASURED VERTICAL TUNE RESPONSE vs. COD

Measurement is not conclusive because of the large errorbars





MEASURED VERTICAL CO





DIFFERENTIAL TUNE RESPONSE: PHYSICS OF OCTUPOLES

Measuring both tunes allows to identify the polarity right away



